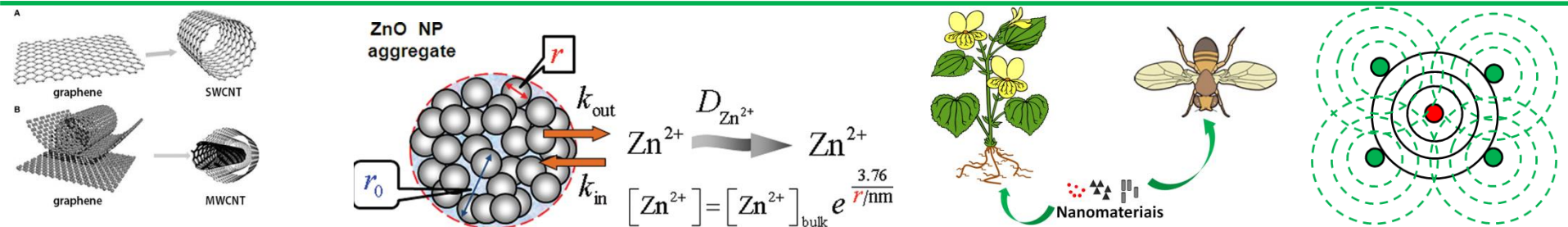


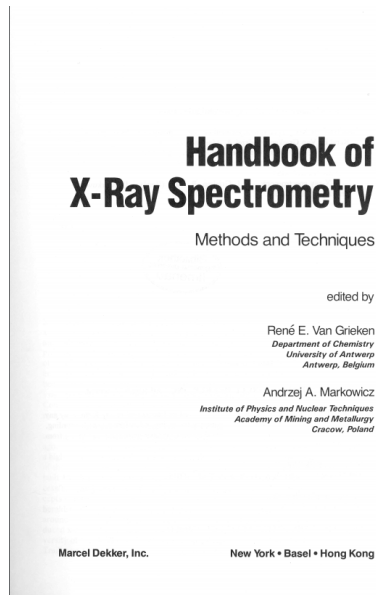


Interaction between radiation & matter

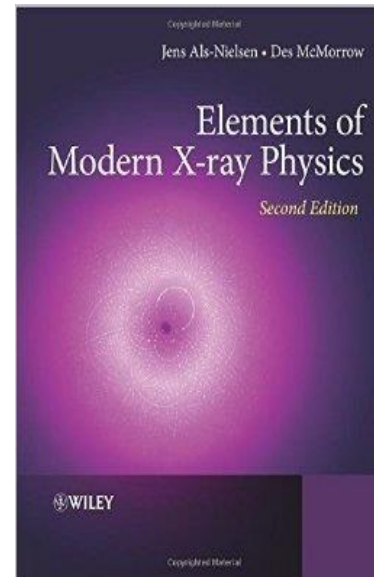
Prof. Hudson W.P. Carvalho
Applied X-ray Spectroscopy
Laboratory of Nuclear Instrumentation



Bibliography



Chapter 1



Chapter 1

Syllabus for today

- Types of interaction
- Cross sections

X-ray regime

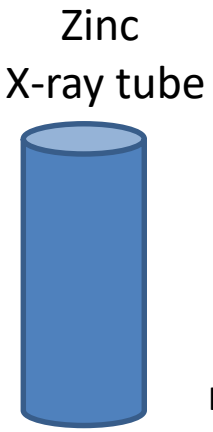
- X-ray Absorption and relaxation processes
- Elastic X-ray scattering
- Inelastic X-ray scattering

5.1 Types of interaction

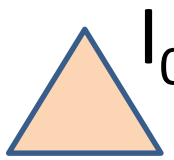
Monochromatic = only one wavelength

- λ_3 Fe = 1.94 A
- λ_3 Cu = 1.55 A
- λ_3 Pb = 0.99 A

X-ray zinc tube = 1.295 Angstroms or 0.1295 nm

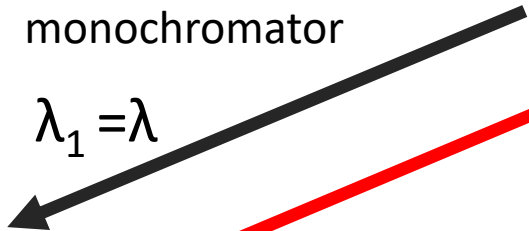


X-rays
 $\lambda = 1.295 \text{ A}$



I_0 Flux = 1000 photons s^{-1}

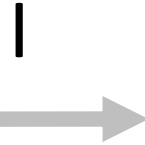
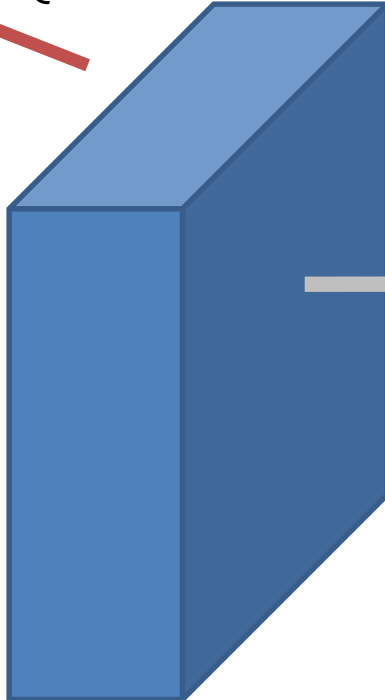
$\lambda_1 = \lambda$



$\lambda_2 > \lambda_1$

Characteristic emission X-ray fluorescence

metal plate



$T = I/I_0$

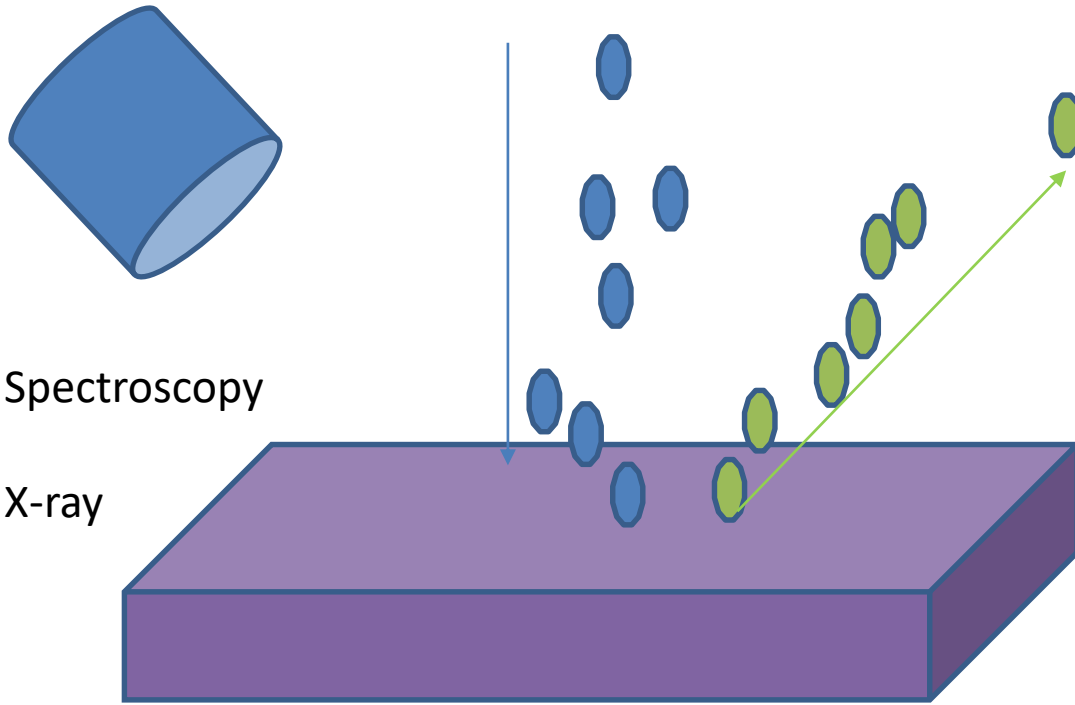
- Flux = 200 photons s^{-1}
- Flux = 150 photons s^{-1}
- Flux = 50 photons s^{-1}

$A = \ln(1/T)$

$\lambda_1 = 1.295 \text{ A}$ – Elastic Scattering or Thomson Scattering

$\lambda_2 = 1.395 \text{ A}$ – Inelastic Scattering or Compton Scattering

EDS = EDX
Energy Dispersive Spectroscopy
Energy Dispersive X-ray



It detects X-ray Fluorescence
It detects all elements above C, the limit of detection for EDS is ca. 1 wt%

$$E = h\nu$$

$$E = (hc) / \lambda$$

E- energy (eV)

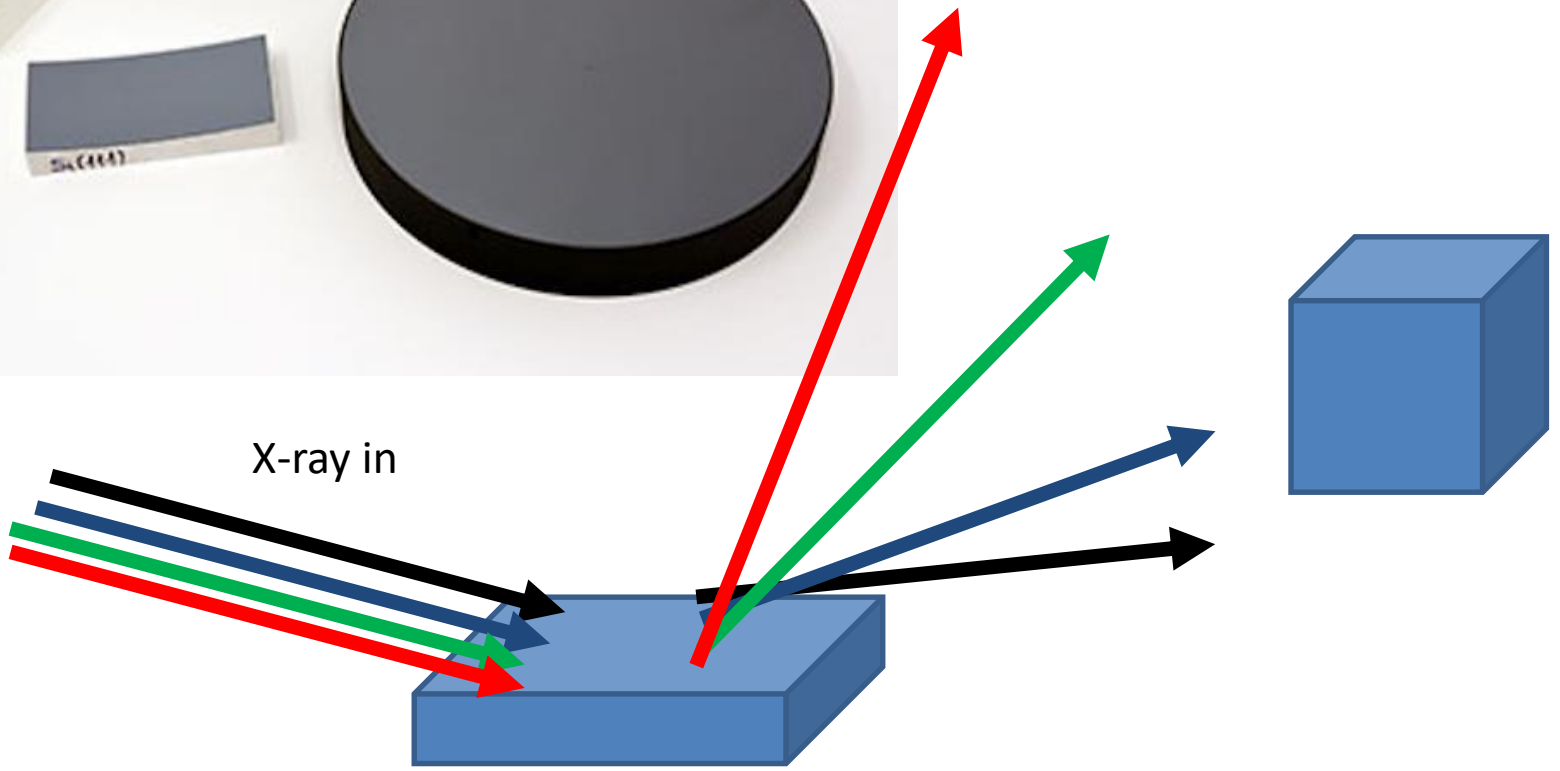
h – Planck's constant (eV)

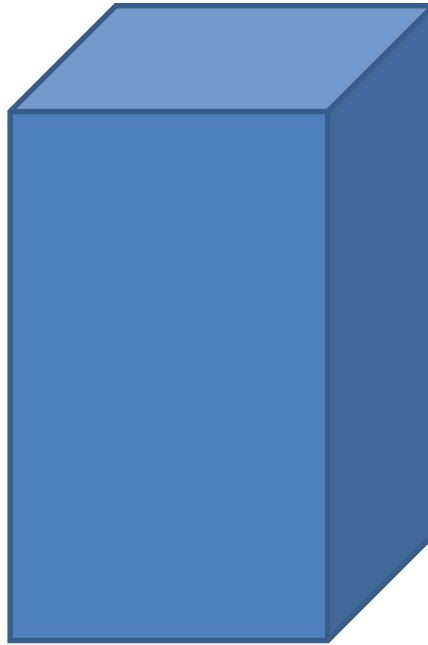
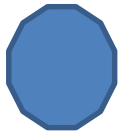
ν – frequency

$$\nu = c/\lambda$$

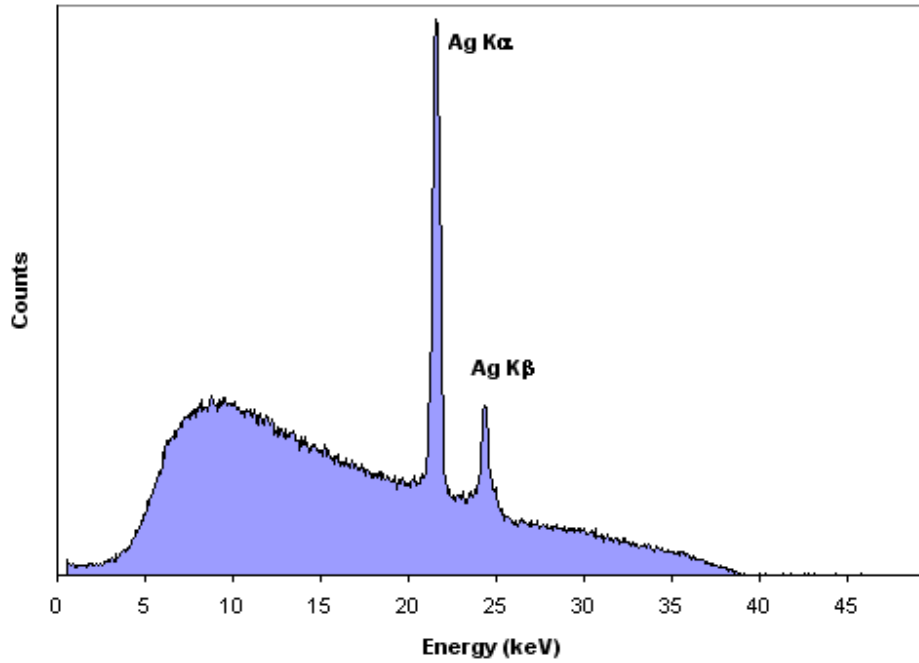
c- speed of light

λ - wavelenght

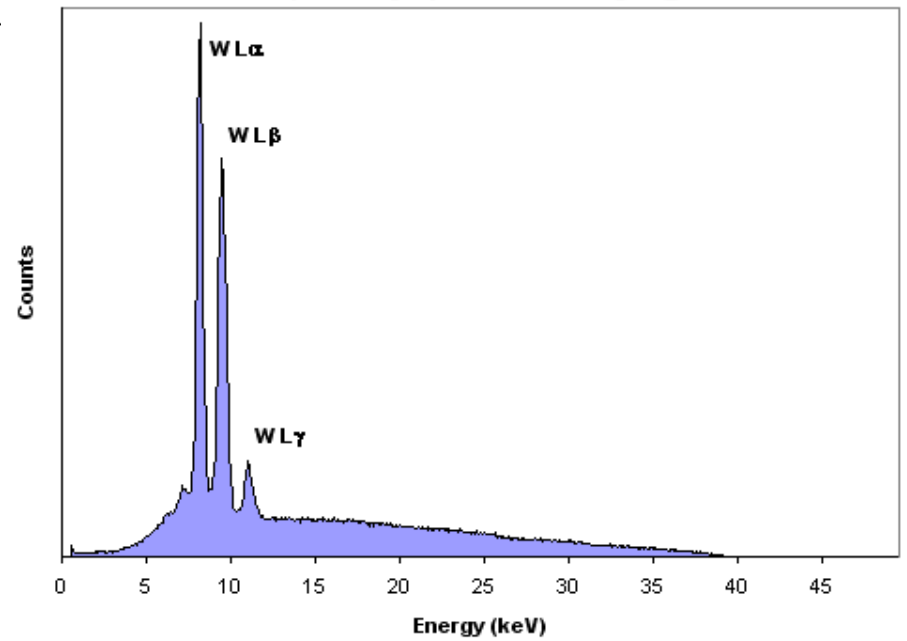


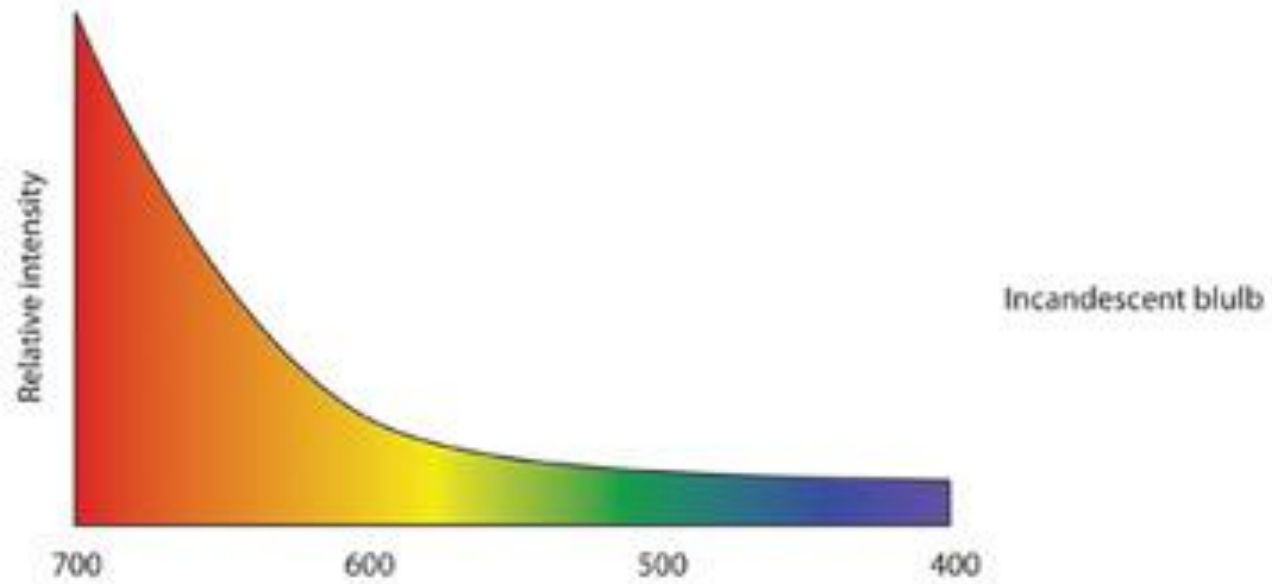
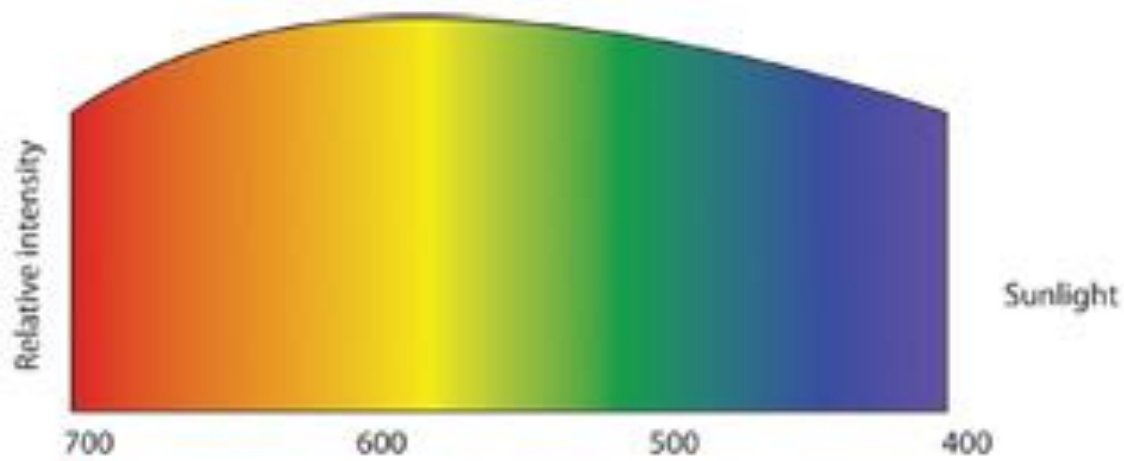


Mini-X Output X-Ray Spectrum: Ag Target @ 40 kV



Mini-X Output X-Ray Spectrum: W Target @ 40 kV





nm

Types of interaction

What happens when one shines matter with X-rays?



Matter

‘It was known that the radiation leaving an X-ray tube was heterogeneous and dependent on the material of the anti-cathode....

When such a homogeneous X-ray beam hits a plate of some element three kinds of *secondary* radiations are emitted....

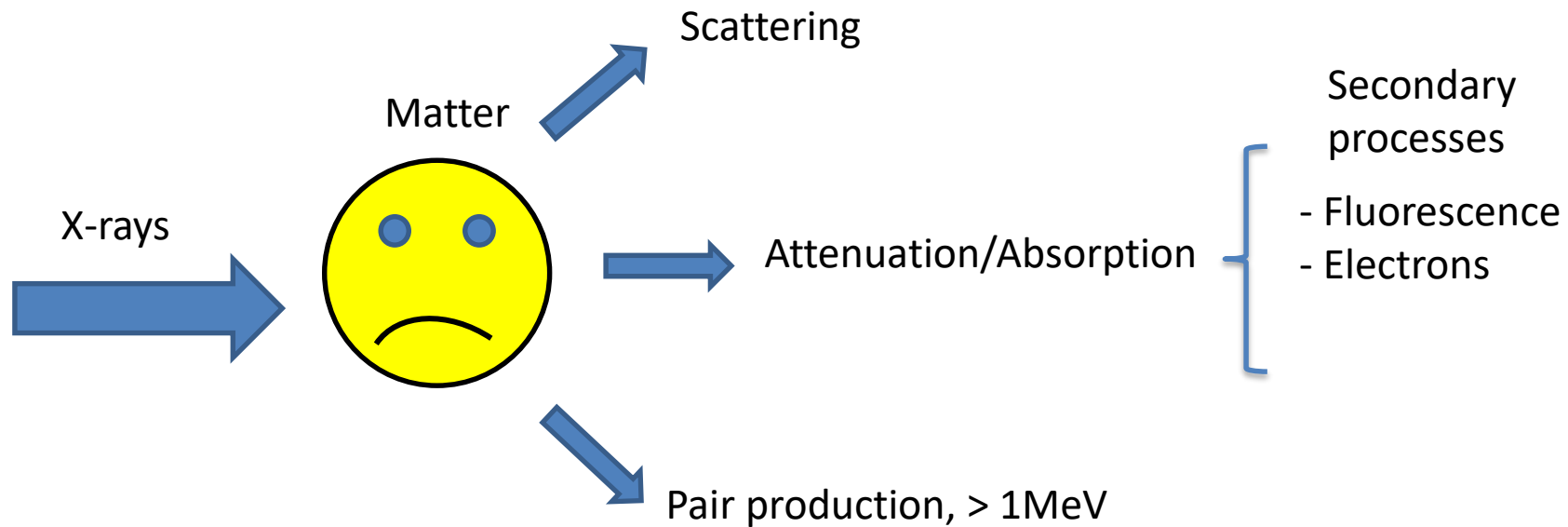
by Manne Siegbahn

- Siegbahn spectrometer, before 1920.

