

15.a Introduction

When renormalizing a model, we must choose a renormalization prescription R . However, this is arbitrary since there are infinite other choices R' . Nevertheless, observables are independent of renormalization prescriptions. A change in renormalization prescription must be compensated by simultaneous changes in the renormalization parameters to have the physical results renormalization-invariant. The renormalization group (RG) studies these changes!

The renormalization group equations (RGE) are extremely useful in summing series of logarithms ~~to~~ to all orders in perturbation theory! Let's elaborate a bit. In QED, the vacuum polarization is given at one-loop by

$$\Pi^{\mu\nu}(p) \equiv \text{diagram} + \text{counterterm} \equiv i\Pi_{\mu\nu}(p) = i(p^2 g_{\mu\nu} - p_\mu p_\nu) \times \left[-\frac{\alpha}{\pi} \int_0^1 dx 2x(1-x) \left(\frac{2}{\epsilon} - \gamma_E + \ln \left(\frac{4\pi\mu^2}{m^2 q^2 x(1-x)} \right) \right) \right] \quad (1)$$

We learnt that for $-q^2 \gg m^2$ we have [(13.18)-(15.1)]

$$\Pi^{\mu\nu}(q^2) \rightarrow \frac{\alpha}{\pi} \left[\frac{1}{3} \ln \frac{|q^2|}{m^2} - \frac{5}{9} + \mathcal{O}\left(\frac{m^2}{q^2}\right) \right] \quad (2)$$

So, there is the appearance of a large logarithm that might spoil perturbation theory. The solution was to sum up the series

$$\text{full} = m + \text{diagram (1)} + \text{diagram (2)} + \dots \Rightarrow e_R^2 \rightarrow \frac{e_R^2}{1 - \Pi^{\mu\nu}(q^2)} \quad (3)$$

The RGE performs the sum of potentially large logarithms ^[10.2] using the one-loop result without worrying with summing a geometrical series or details!

Notice that the appearance of logarithms is a feature of loop contributions. For instance a divergent integral like

$$I(p) = \int_0^\infty \frac{k dk}{k+p} \Rightarrow \frac{dI}{dp^2} = \int_0^\infty \frac{2k dk}{(k+p)^2} = \frac{1}{p}$$

Now, integrating twice $\frac{dI}{dp^2}$ up to Λ to regulate it, we get

$$I(p) = p \ln p - p (\ln \Lambda + 1) + C, \Lambda \Rightarrow \text{there are log's in the result!}$$

In QED we have, at large $-p^2$,

$$\Pi(p^2) = -\frac{e^2}{12\pi^2} \left[\frac{2}{\epsilon} + \ln \frac{\mu^2}{-p^2} + \dots \right] + \delta Z_3^{(4)} \text{ in dimensional regularization.}$$

that we used a cut off as in Pauli-Villars we would get

$$\Pi(p^2) = -\frac{e^2}{12\pi^2} \left[\ln \frac{\Lambda^2}{-p^2} + \dots \right] - \delta Z_3 \quad (5)$$

notice that the $\ln(-p^2)$ in (4) and (5) comes accompanied by μ^2 or Λ^2 (unphysical parameters) due to dimensional analysis! When we focus on μ^2 we have the continuous RG while focusing on Λ^2 leads to the Wilsonian RG! We will start with the former and comment on the latter after.

Stueckelberg + Petermann 1953

Gell-man - Low 1959 $\rightarrow$ behavior of QED at short distances

block-spin RG Kadanov 1966

$\Downarrow$

Continuum $\Rightarrow$ dif. eq. Wilson 1971

F.T. : lattice or path integral: widely

note : $\left\{ \begin{array}{l} \text{Callan} \\ \text{Symanzik} \end{array} \right.$

widespread applications.

15.b Renormalization Group Equation.

15.3

Let's obtain the RGE for the $\lambda\phi^4$ model

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} \partial_\mu \phi_B \partial^\mu \phi_B - \frac{1}{2} m_B^2 \phi_B^2 - \frac{\lambda_B}{4!} \phi_B^4 \\ &= \frac{1}{2} \partial_\mu \phi_R \partial^\mu \phi_R - \frac{1}{2} m_R^2 \phi_R^2 - \frac{\lambda_R}{4!} \phi_R^4 \\ &\quad + \frac{1}{2} (Z_3^{-1}) \partial_\mu \phi_R \partial^\mu \phi_R - (Z_0^{-1}) \frac{1}{2} m_R^2 \phi_R^2 - (Z_1^{-1}) \frac{\lambda_R}{4!} \phi_R^4 \end{aligned} \quad (6)$$

where

$$\phi_B = Z_3^{1/2} \phi_R$$

$$\mu^{-\epsilon} \lambda_B = Z_3^{-2} Z_1 \lambda_R \quad (7)$$

$$m_B^2 = Z_3^{-1} Z_0 m_R^2$$

We assume that the renormalization scheme is such that Z_i 's are mass independent. For instance, we are using the MS scheme. Working in $d = 4 - \epsilon$ dimensions we know explain why!

$$\Gamma_R^{(n)}(p_i, \lambda_R, m_R, \mu, \epsilon) = Z_3^{n/2} \Gamma_B^{(n)}(p_i, \lambda_B, m_B, \epsilon) \quad (8)$$

where

$$Z_i = Z_i(\lambda_R, \epsilon) \quad (9)$$

Notice that $\Gamma_B^{(n)}$ does not depend on μ ! So, $\mu \frac{d}{d\mu}$ (8) leads to

$$\begin{aligned} \mu \left(\frac{\partial}{\partial \mu} + \frac{d\lambda_R}{d\mu} \frac{\partial}{\partial \lambda_R} + \frac{dm_R}{d\mu} \frac{\partial}{\partial m_R} \right) \Gamma_R^{(n)}(p_i, \lambda_R, m_R, \mu, \epsilon) &= \\ &= \frac{n}{2} Z_3^{n/2-1} \mu \frac{dZ_3}{d\mu} \Gamma_B^{(n)} = \frac{n}{2} \frac{\mu}{Z_3} \frac{d}{d\mu} (Z_3) \Gamma_R^{(n)} \end{aligned} \quad (9)$$

Now we define in order to ⁽⁷⁾ use the method of characteristics [15.4]

$$\beta(\lambda_R, \epsilon) \equiv \mu \frac{d\lambda_R}{d\mu} \stackrel{(7)}{=} \mu \frac{d}{d\mu} (\mu^{-\epsilon} z_3^2 \bar{z}_1^{-1} \lambda_0) \quad (10)$$

$$= -\epsilon \lambda_R + \lambda_R \underbrace{\mu \frac{d \ln(z_3^2 \bar{z}_1^{-1})}{d\mu}}_{\frac{1}{z_\lambda}} \xrightarrow{\epsilon \rightarrow 0} \beta(\lambda_R)$$

$$\beta(\lambda_R, \epsilon) \equiv \frac{1}{2} \mu \frac{d}{d\mu} \ln(z_3) \xrightarrow{\epsilon \rightarrow 0} \gamma(\lambda_R) \quad (11)$$

$$\gamma_w(\