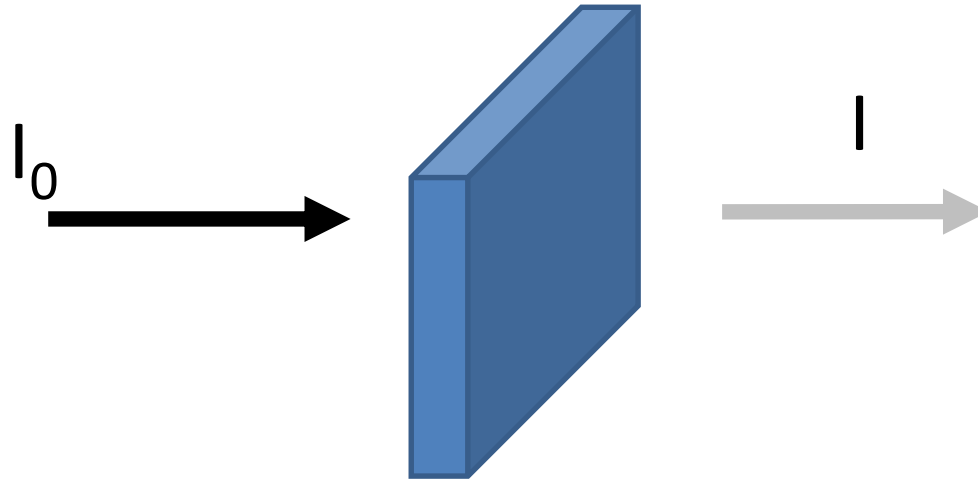
Extract from *50 Years of X-ray Diffraction*, edited by P. P. Ewald

Types of interaction

- What happens when one shines matter with X-rays?



Atenuation = Absorption + Elastic Scattering + Inelastic Scattering



How important are these phenomenon?

τ Photoelectric absorption

σ_{coh} Elastic scattering

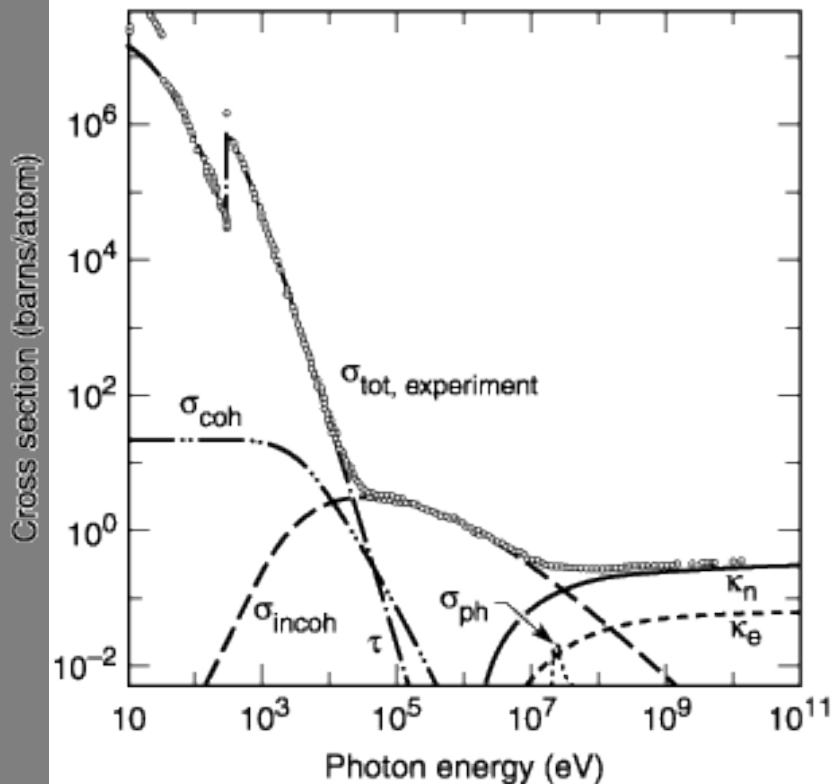
σ_{incoh} Inelastic scattering

κ_e Pair production in electron field

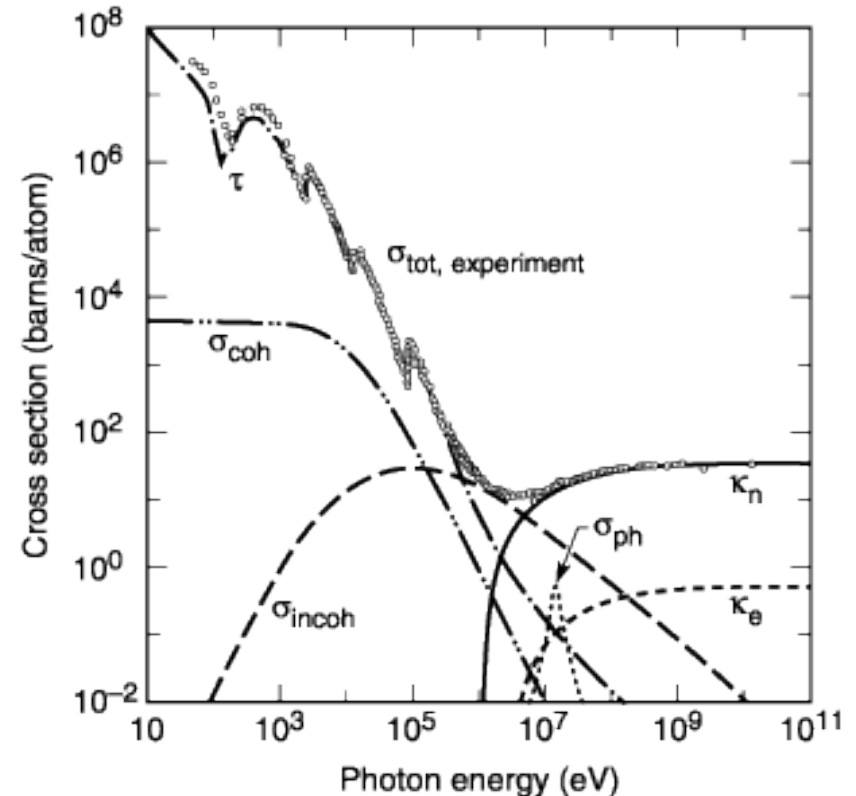
κ_n Pair production in nuclear field

σ_{ph} Photonuclear absorption

Carbon

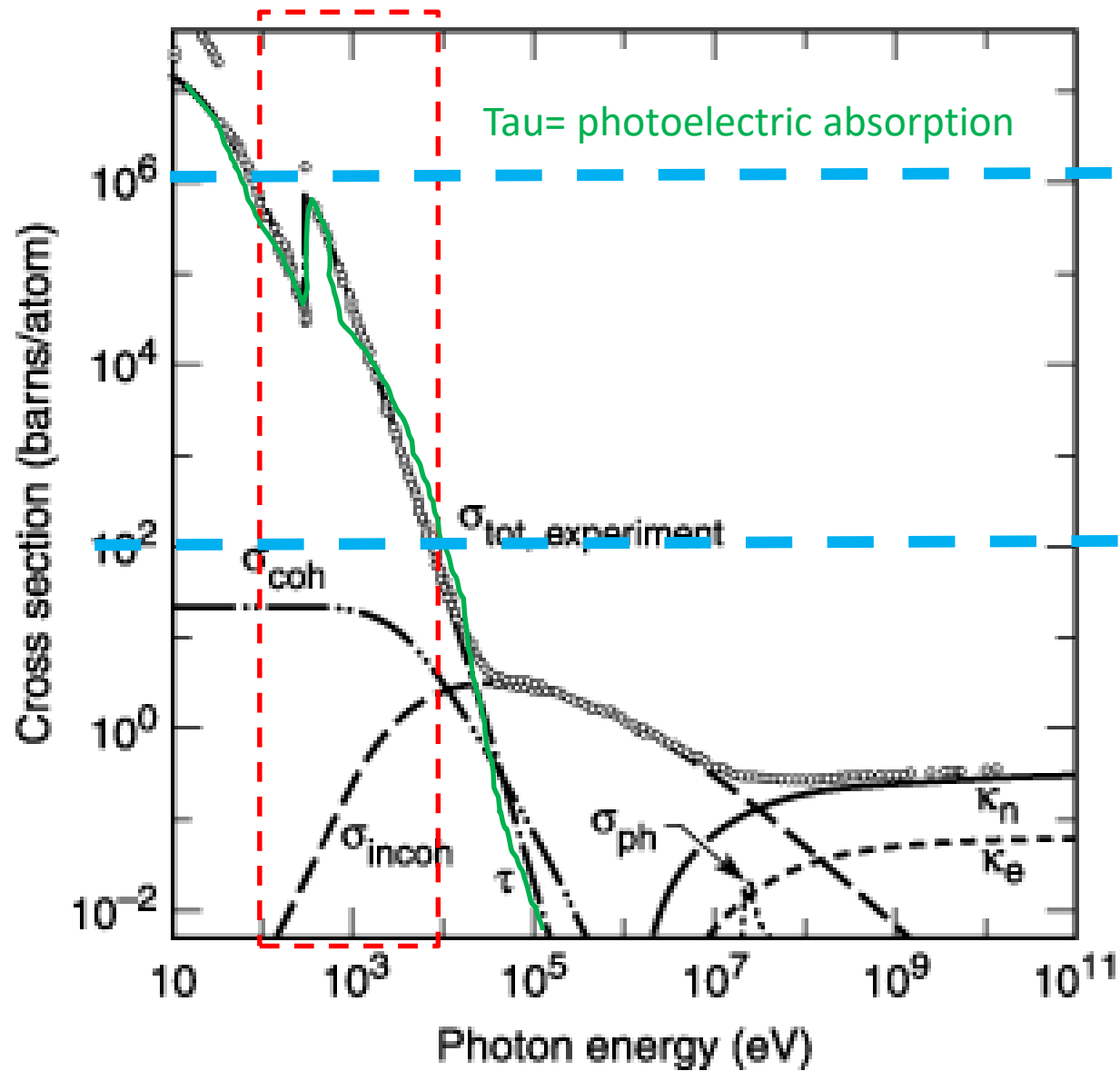


Lead

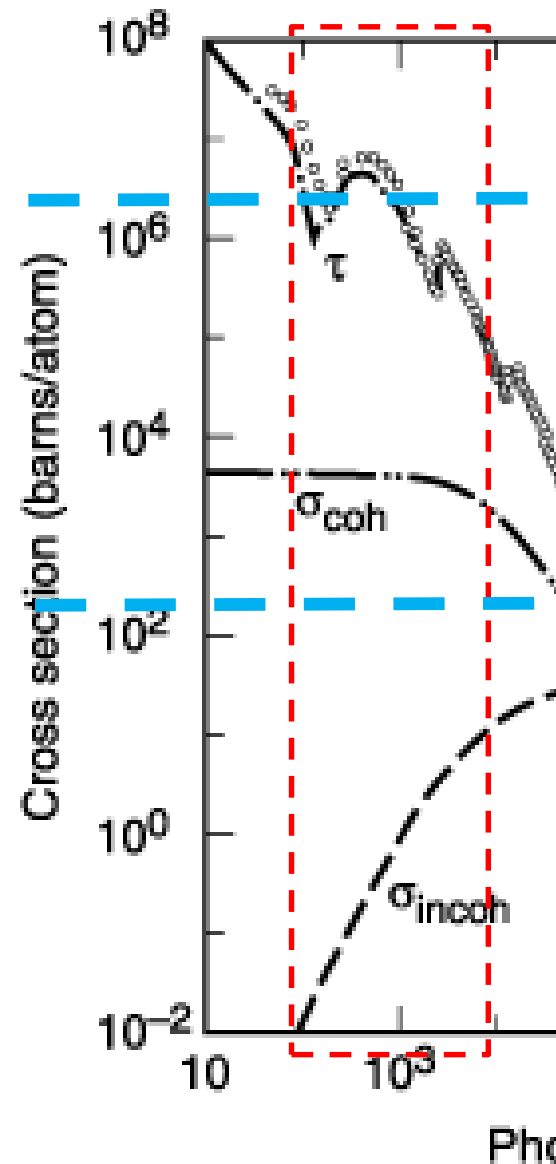


Attenuation

Organic sample



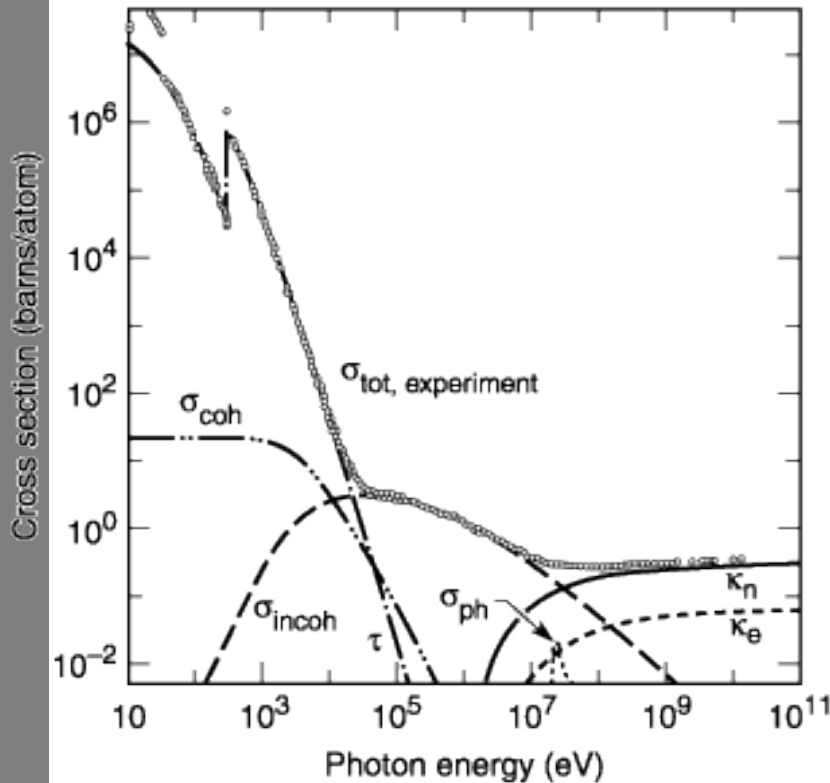
Lead



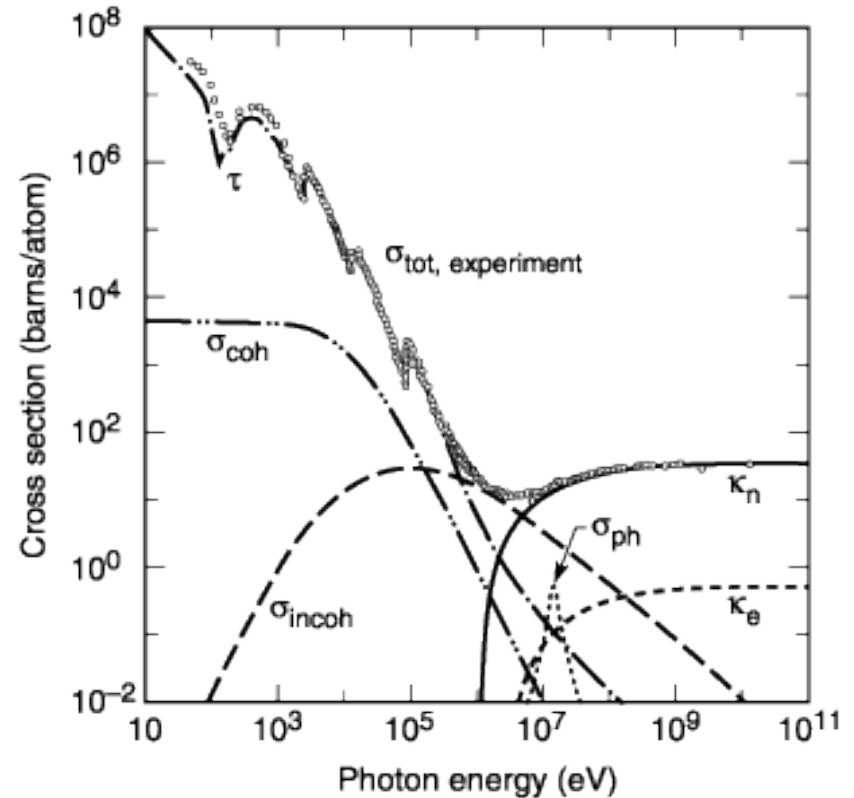
How important are these phenomenon?

- ❑ What singularities one may capture from C and Pb?
- ❑ Why has C only one sharp increase of τ whereas Pb has several?

Carbon



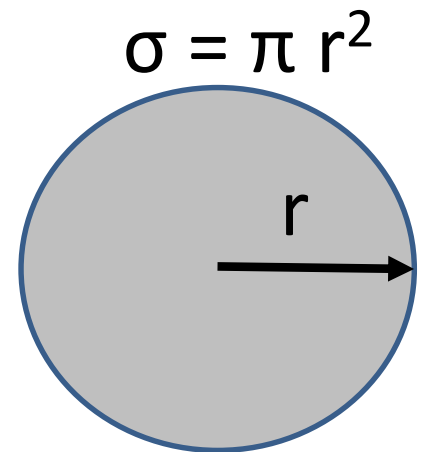
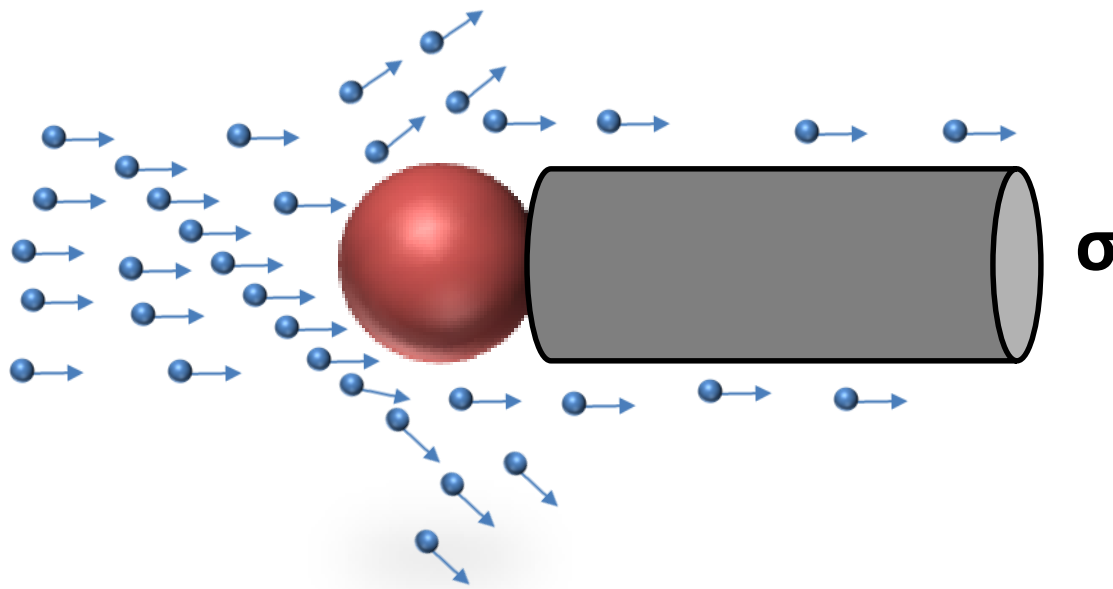
Lead



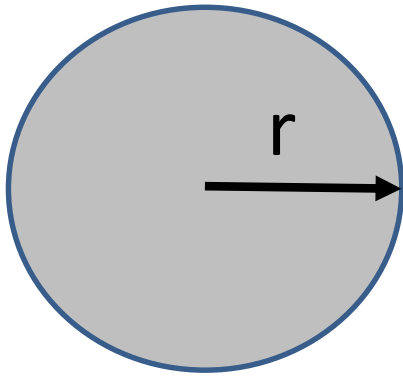
5.2 Cross sections

Cross sections

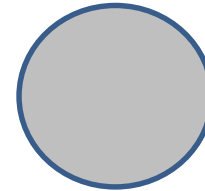
- ❑ Is the effective area that express the probability of an scattering event to occur
- ❑ The higher the cross section, the higher is the probability
- ❑ The cross section is expressed in m^2 or in Barn = $10^{-28} m^2$



$$\sigma = \pi r^2$$

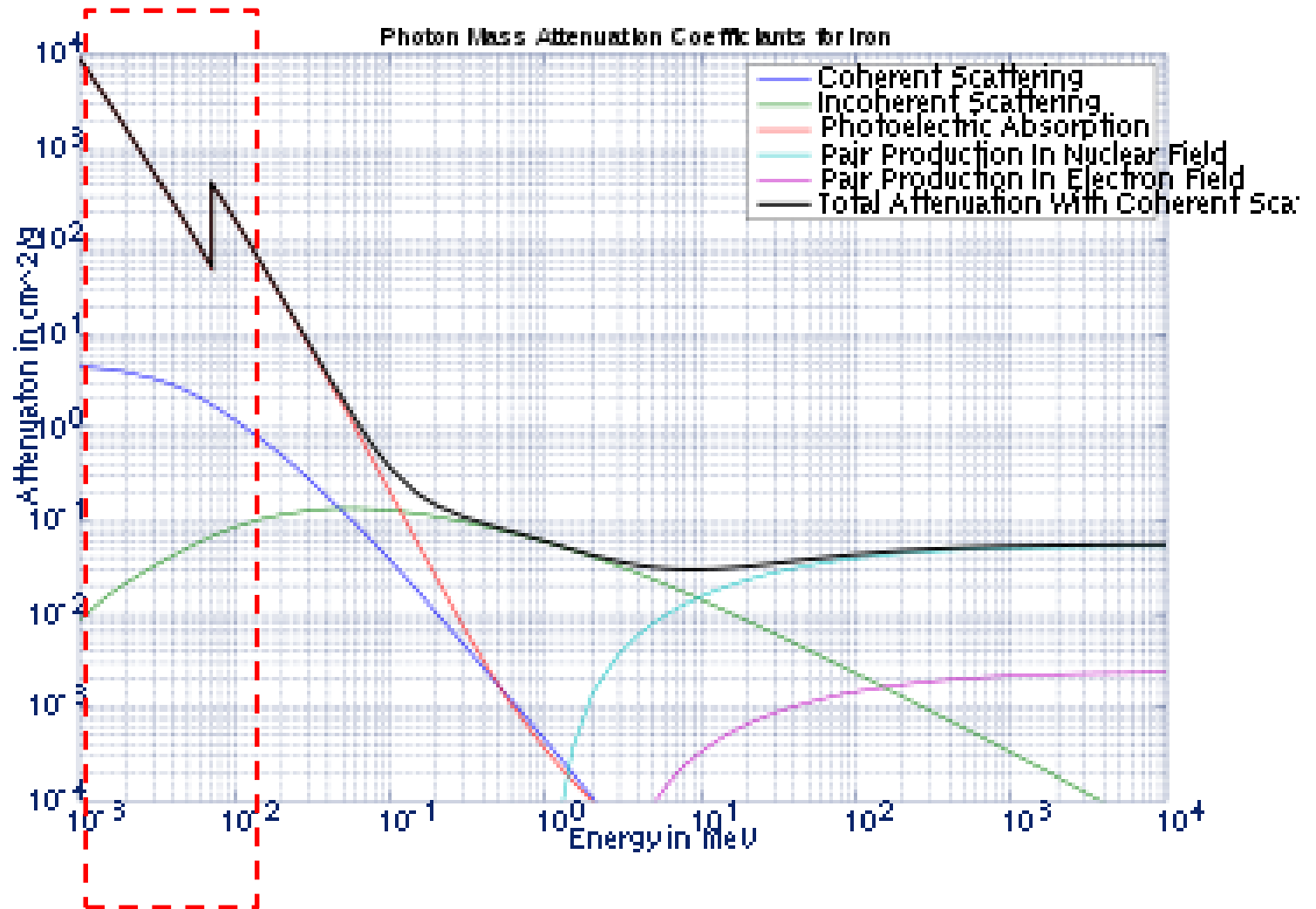


1740 eV
Si K edge



7112 eV
Fe K edge

Cross section for Copper



https://upload.wikimedia.org/wikipedia/commons/thumb/e/ed/Attenuation_Coefficient_Iron.svg/400px-Attenuation_Coefficient_Iron.svg.png

Cross sections and atomic raddii

□ Attenuation cross sections represent how 'large' is atom from the photon point of view

Atomic raddii in 10^{-12} m

http://images.flatworldknowledge.com/averillfwk/averillfwk-fig07_007.jpg

X-ray attenuation image

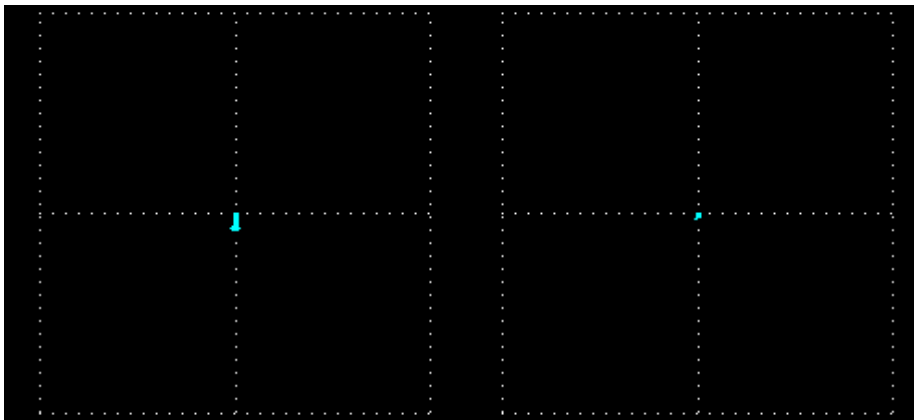
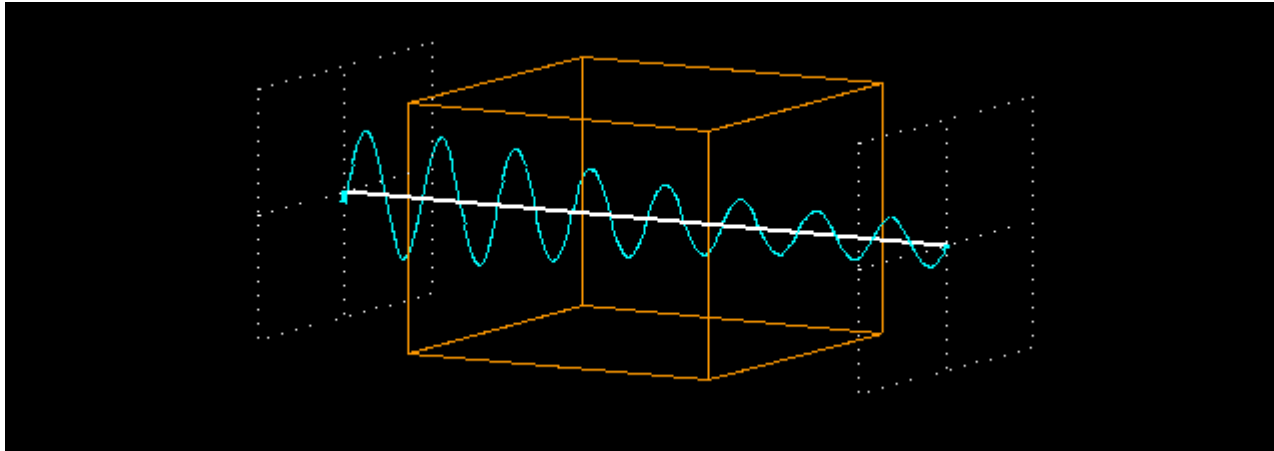


<http://www.sbcc.edu/careercenter/images/Radiography%20-%20Hands.jpg>

5.3 X-ray photoelectric absorption and relaxation processes

Absorption of a polarized

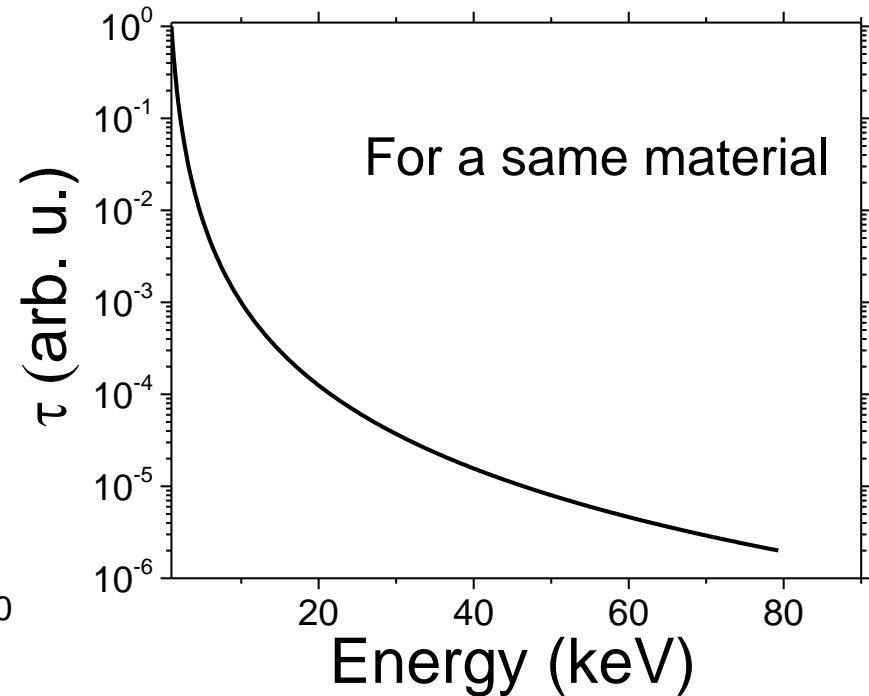
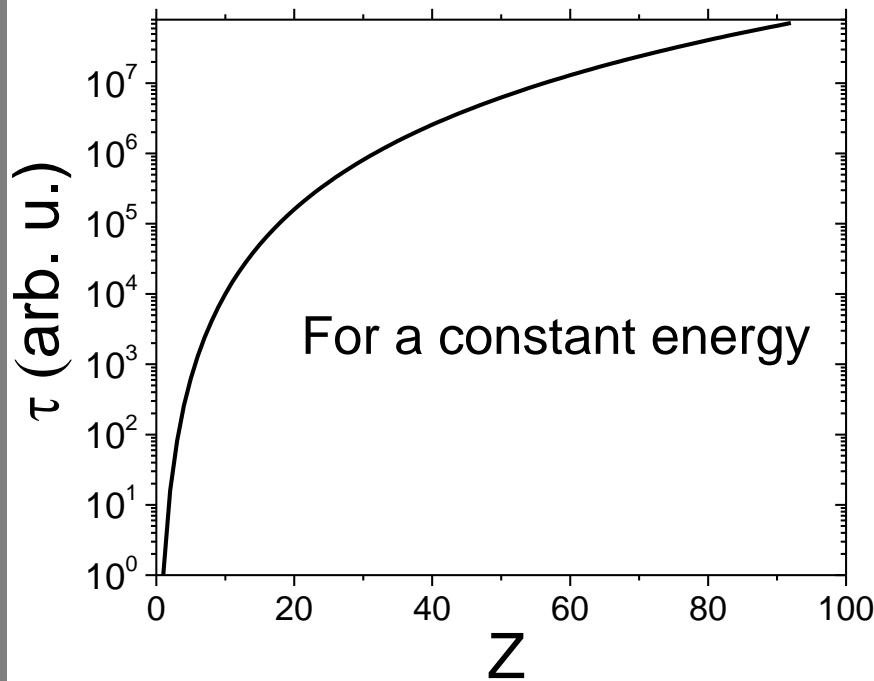
- The reduction in the number of photons is represented as a decrease in the amplitude of the wave



<http://cddemo.szialab.org/>

X-ray photoelectron absorption

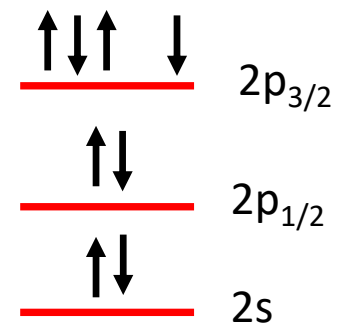
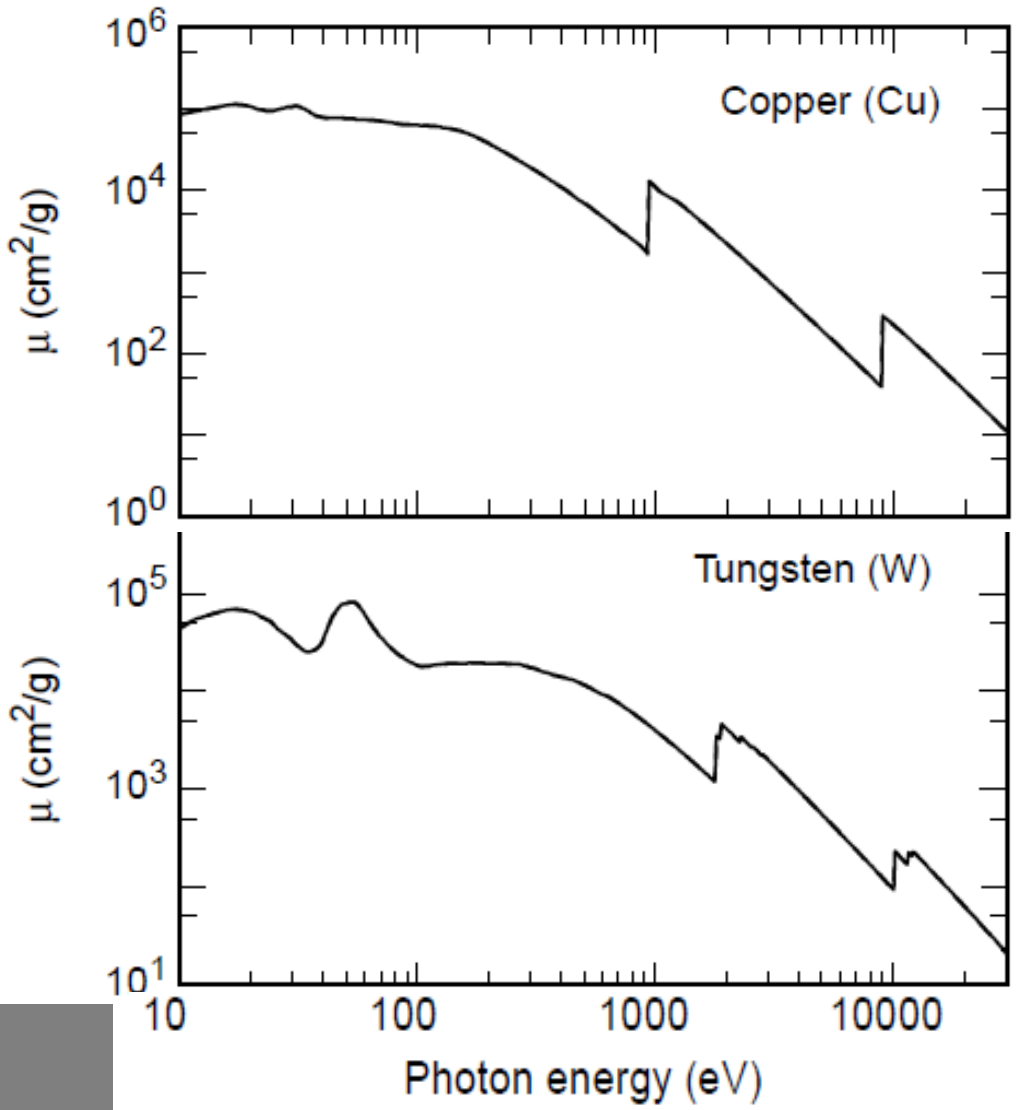
- The photoelectron absorption increases as function of Z and decreases as function of E



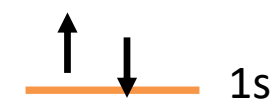
Attenuation cross section discontinuities

□ What are these sharp increases in μ ?

Cu

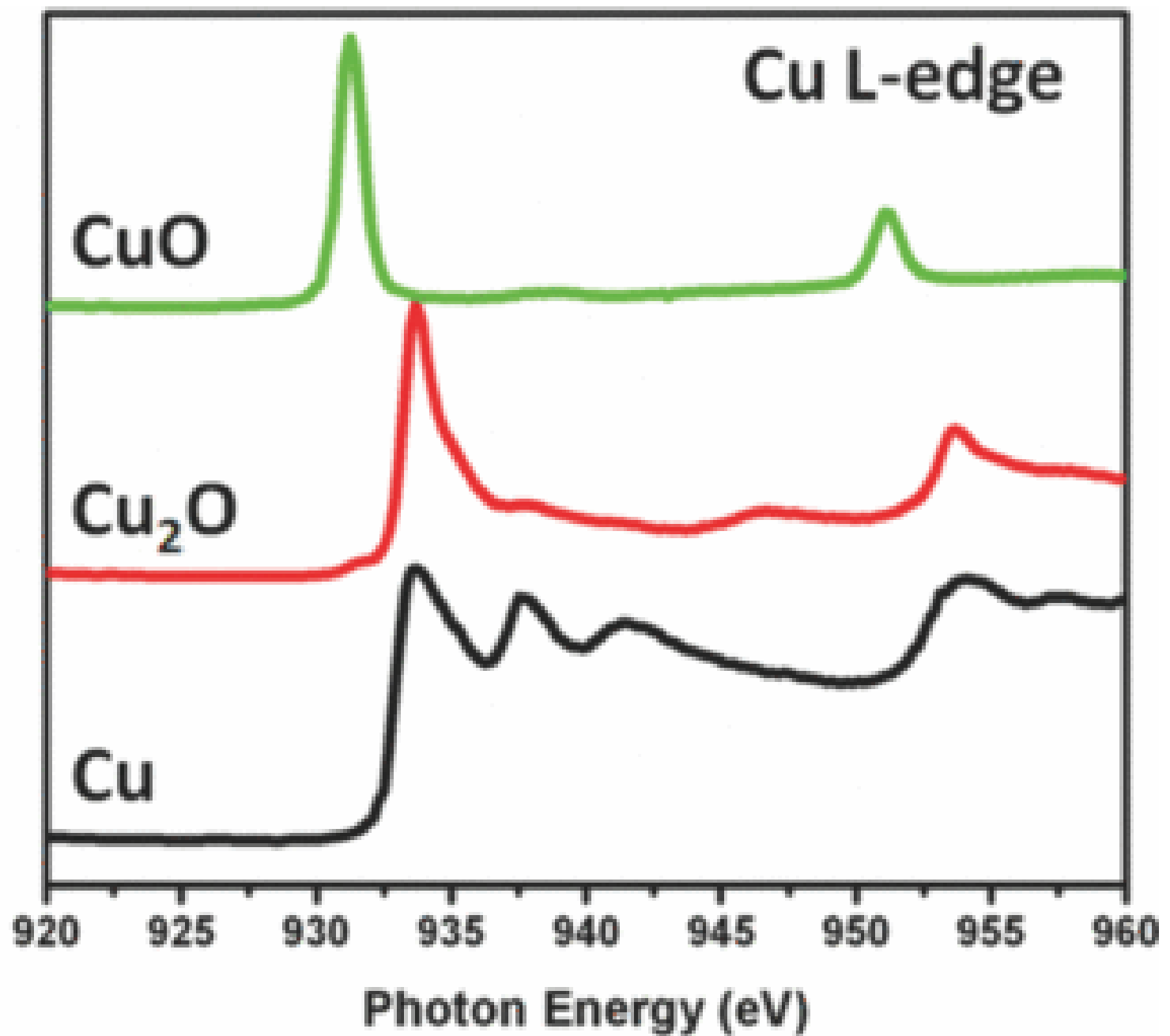


L edges



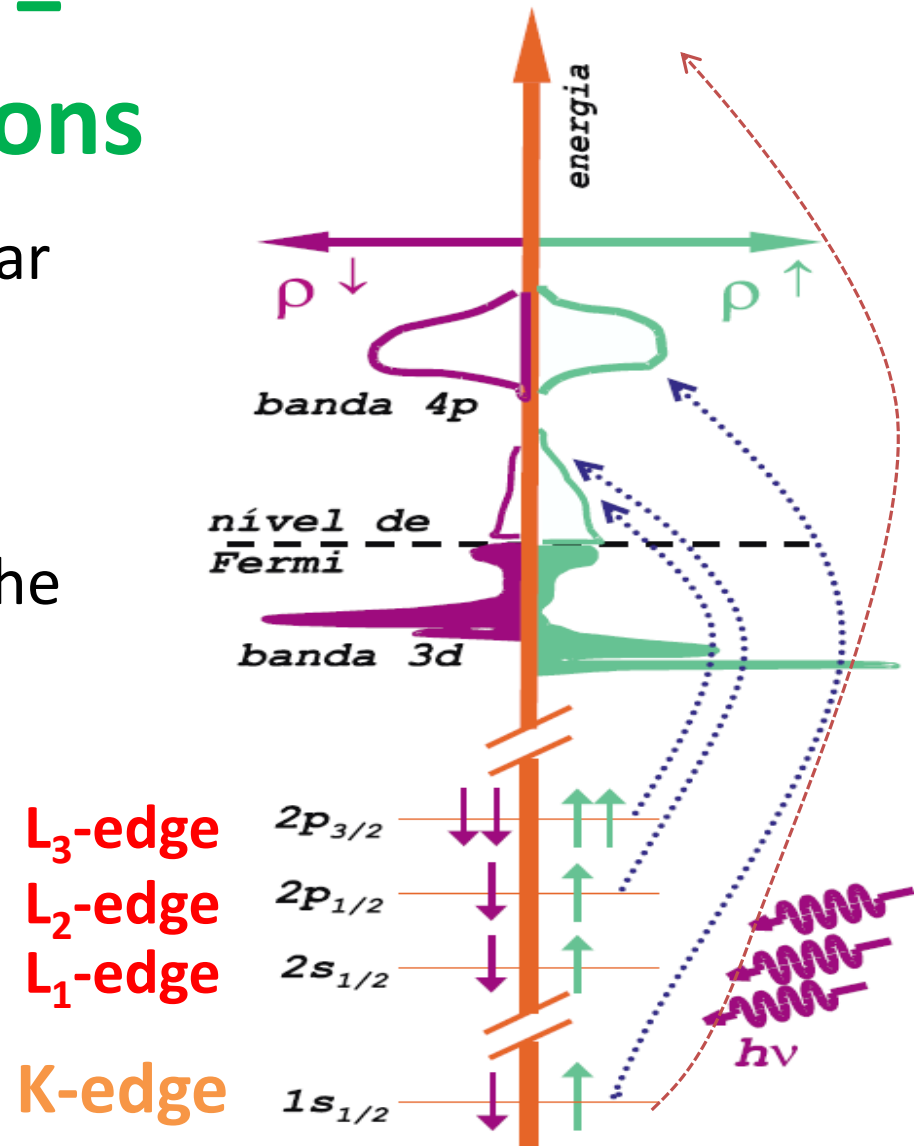
K edge

Cu L-edge

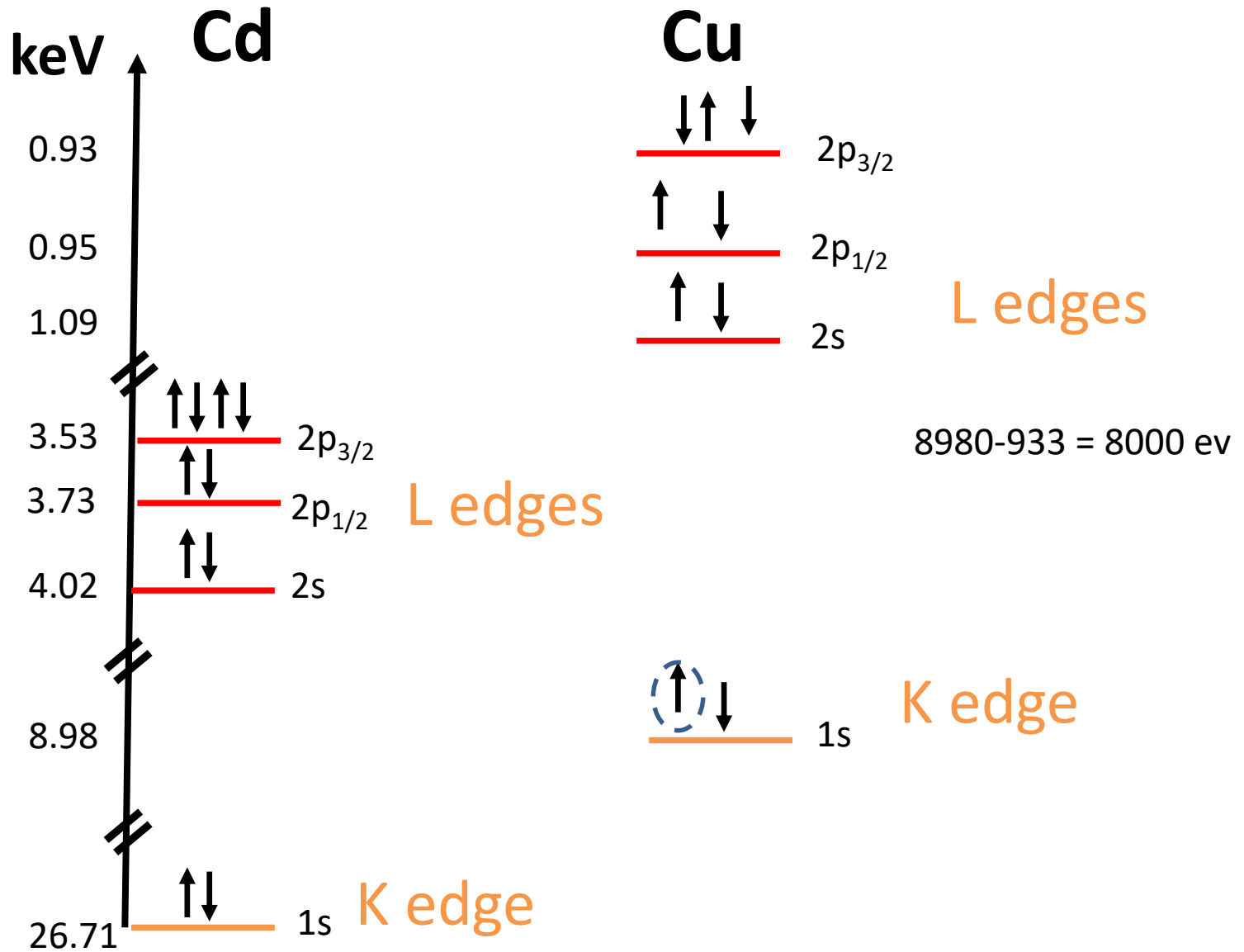


Absorption edges = Electronic transitions

- The orbitals have particular energies
- When the photon has energy slightly higher than the orbital's, the probability of interaction increases



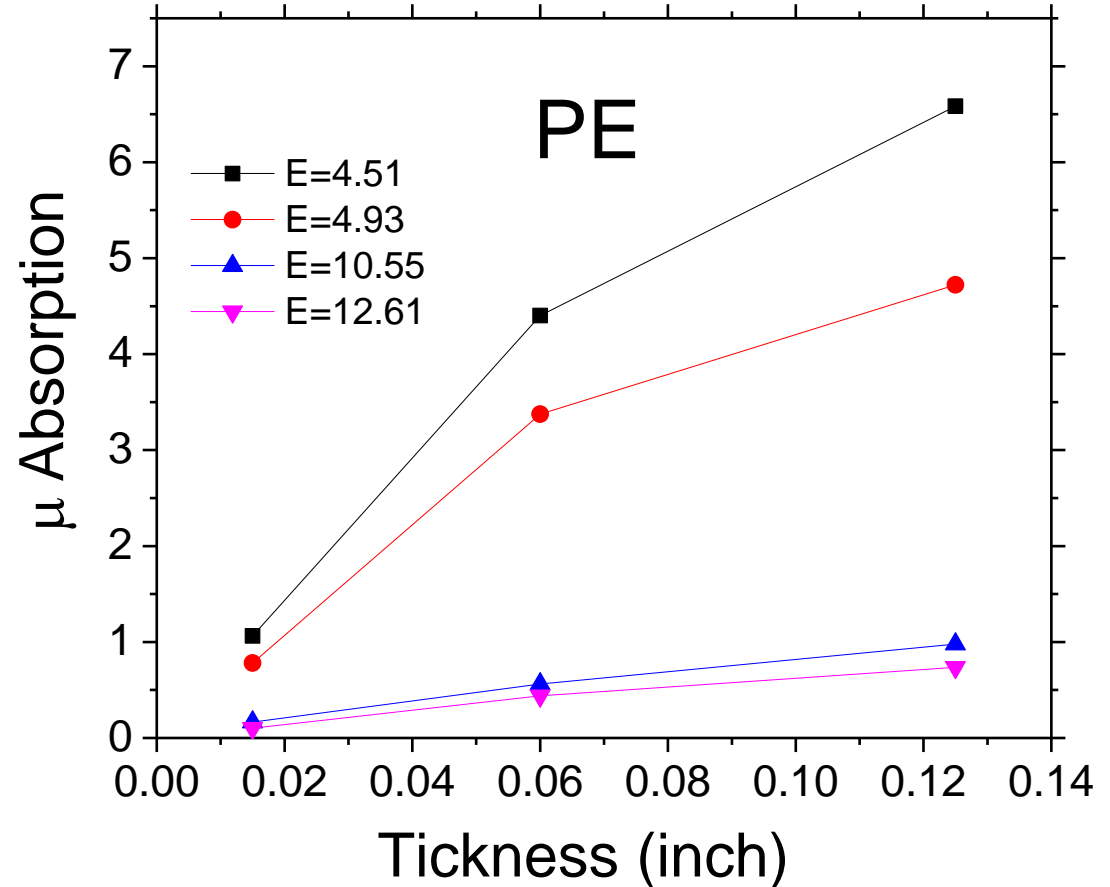
Cd and Cu absorption edges



Absorption as function of thickness

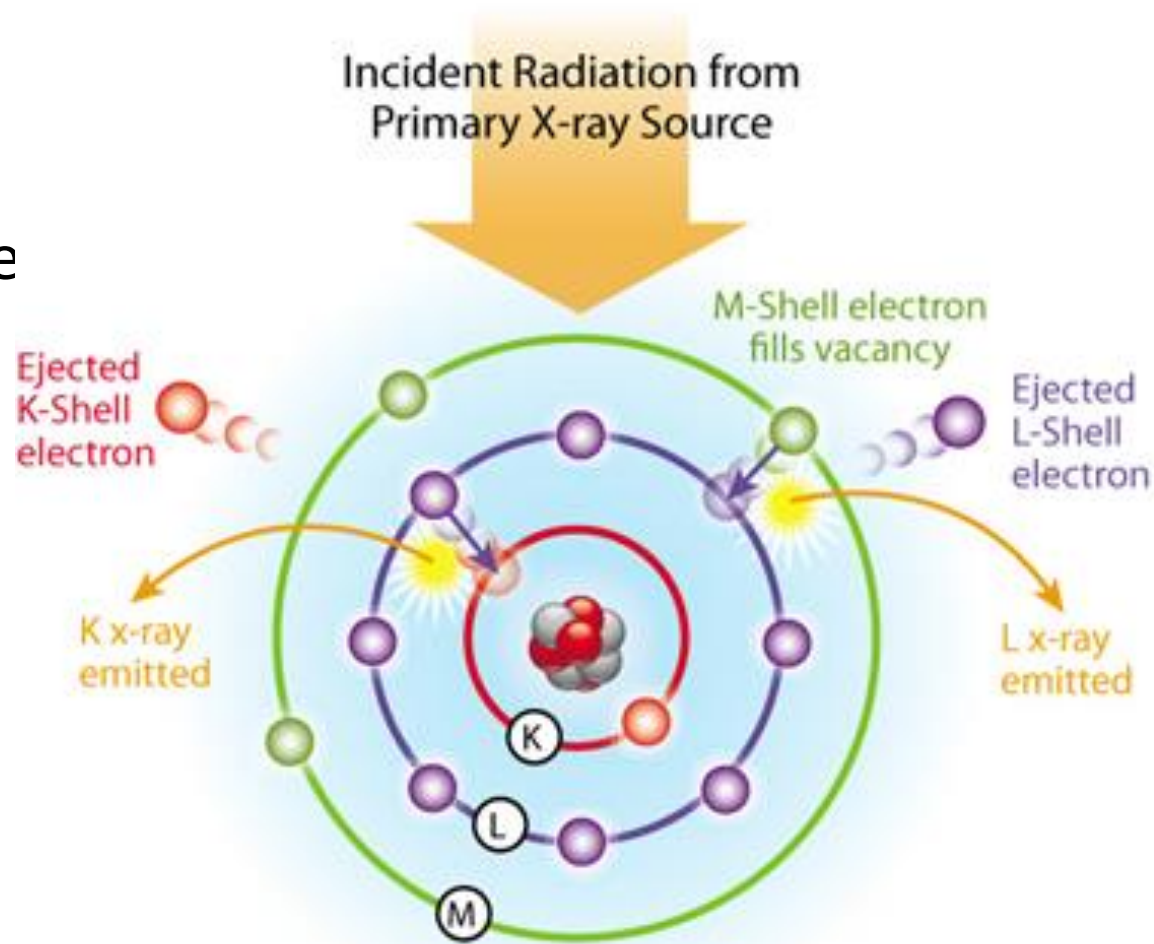
X-ray attenuation for a polyethylene sample

Paramos aqui
28/10/2022



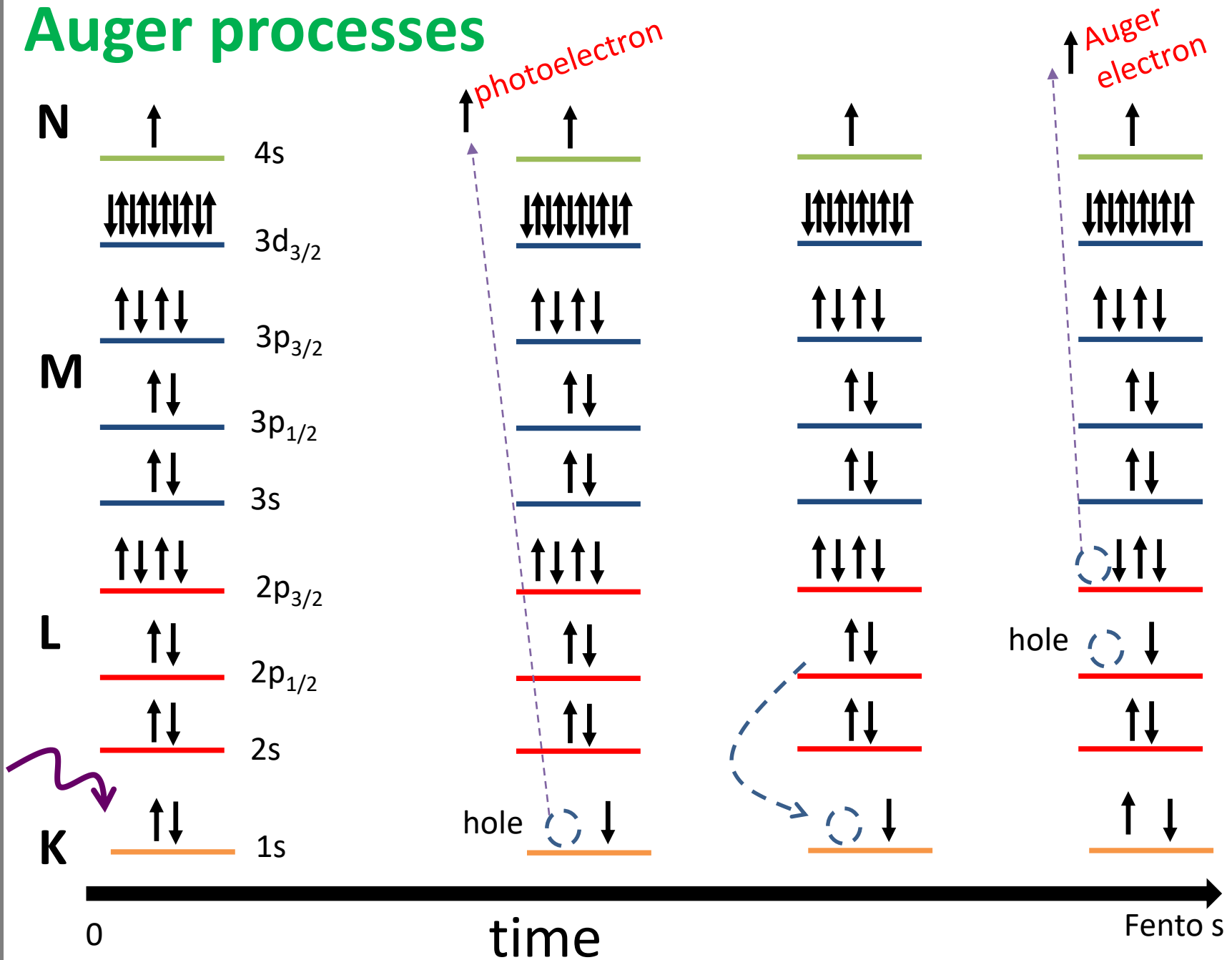
Relaxation processes

- ❑ Auger emission
- ❑ X-ray fluorescence

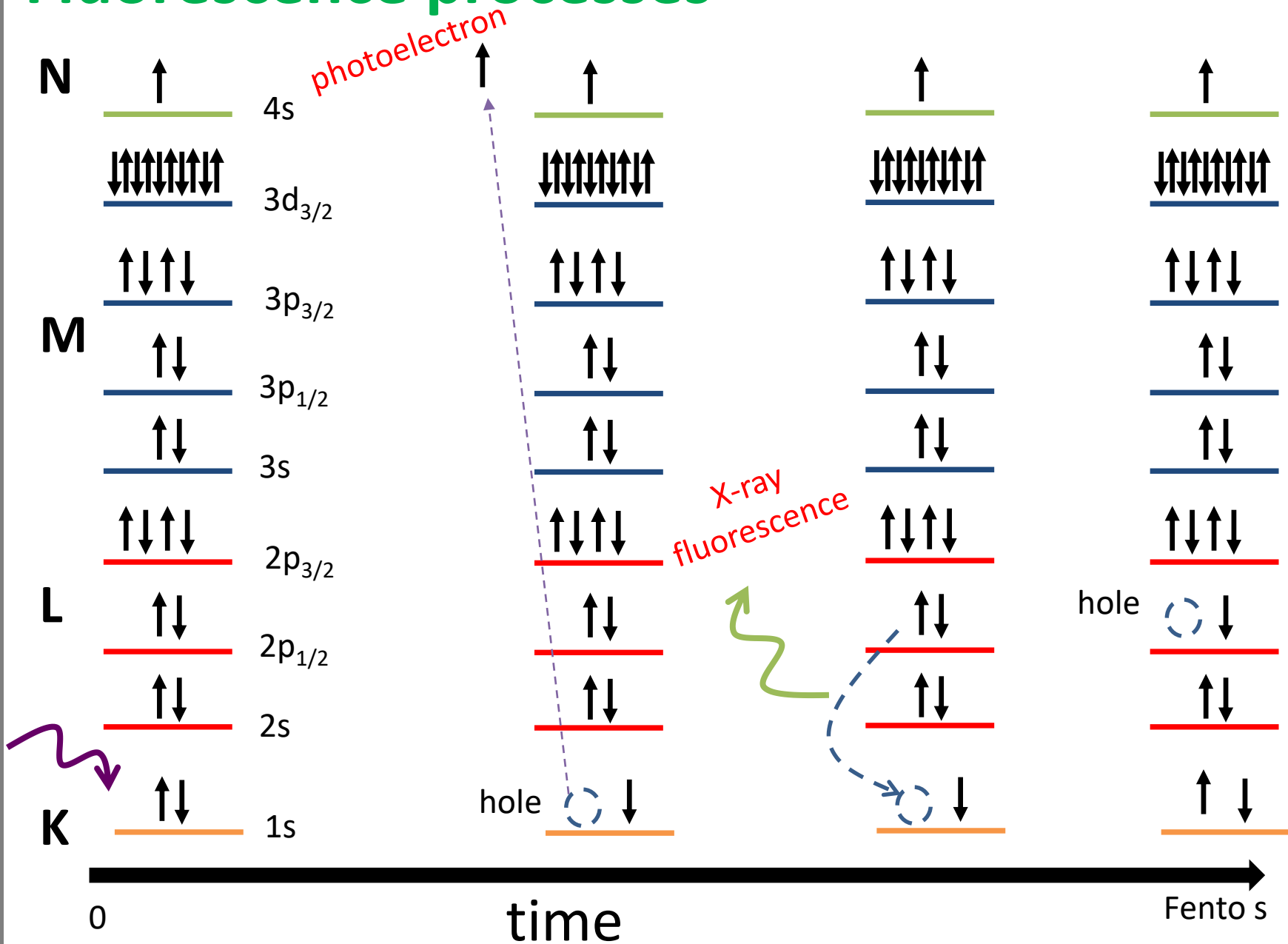


<https://wiki.utep.edu/download/attachments/67371770/XRF%20fig%200.png?version=1&modificationDate=1384724989903&api=v2>

Auger processes

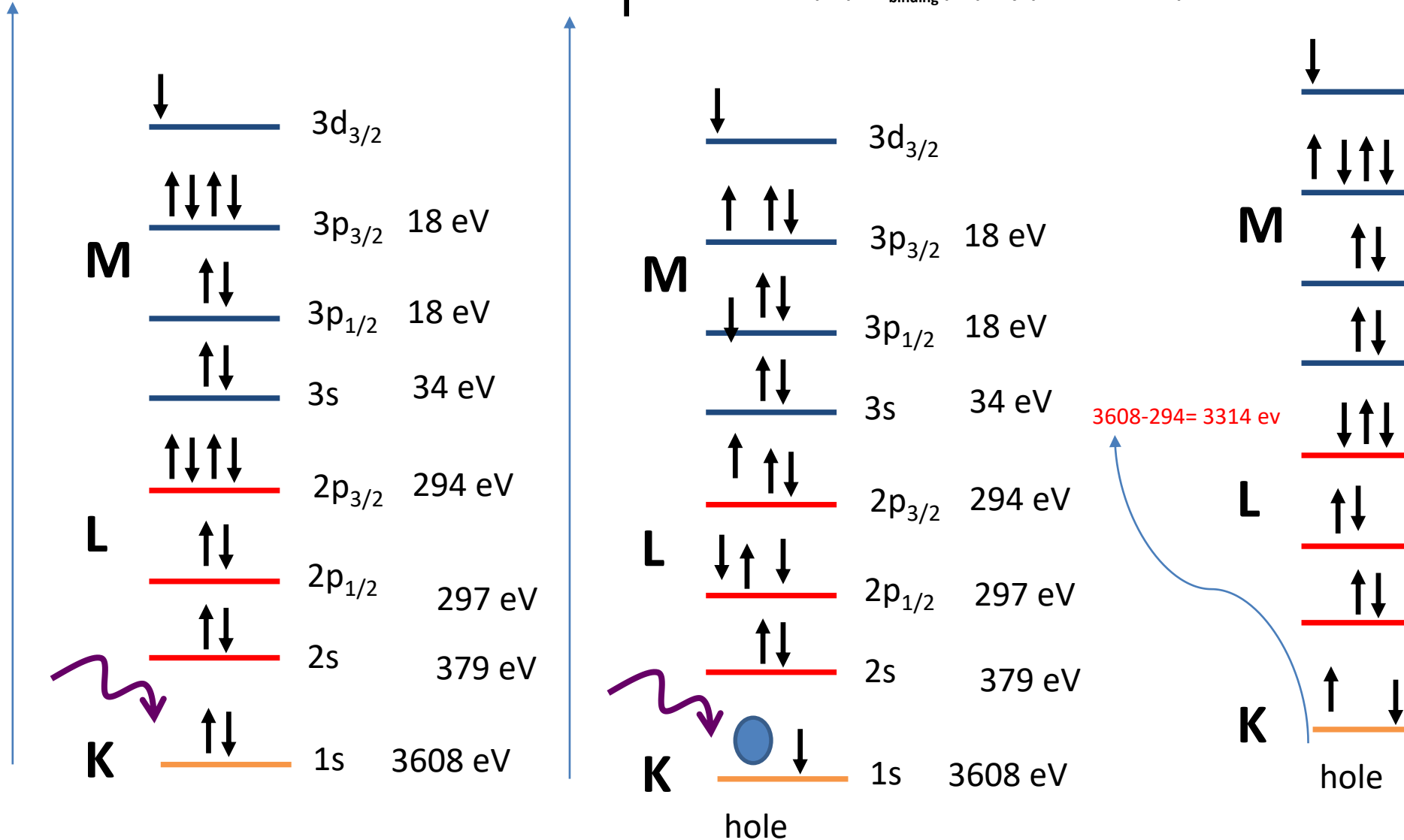


Fluorescence processes



Potassium atom

$$\uparrow KE = \frac{1}{2}mv^2 = h\nu \text{ (eV)} - E_{\text{binding}} \text{ (eV)} - \phi \text{ (work function)}$$



X-ray Absorption and Emission Energies of the Elements

Hydrogen 1																		Helium 2																																	
1.0079																		4.0026																																	
Li 3																		Be 4		B 5		C 6		N 7		O 8		F 9		Ne 10																					
6.941																		9.0122		10.81		12.011		14.0067		15.9994		18.9984		20.1797																					
Na 11																		Mg 12		Al 13		Si 14		P 15		S 16		Cl 17		Ar 18																					
22.9898																		24.305		26.9815		28.0855		30.9738		32.06		35.453		39.948																					
K 19																		Ca 20		Sc 21		Ti 22		V 23		Cr 24		Mn 25		Fe 26		Co 27		Ni 28		Cu 29		Zn 30		Ga 31		Ge 32		As 33		Se 34		Br 35		Kr 36	
39.0983																		40.08		44.9559		47.88		50.9415		51.996		54.938		55.847		58.9332		58.69		63.546		65.38		69.72		72.59		74.9216		78.96		81.904		83.8	
Rb 37																		Sr 38		Y 39		Zr 40		Nb 41		Mo 42		Tc 43		Ru 44		Rh 45		Pd 46		Ag 47		Cd 48		In 49		Sn 50		Sb 51		Te 52		I 53		Xe 54	
85.4678																		87.62		88.9059		91.22		92.9064		95.94		97.907		101.07		102.906		106.42		107.868		112.412		114.82		118.69		121.75		127.6		129.905		131.29	
Cs 55																		Ba 56		La 57		Hf 58		Ta 59		W 60		Re 61		Os 62		Ir 63		Pt 64		Au 65		Hg 80		Tl 81		Pb 82		Bi 83		Po 84		At 85		Rn 86	
132.905																		137.33		138.906		178.49		180.948		183.85		186.207		190.2		192.22		195.08		196.967		200.59		204.38		207.2		208.982		209.987		222.018			
Fr 87																		Ra 88		Ac 89																															
1013.37																		1063.00		1064.83																															
188.39																		190.39		190.39																															
1700.7																		1700.7		1700.7																															
1501.3																		1501.3		1501.3																															
223.02																		223.02		223.02																															

Atomic Data and Energies from
W. T. Elam, B. D. Ravel and J. R. Sieber,
Radiation Physics and Chemistry 63, pp 121-128 (2002)

Common oxidation states from wikipedia.org, after
N. N. Greenwood and A. Earnshaw,
Chemistry of the Elements, 2nd ed. (1997).

All energies in eV.
Emission line strengths are approximate, and vary with element.

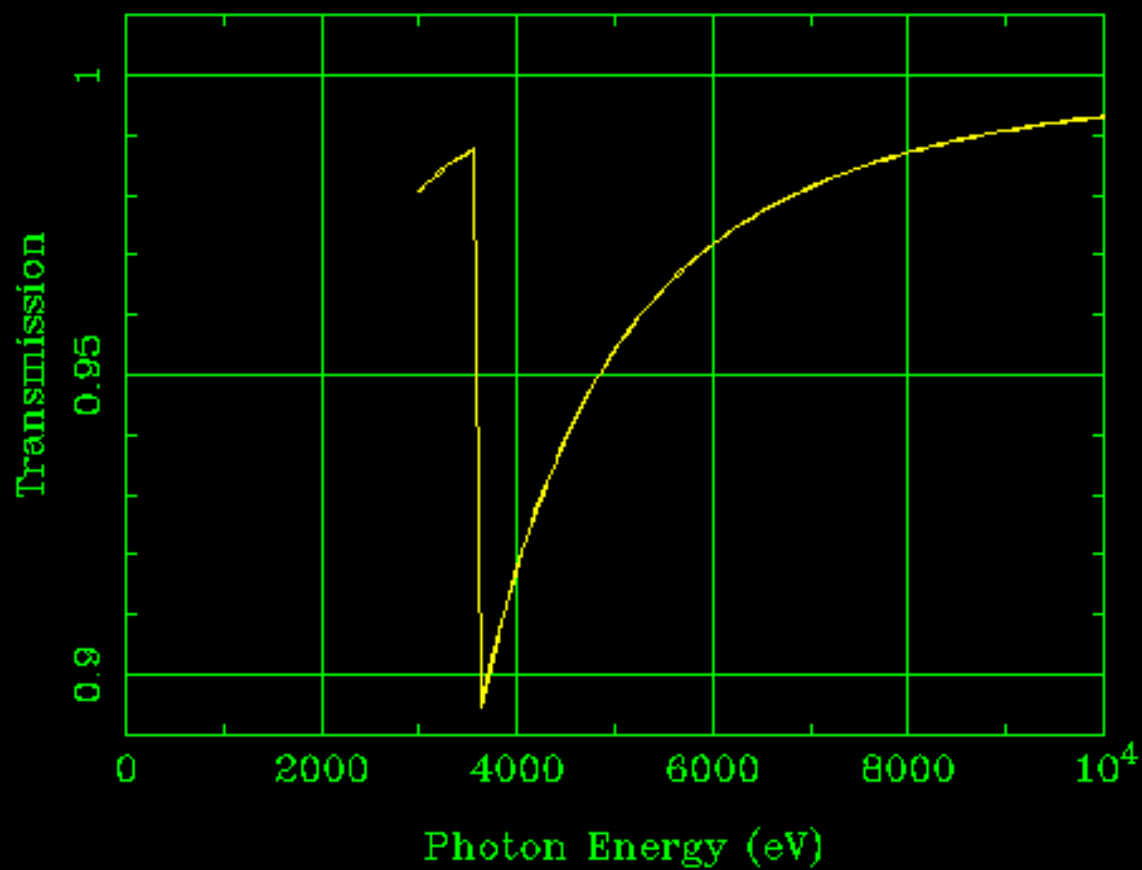


https://xrayabsorption.org/xraytable
Version 4, 2020-April-19

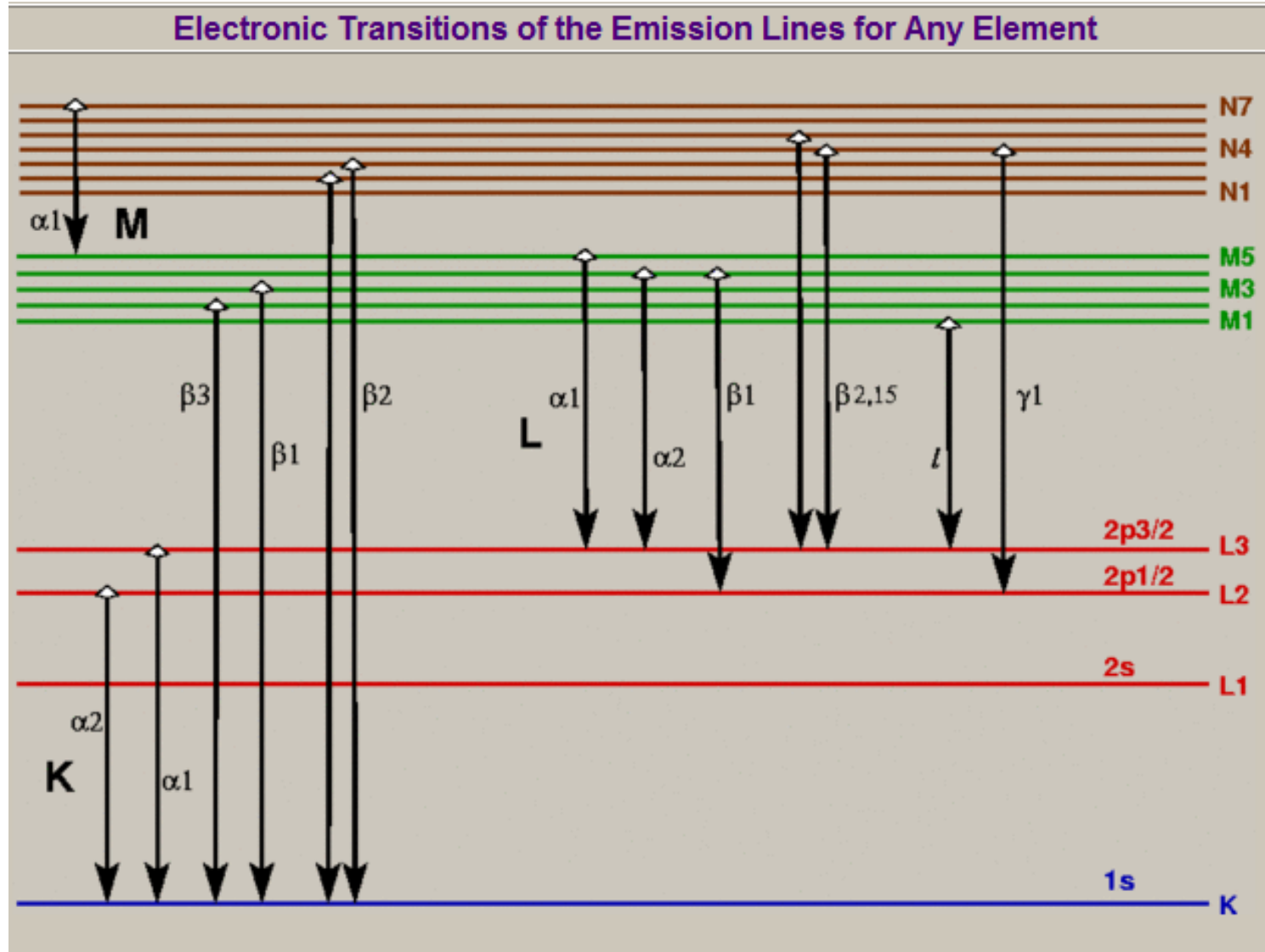


Marie Skłodowska Curie

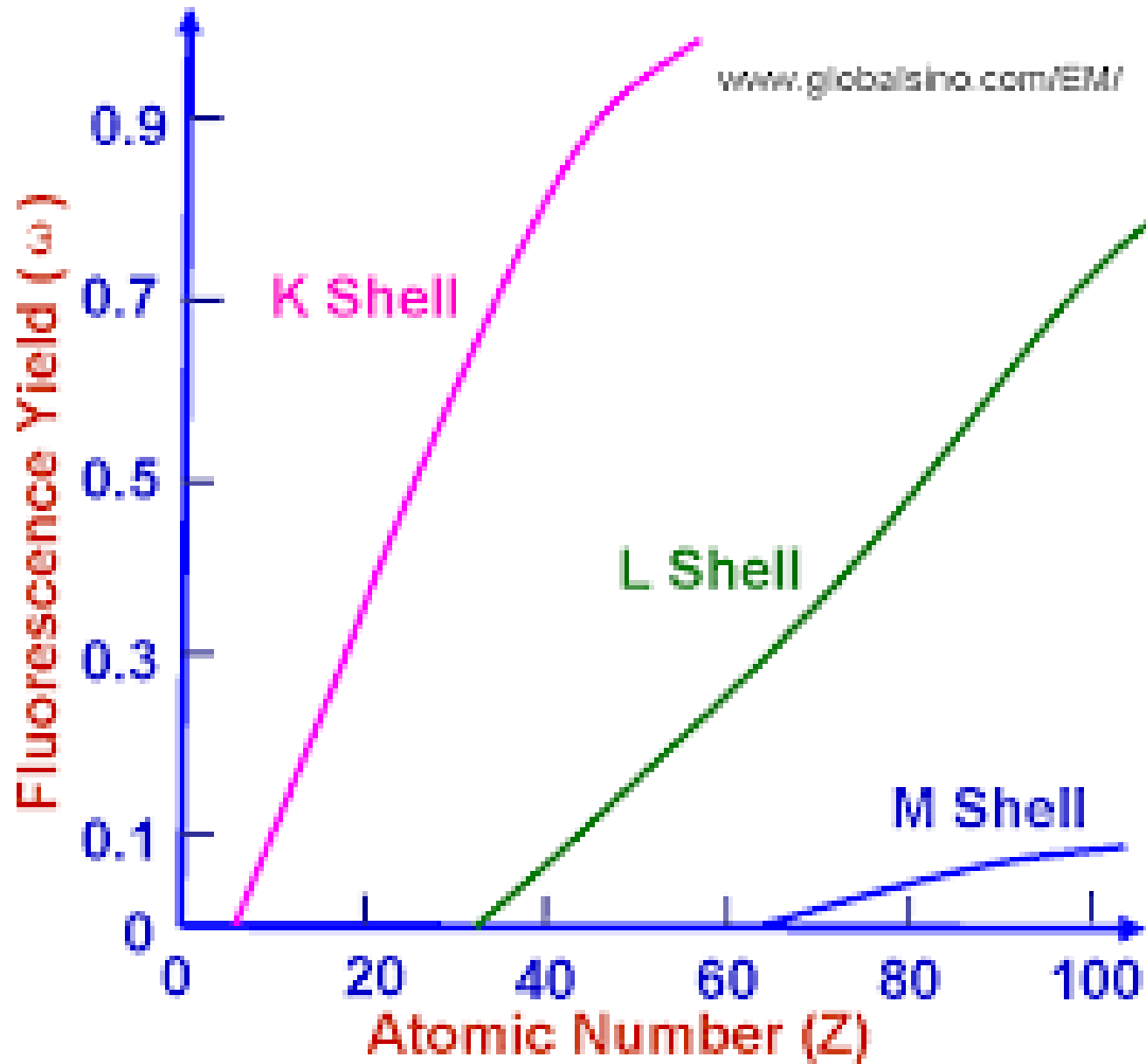
K Density=0.862 Thickness=1. microns



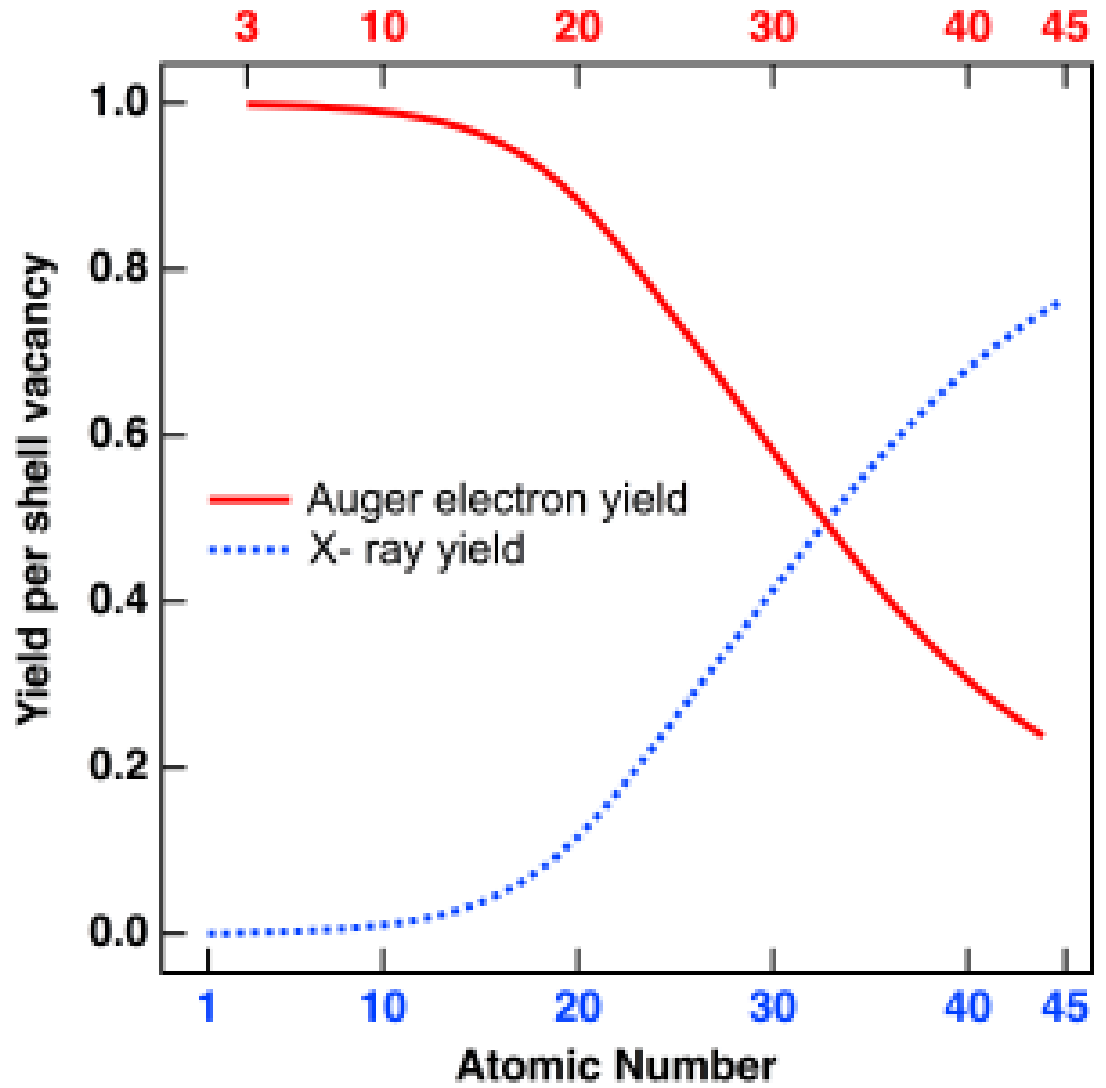
XRF Relaxation processes



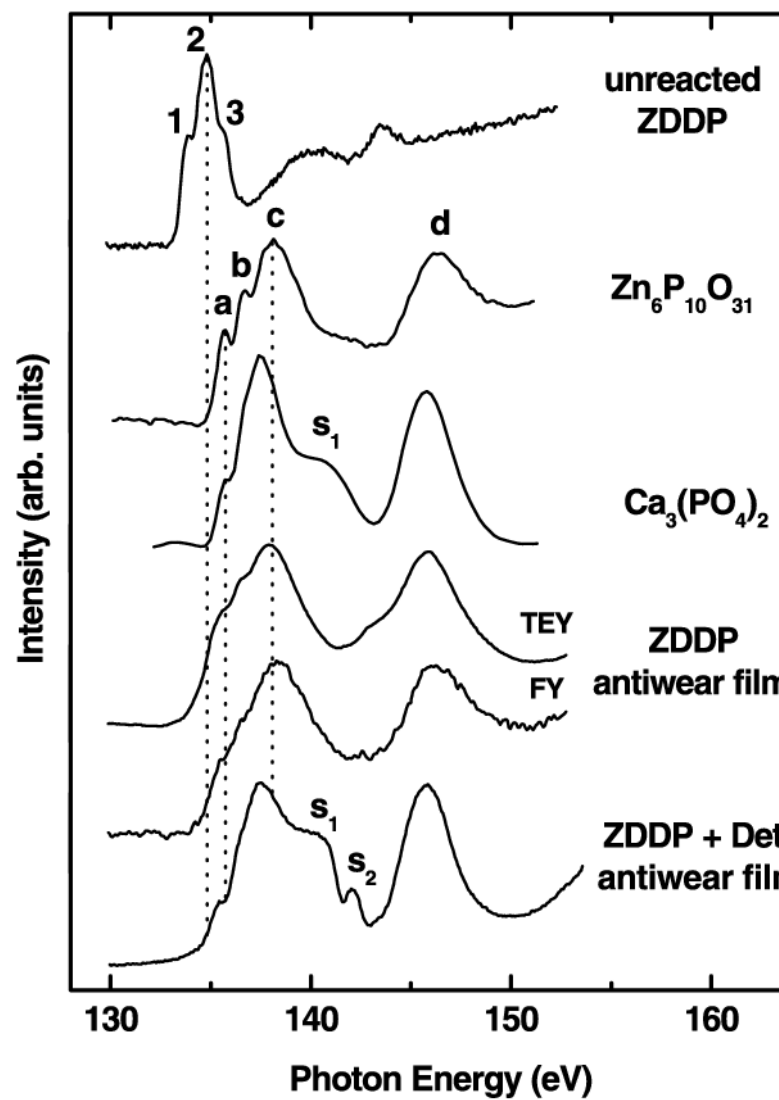
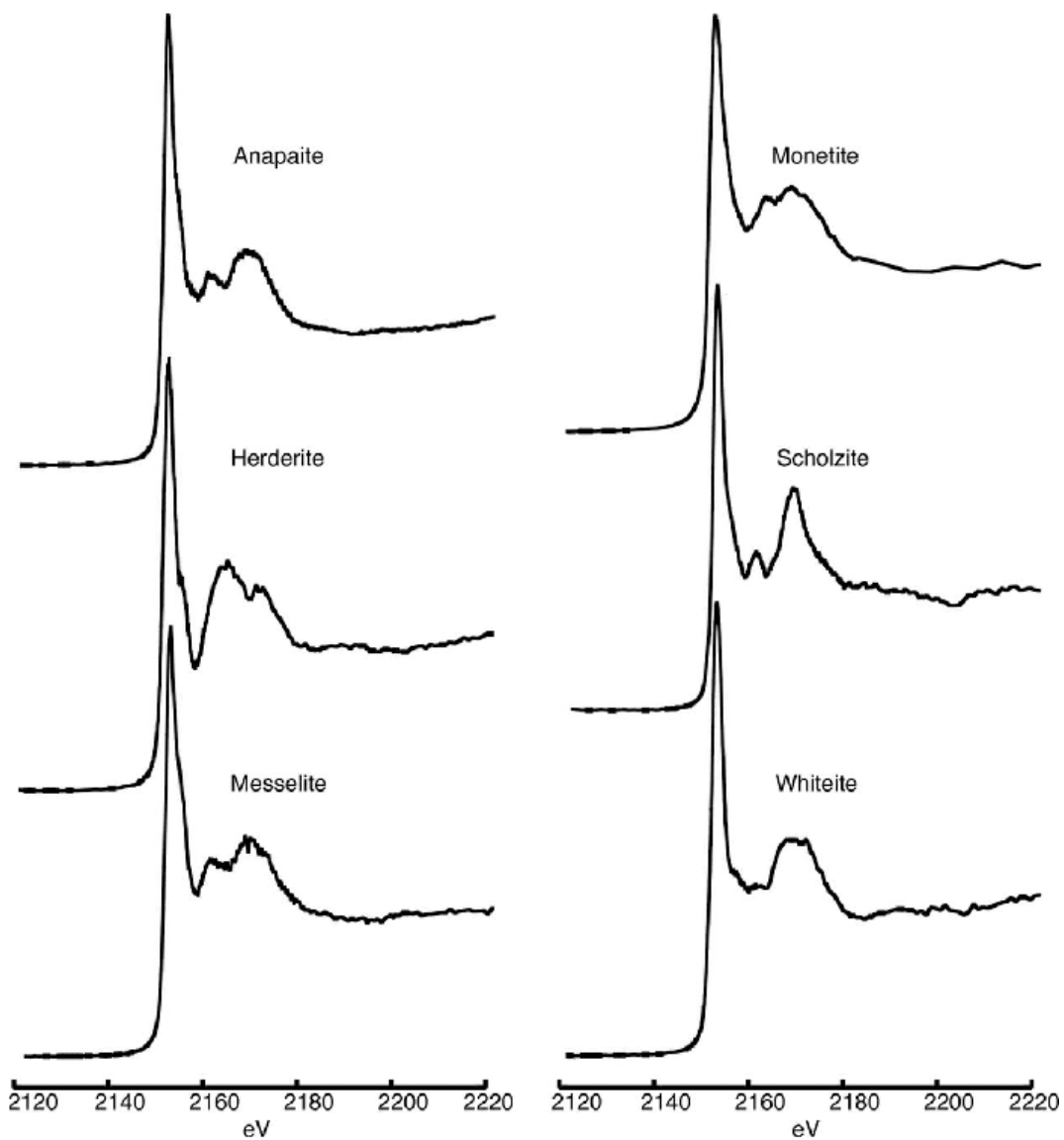
Hephasestus, By B. Ravel



XRF versus Auger competition



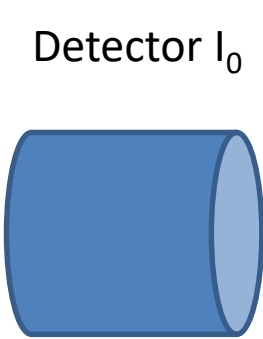
https://upload.wikimedia.org/wikipedia/commons/thumb/5/54/Auger_xray_wiki_in_png_format.png/340px-Auger_xray_wiki_in_png_format.png



Monochromatic
incident radiation
10,000 eV



1000
photons

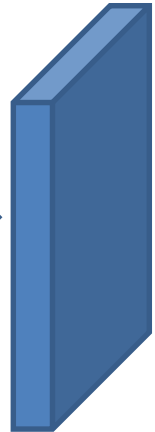


Detector I_0



I_0

1000
photons

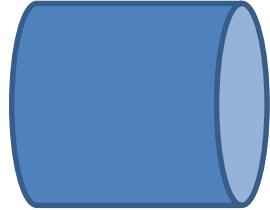


Material copper



I

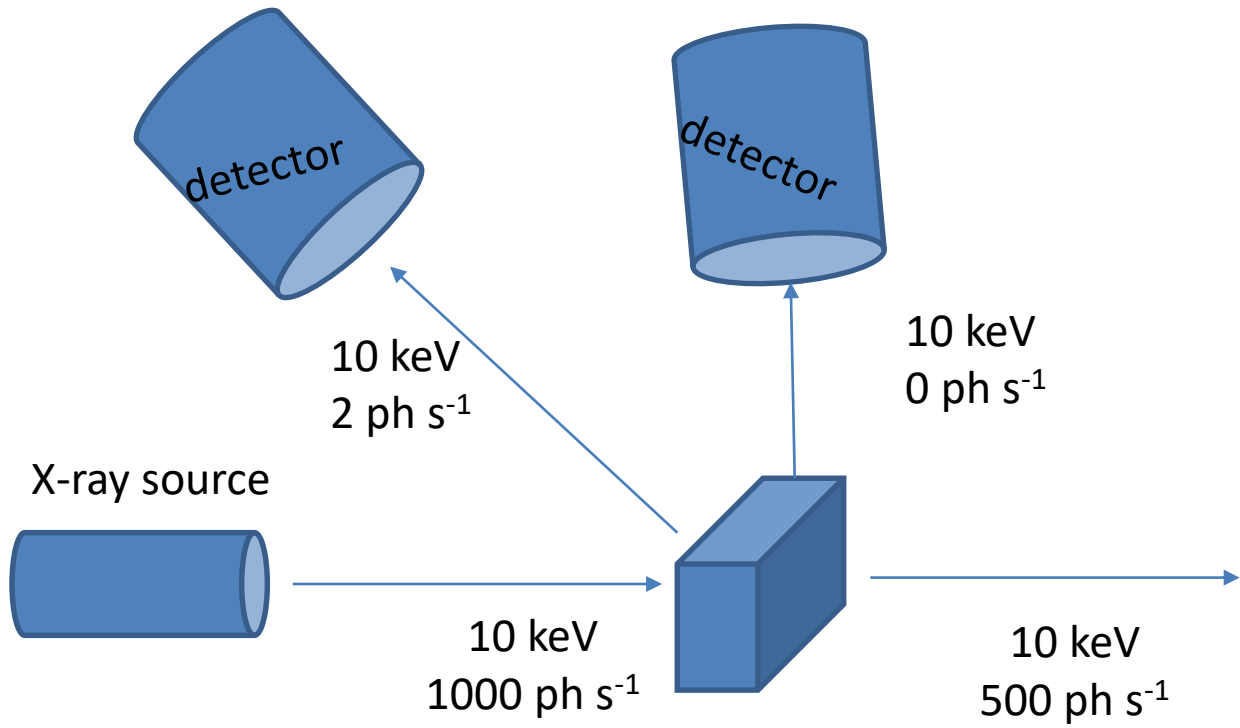
500
photons



Detector I

Transmittance = I/I_0

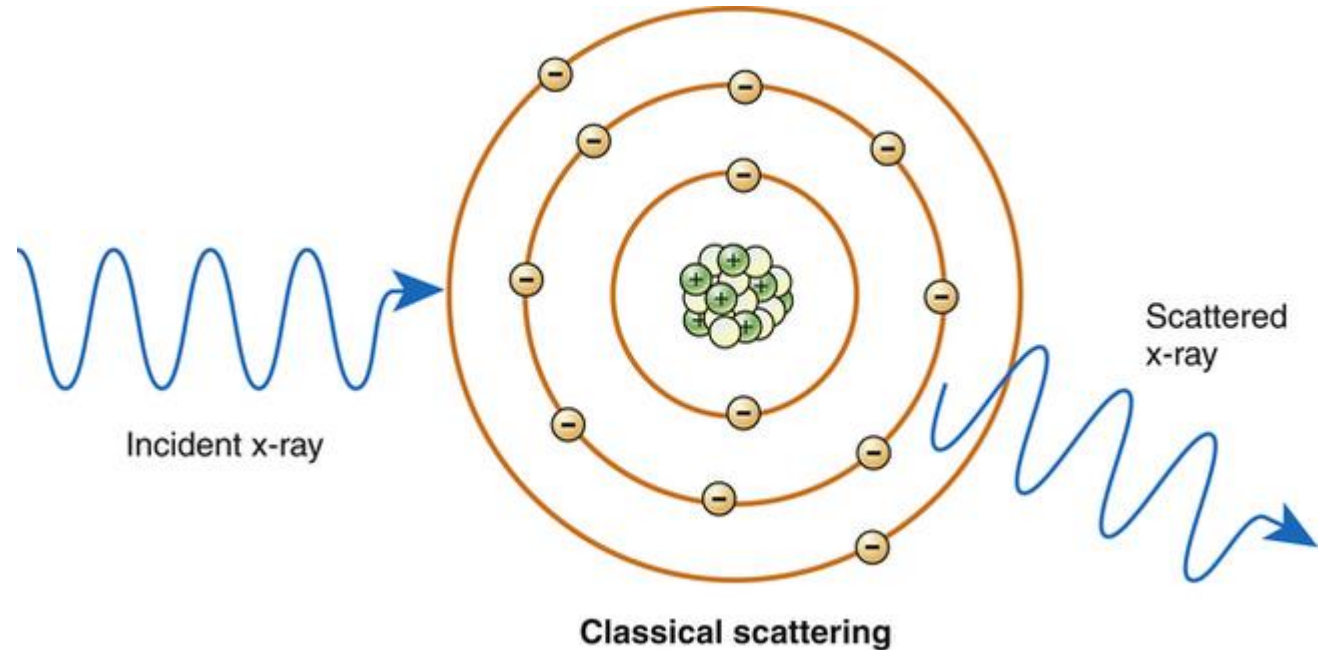
$T = 0.5$ or 50 %



5.4 X-ray Thomson scattering

Thomson Scattering

□ In Thomson scattering the energy of the photon does not change



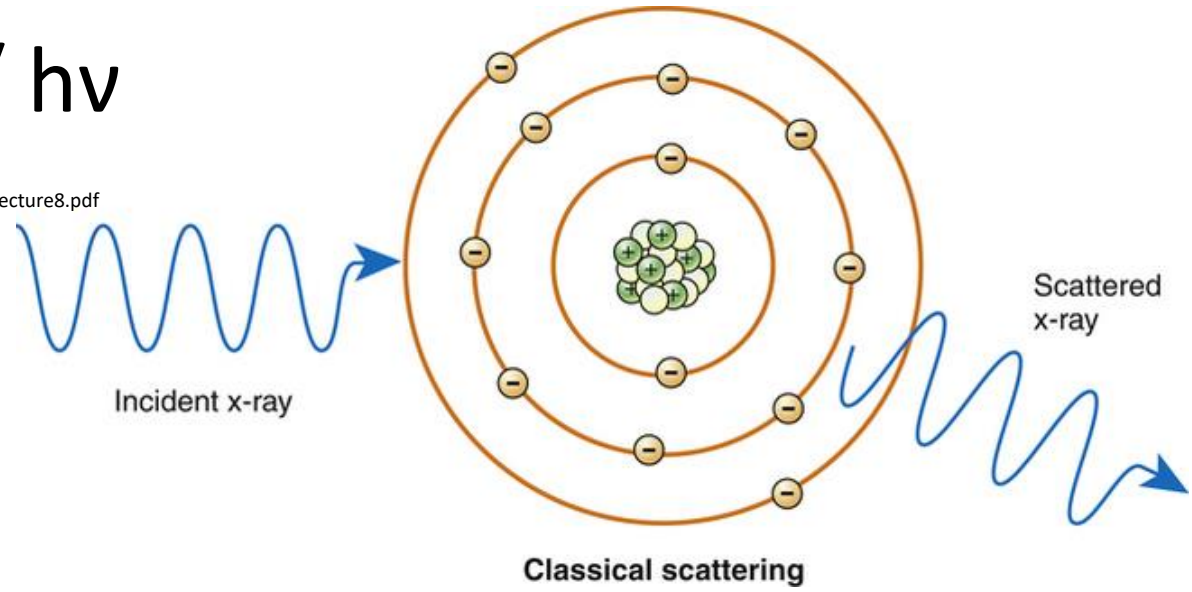
http://radiologykey.com/wp-content/uploads/2016/02/B9780323069748100074_f07-02-9780323069748.jpg

Dependence of E and Z

- For the energy range we are mostly interested, 1-100 keV, it is nearly constant up to 10 keV and then starts to decrease
- Since Thomson scattering is the result of the interaction between the photon and the electron it will increase is function of Z

$$\sigma_{\text{Thomson}} \propto Z / h\nu$$

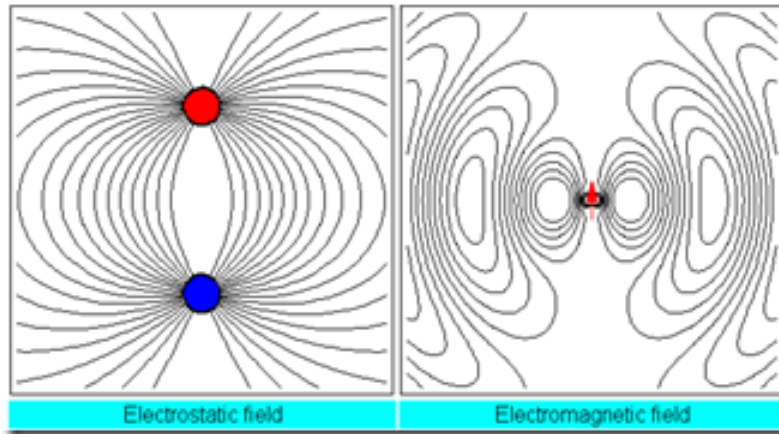
https://www.medphysics.wisc.edu/courses/mp501/501_lecture8.pdf



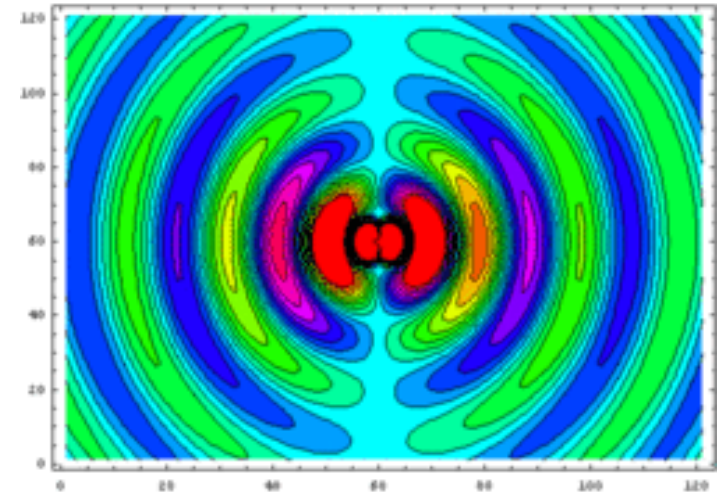
http://radiologykey.com/wp-content/uploads/2016/02/B9780323069748100074_f07-02-9780323069748.jpg

How does it take place?

□ Dipole radiation



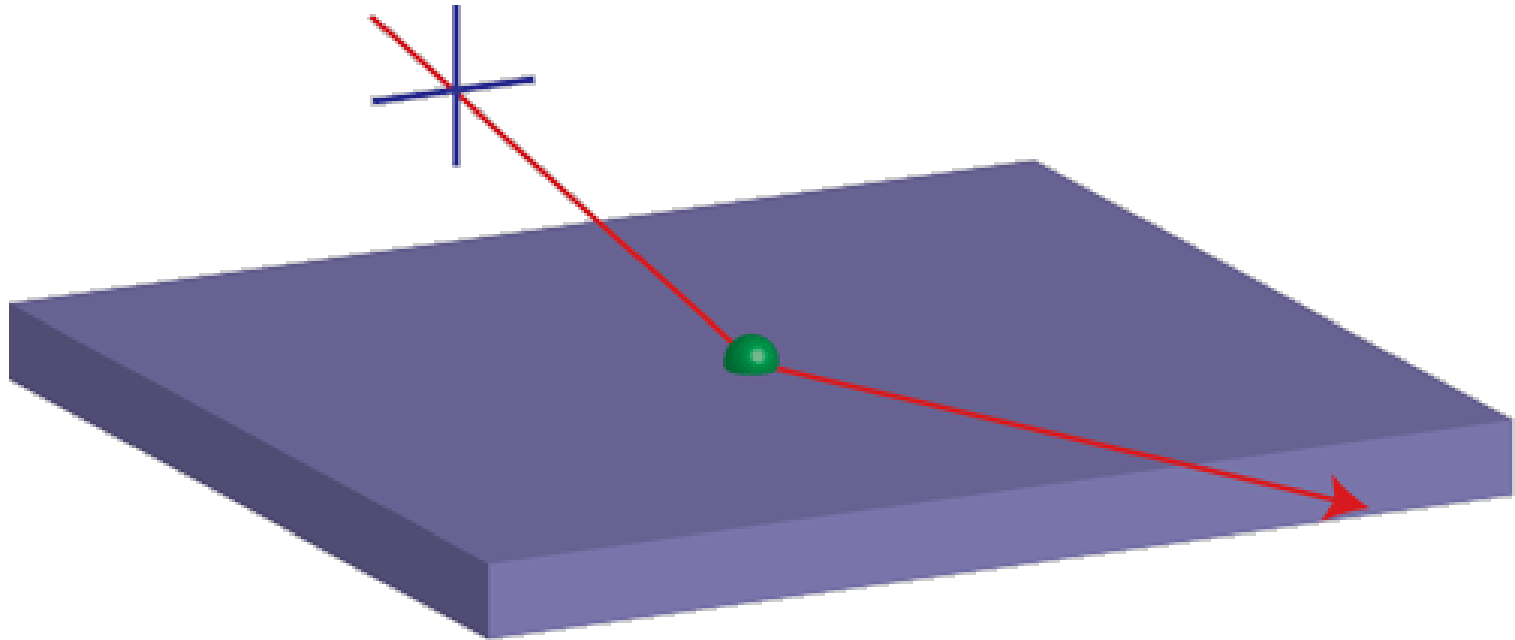
□ In which direction is the electron moving?



<http://allenluadvance.blogspot.com.br/2009/02/dipole-and-monopole-antenna-modeling.html>

Thomson Scattering by metal surfaces

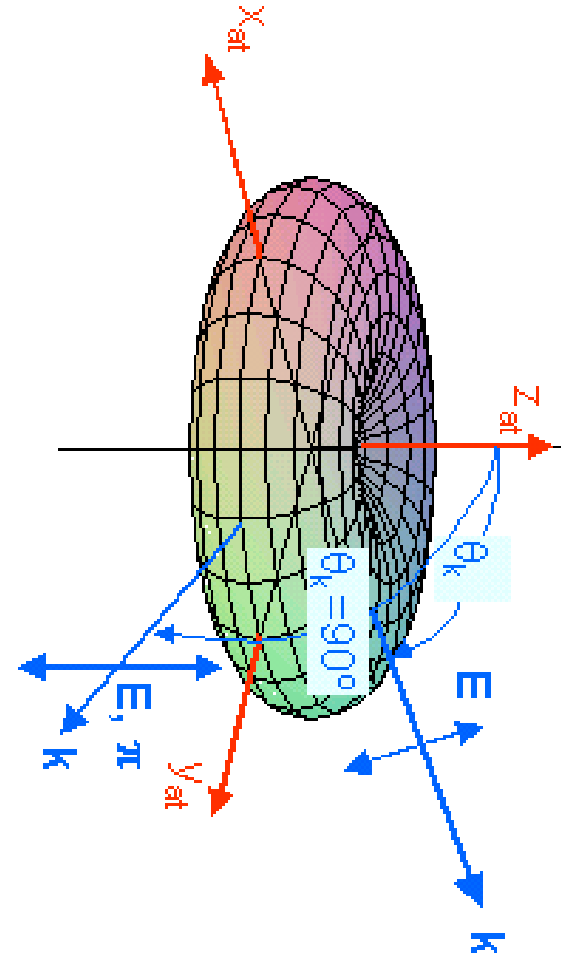
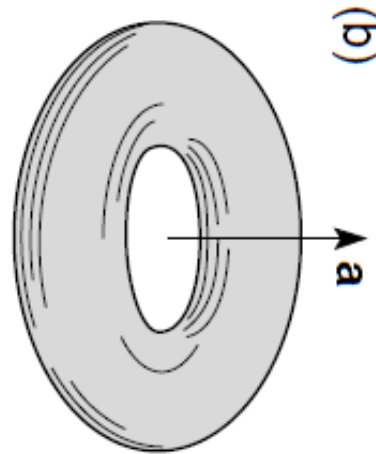
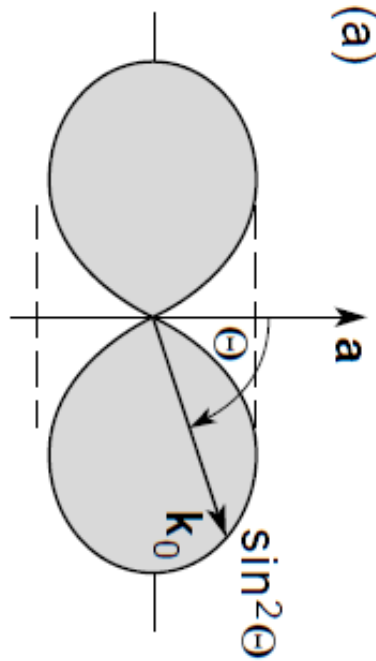
□ Polarization by reflection



<http://background.uchicago.edu/~whu/intermediate/Polarization/polscat.gif>

Thomson Scattering of polarized radiation

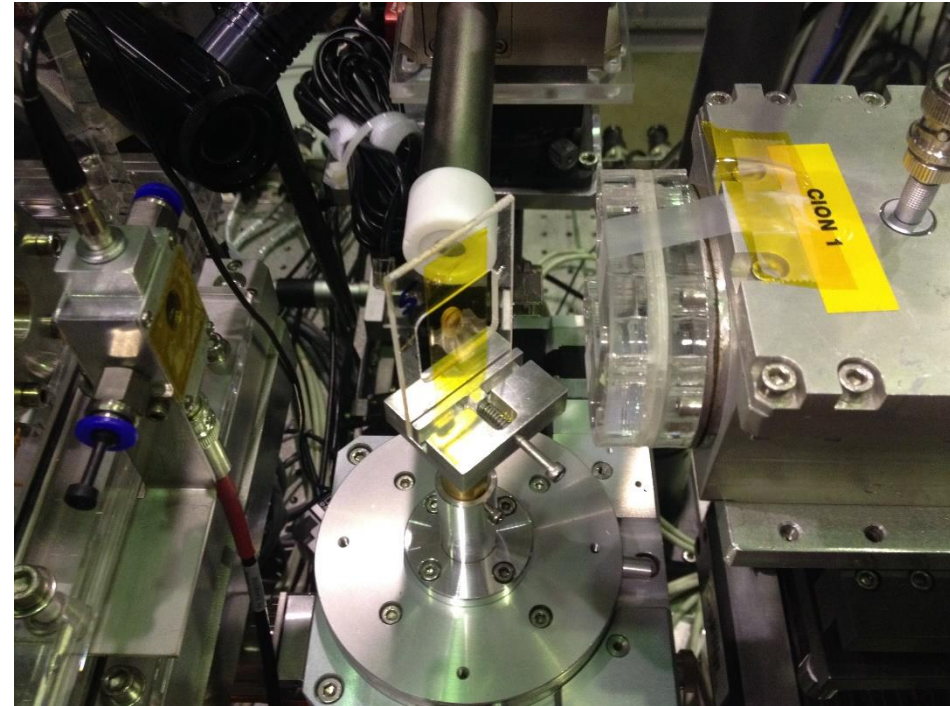
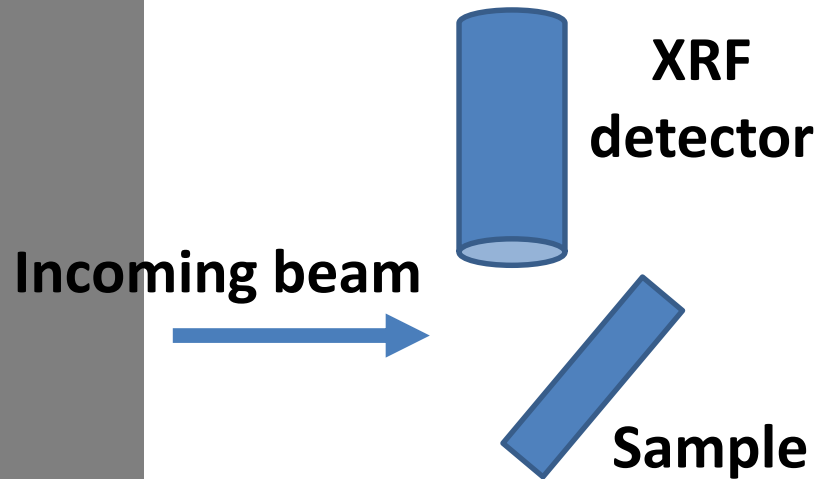
□ Dipole



<http://staff.mbi-berlin.de/hertel/physik3/chapter8/8.7html/01.htm>

Thomson Scattering of polarized radiation

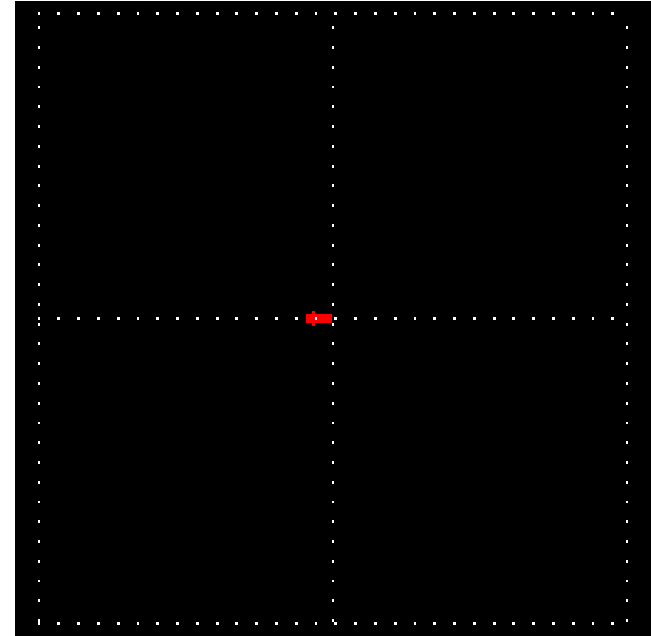
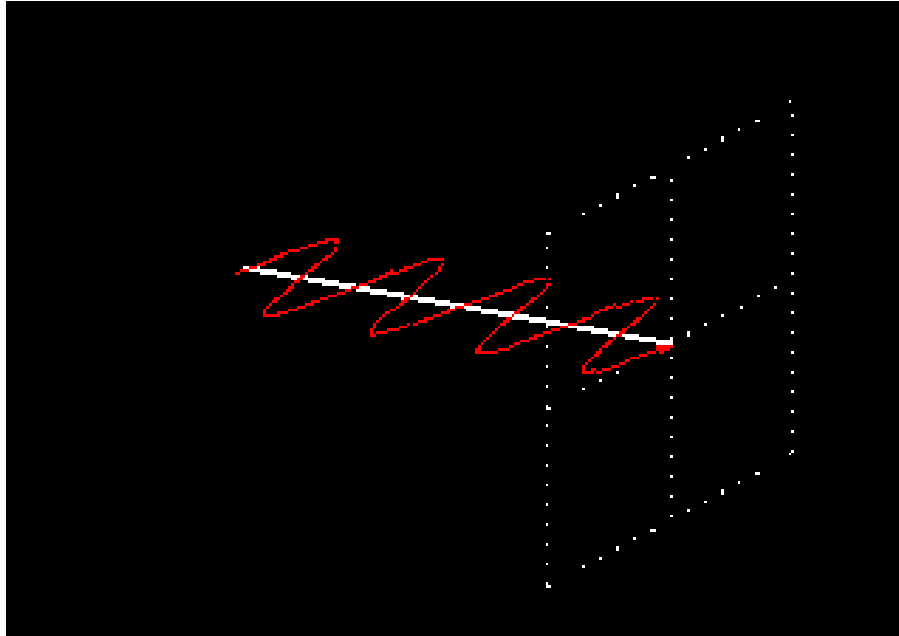
❑ XRF beamline@LNLS



❑ Why is the detector placed at this position?

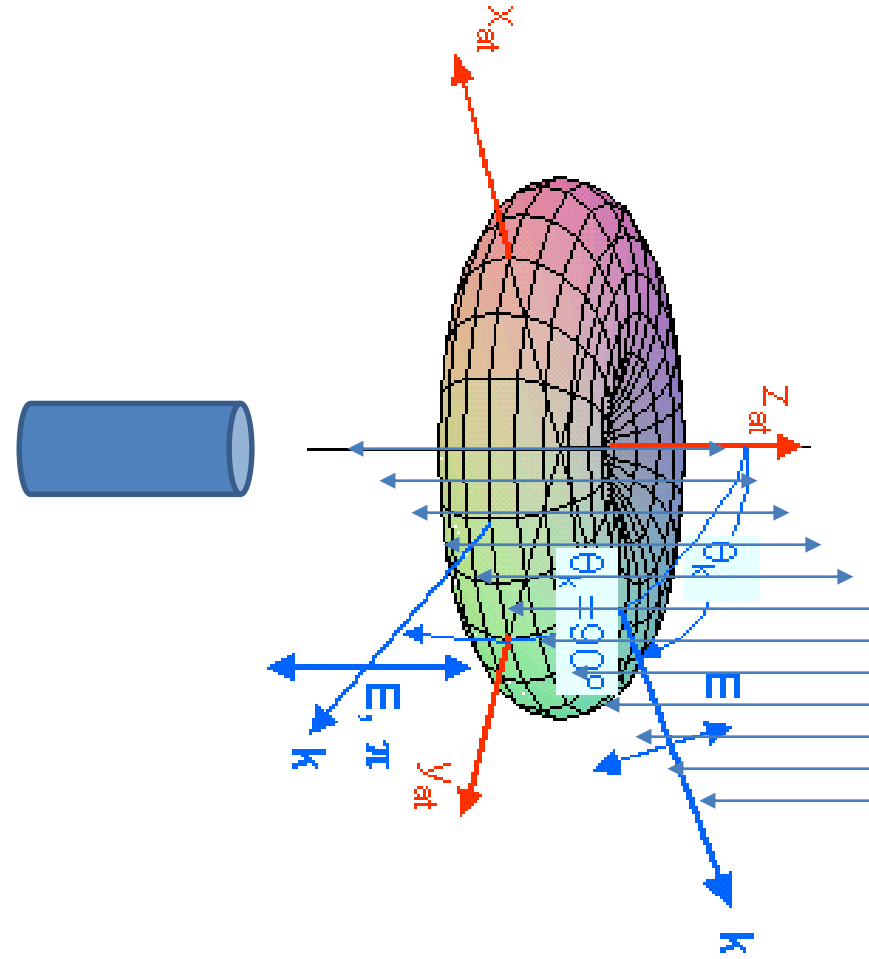
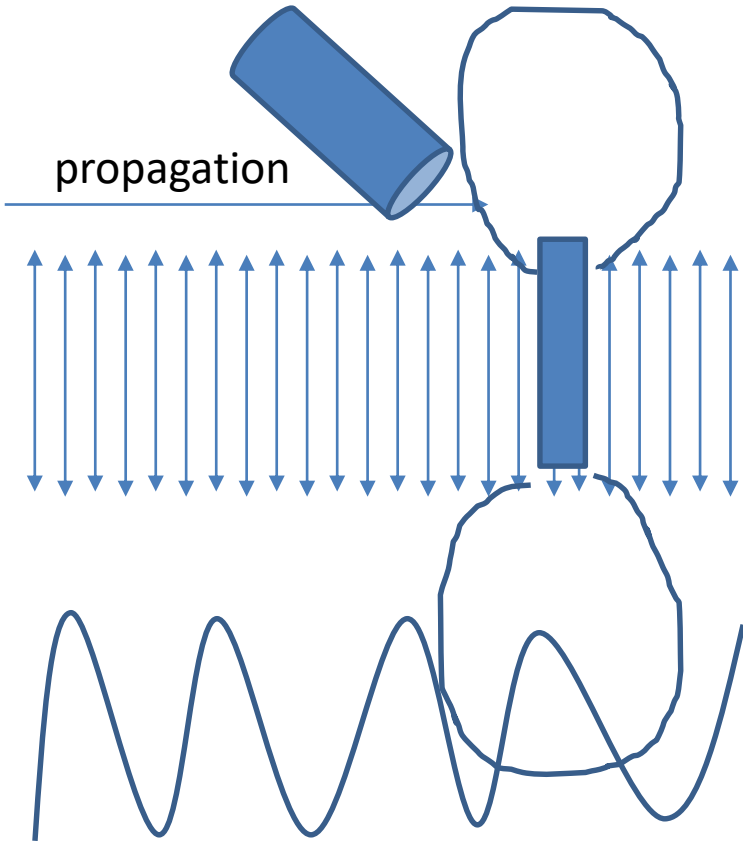
Polarization

□ Horizontal polarization



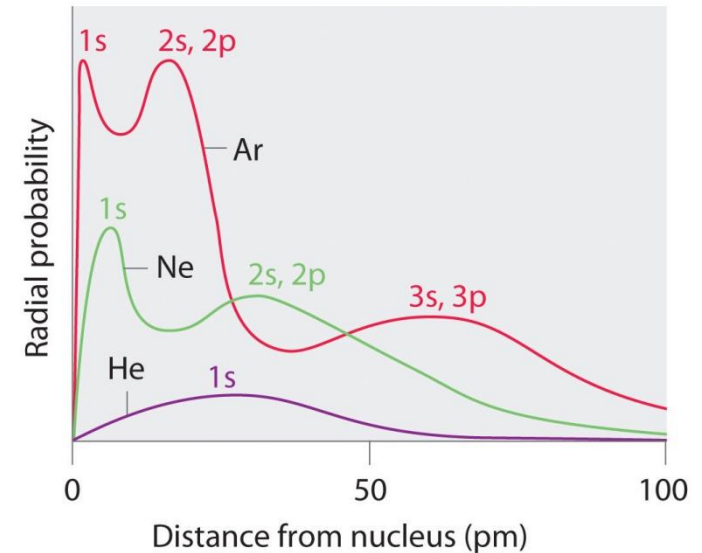
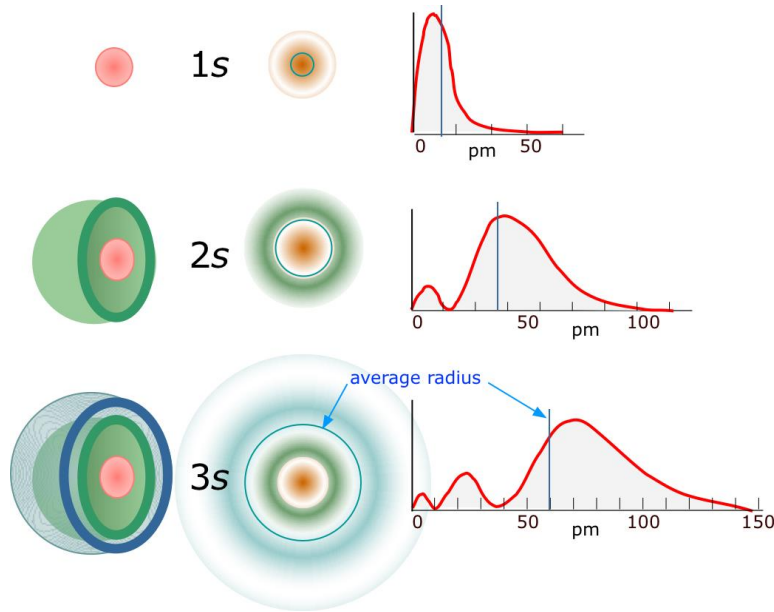
<http://cddemo.szilab.org/>

Top view



Electronic Structure

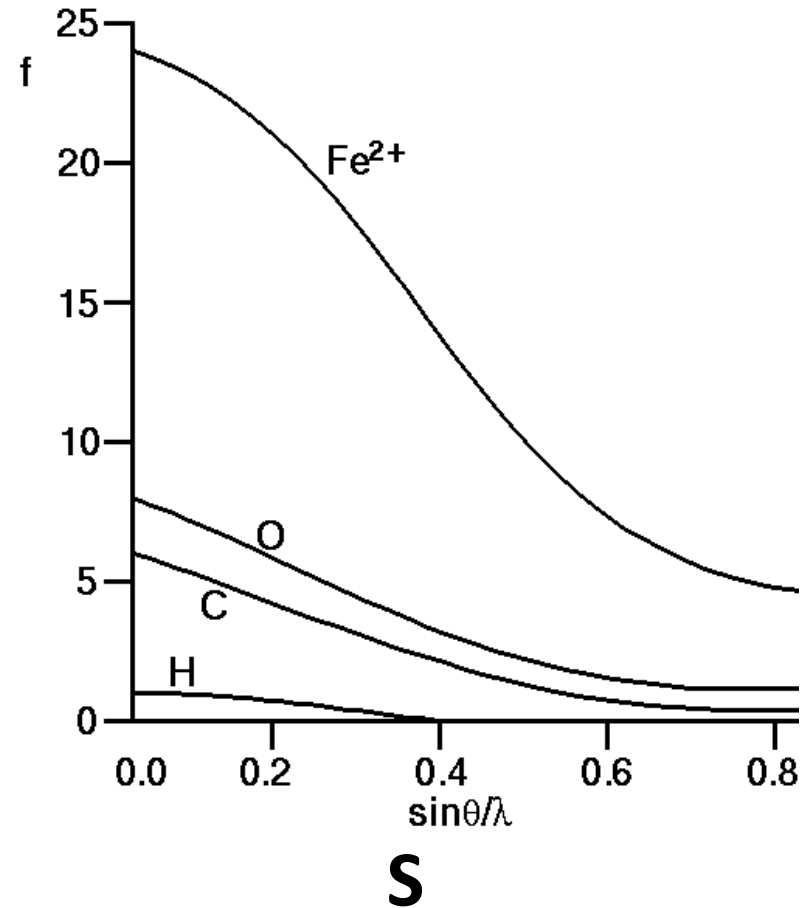
Radial electronic distribution



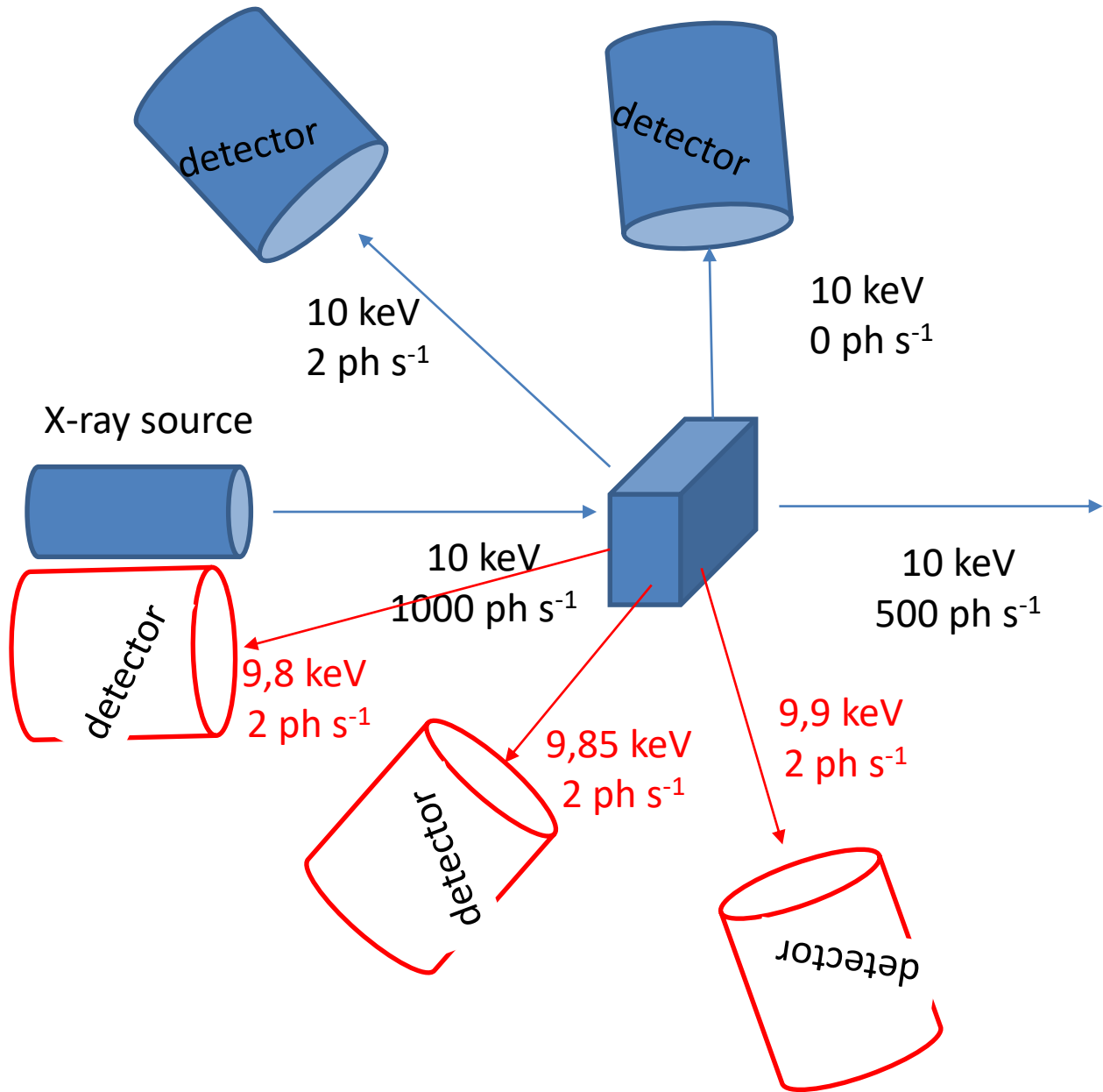
<http://chemwiki.ucdavis.edu/@api/deki/files/27872/e730a31a708c5fe85622cbfd31bba9ca.jpg?revision=1>

Atomic Scattering Factors

Angular dependence of F

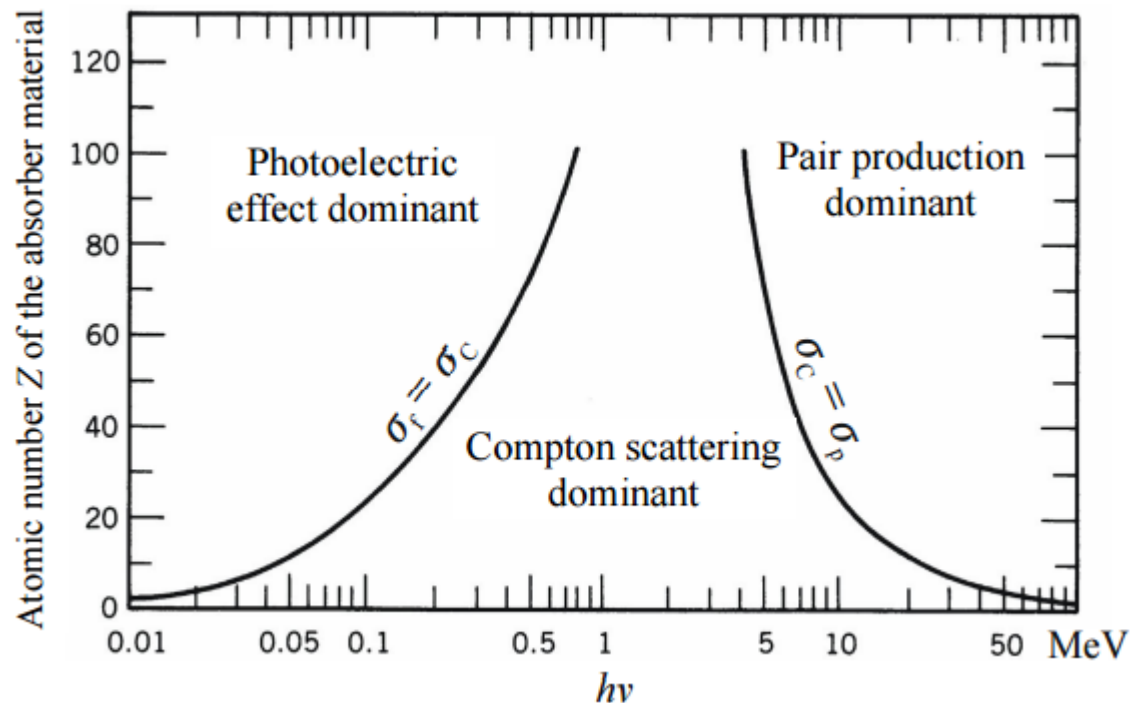


5.5 Compton Scattering



Compton Scattering is more pronounced at high energies

- Compton increases as the photoelectric absorption decreases



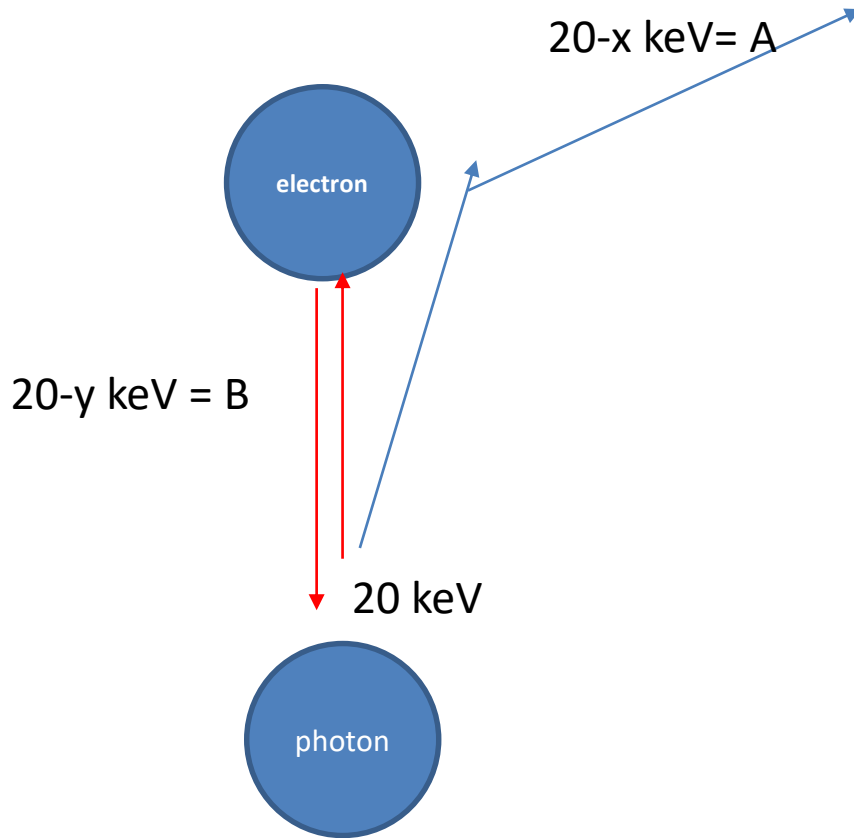
<http://www.nuclear-power.net/nuclear-power/reactor-physics/interaction-radiation-matter/interaction-gamma-radiation-matter/gamma-ray-attenuation/>

A > B

A=B

A < B

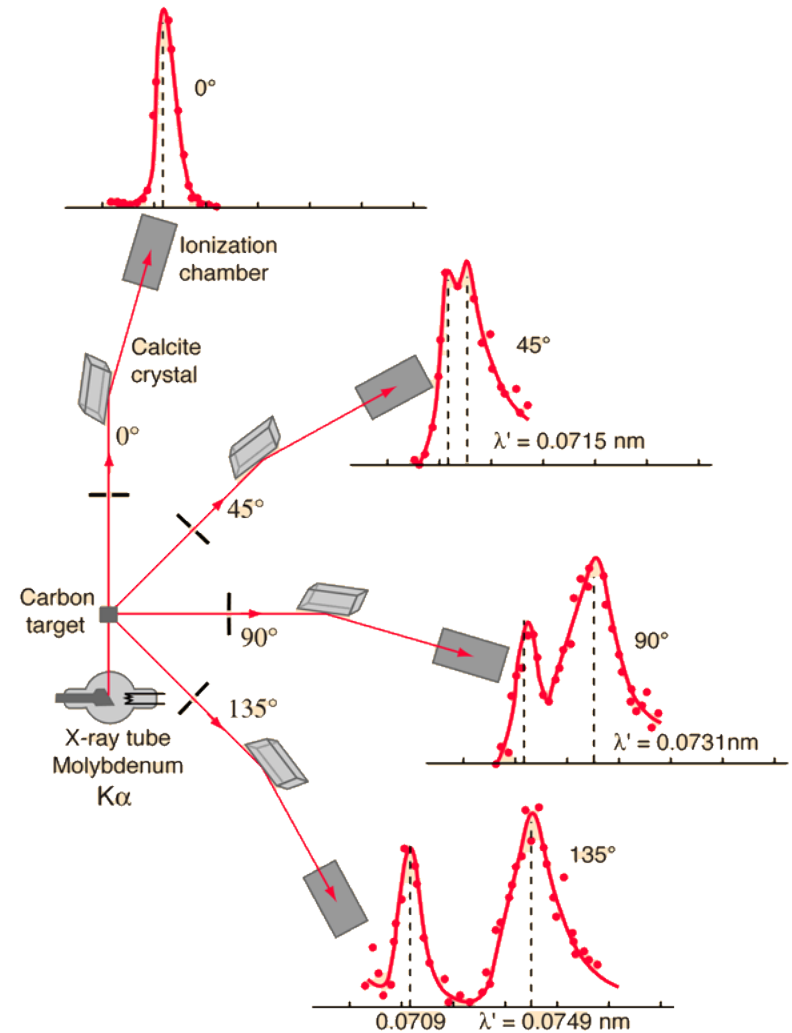
x < y



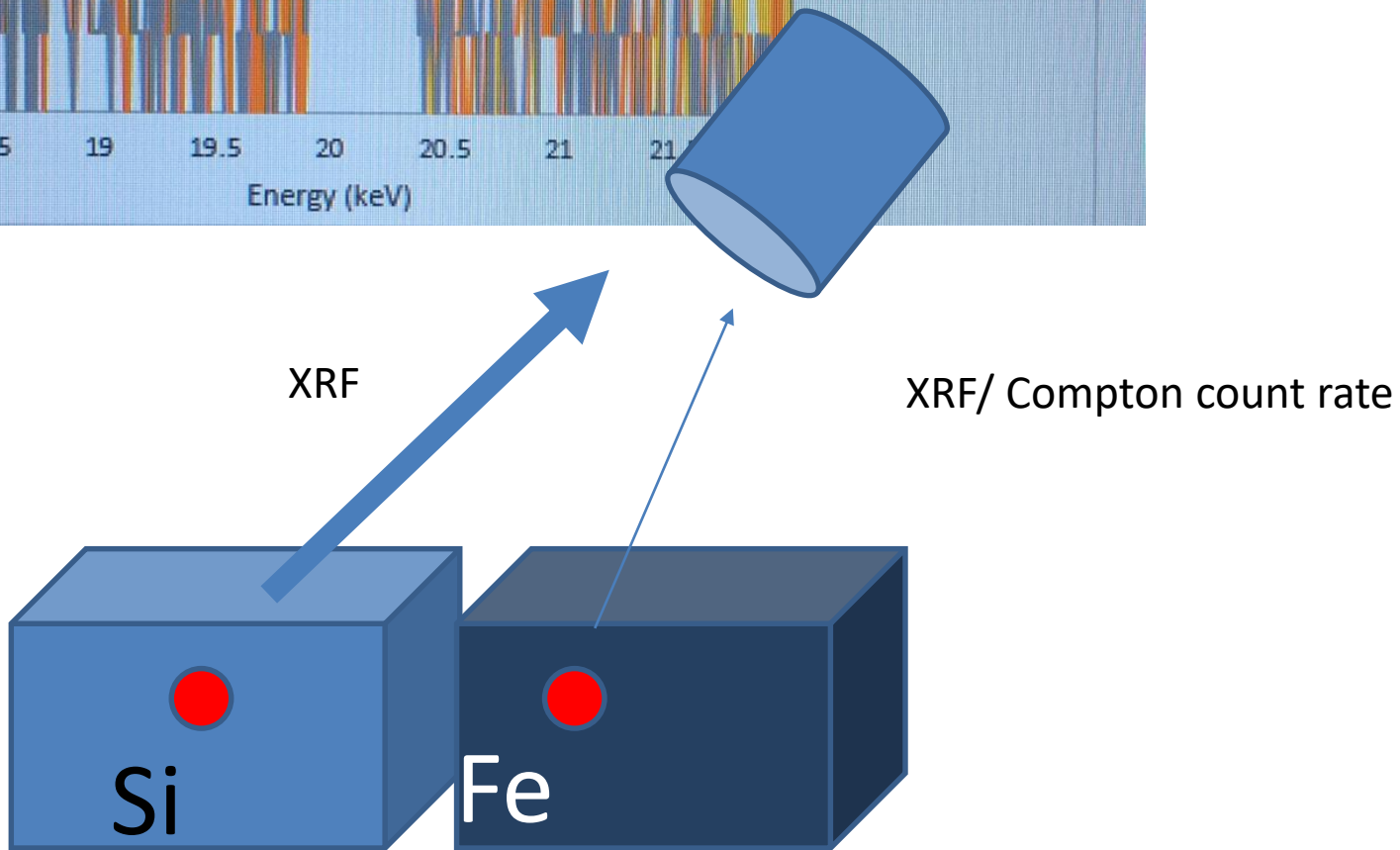
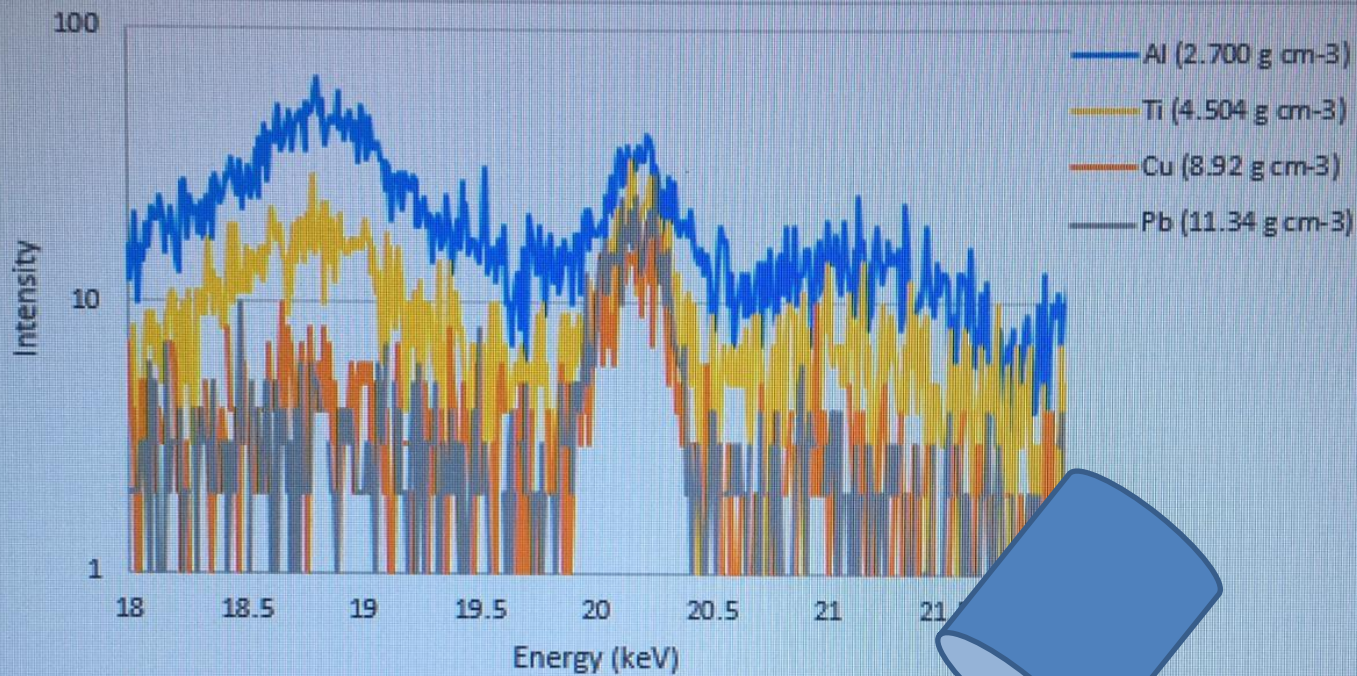
Compton Scattering

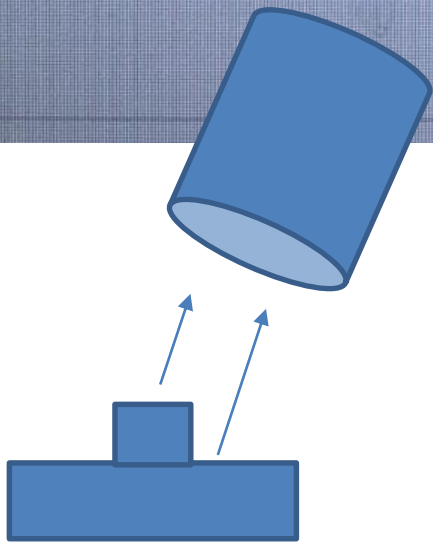
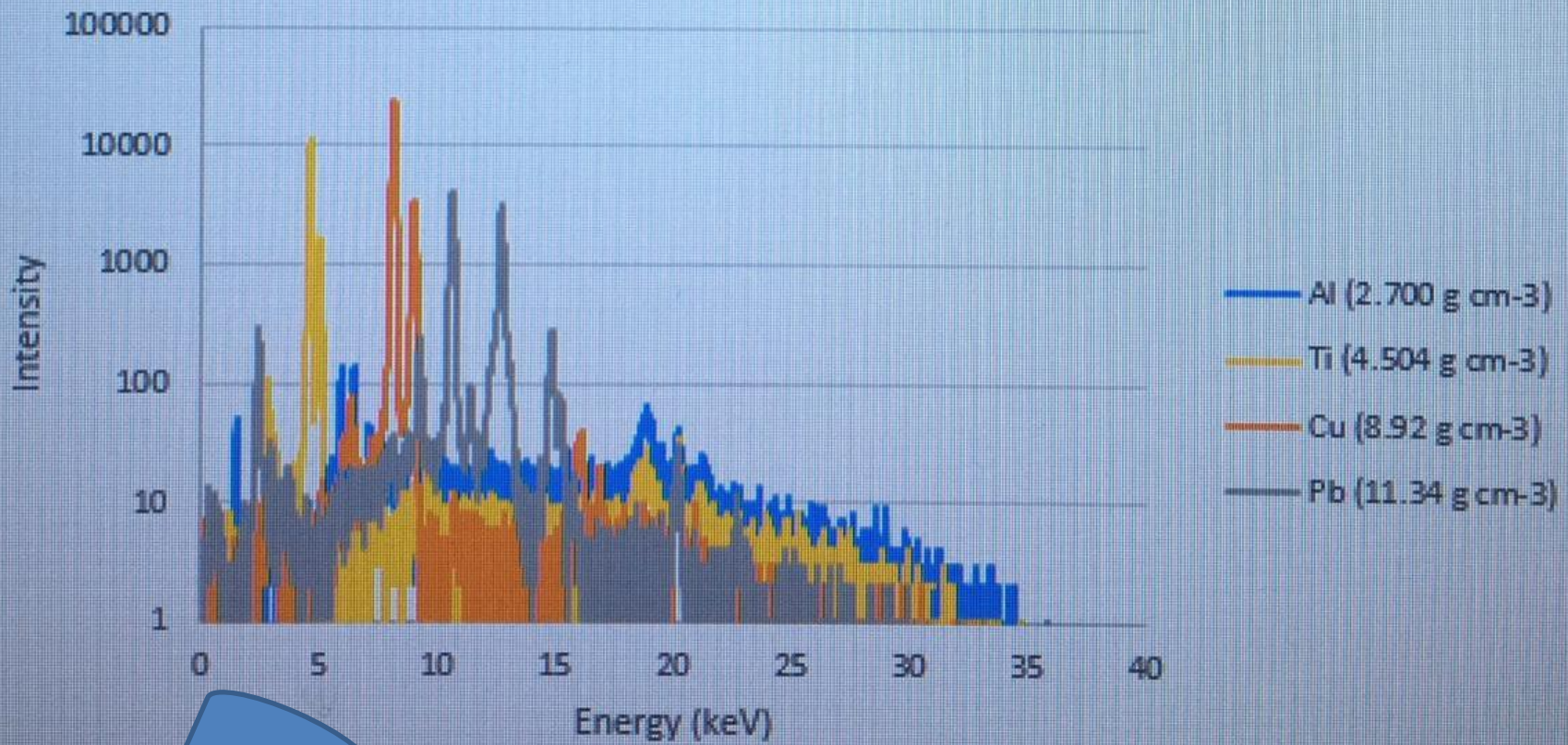
□ angle dependence

<http://hyperphysics.phy-astr.gsu.edu/hbase/quantum/compton.html#c1>

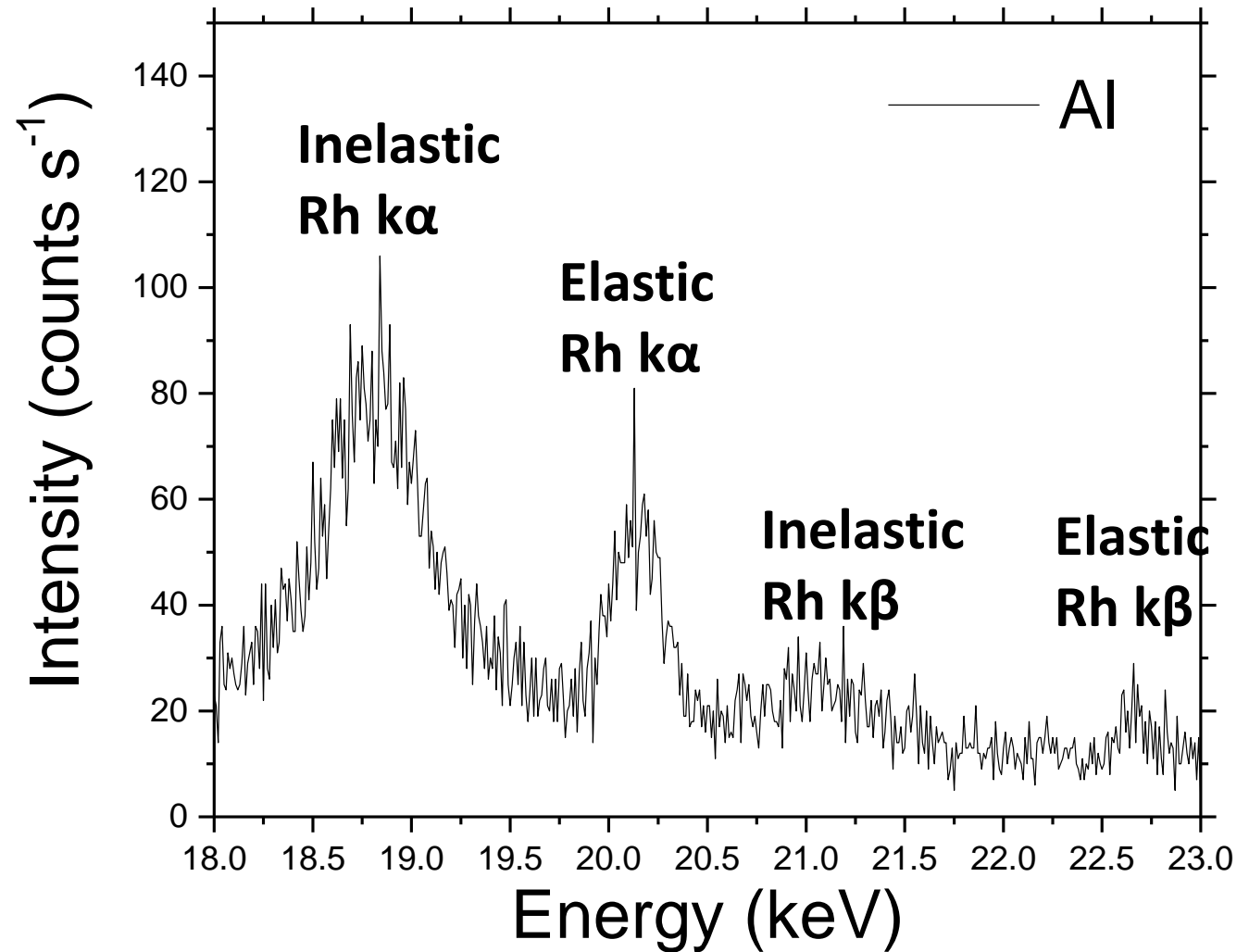


- $E_i = 20200 \text{ eV @ } 45^\circ$ $E_f = 19988 \text{ eV}$
- $E_i = 20220 \text{ eV @ } 90^\circ$ $E_f = 19450 \text{ eV}$
- $E_i = 20200 \text{ eV @ } 180^\circ$ $E_f = 18737 \text{ eV}$



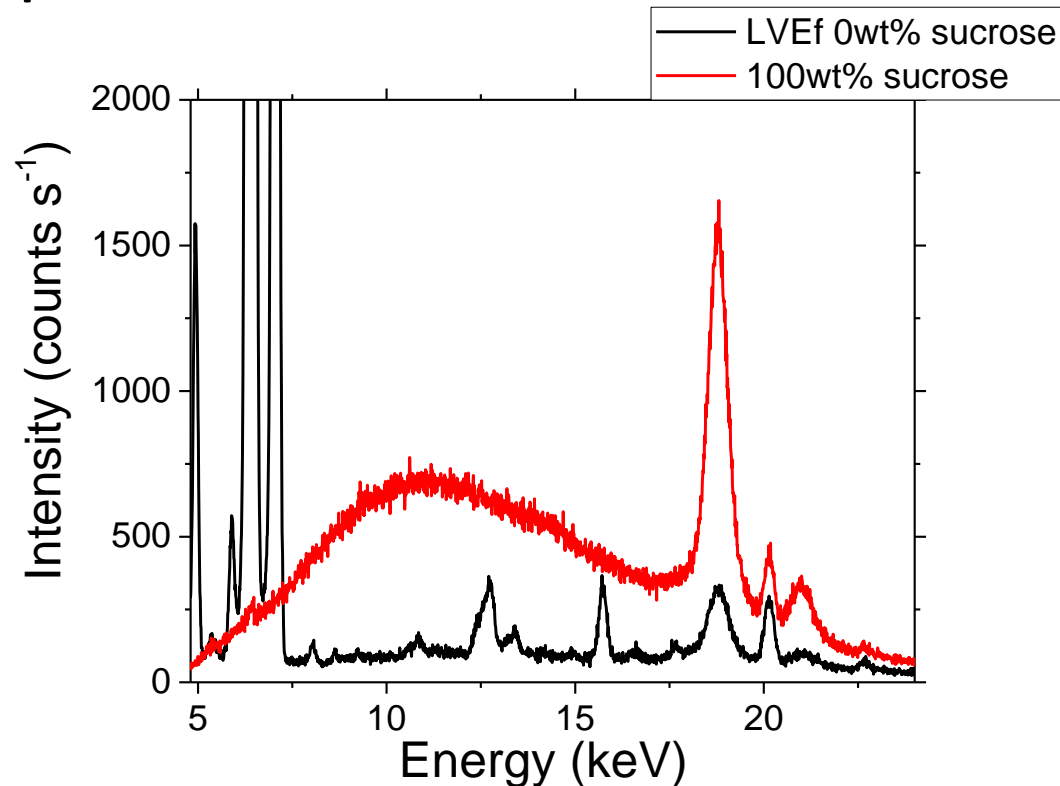


Rh K Compton Scattering



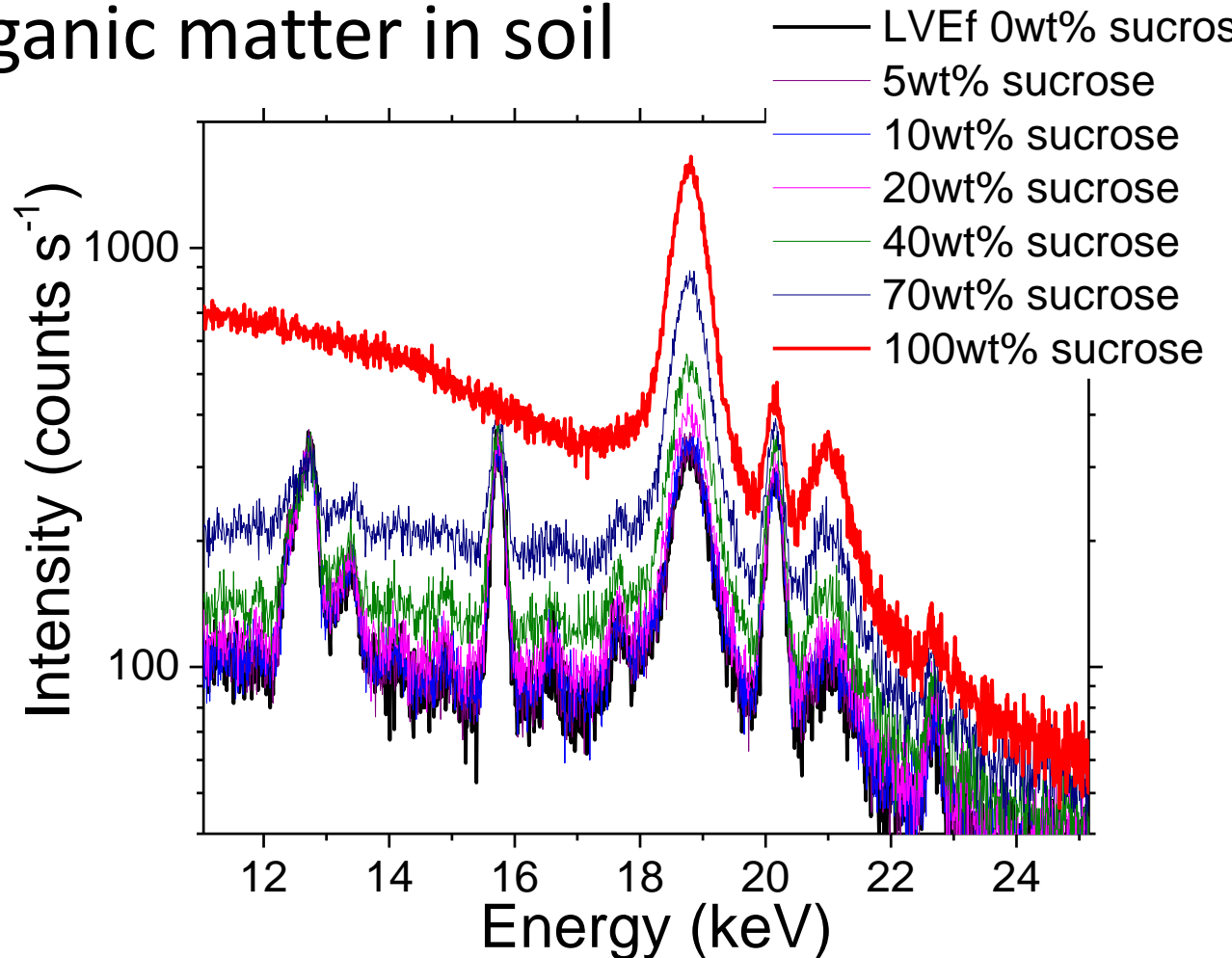
Compton scattering from light and heavy samples

- Black: soil sample
- Red: sucrose sample



Compton Scattering

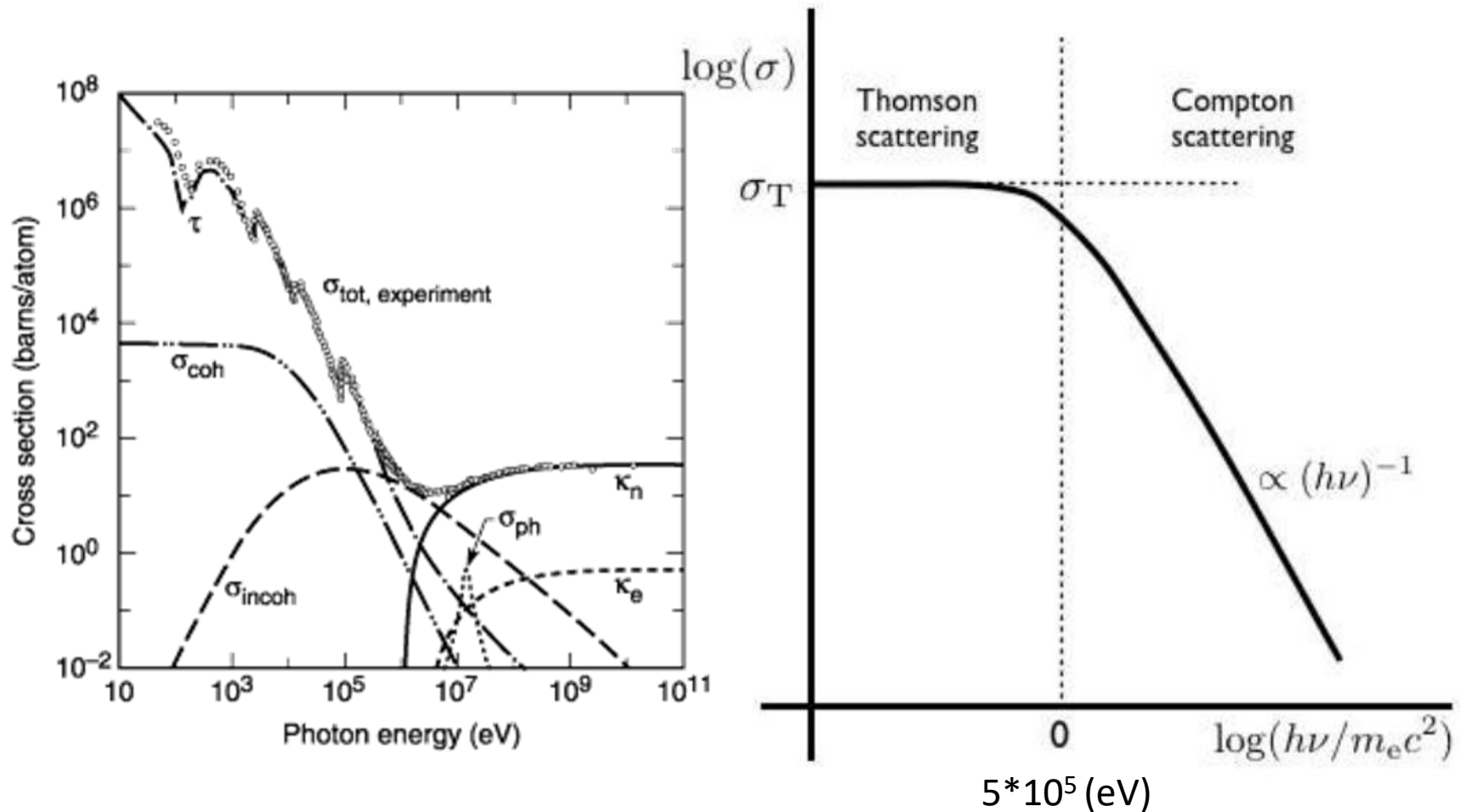
Compton response to increasing organic matter in soil



Exploring the competition between the process

The Compton vs Thomson Scattering

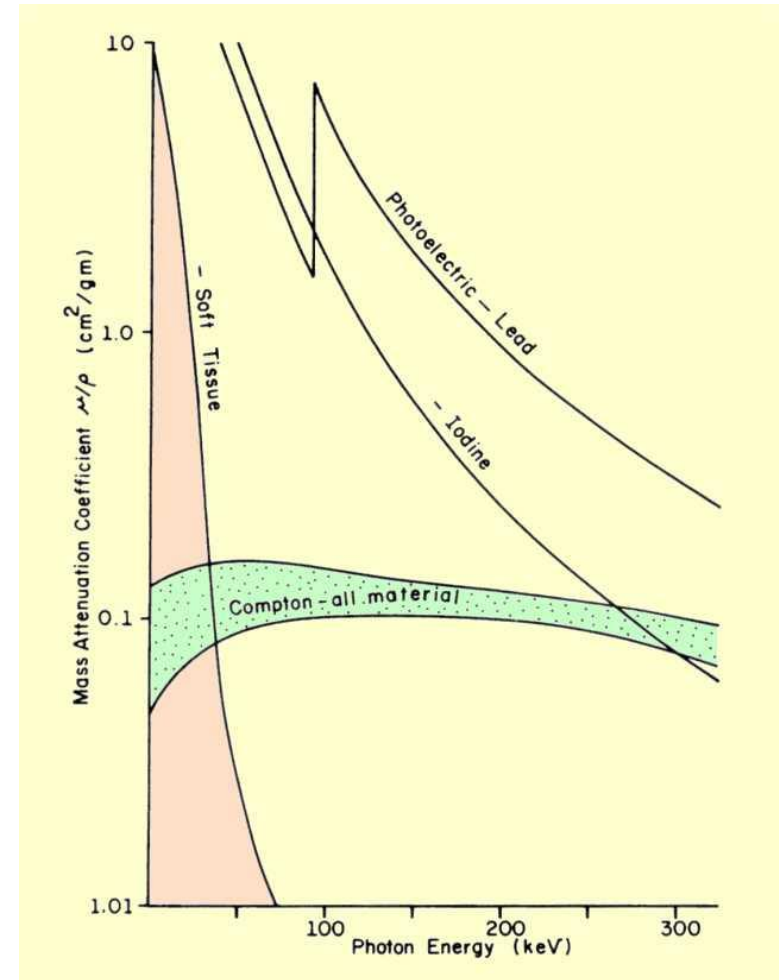
- Compton scattering is quantum mechanical process



http://www.astro.yale.edu/vdbosch/astro320_summary23.pdf

□ The reason why light elements scatters more Rh $K\alpha$ Compton is because it stops absorbing!

□ For each energy there is a probability that one of several phenomena takes place



<http://www.sprawls.org/ppmi2/INTERACT/#Compton Rates>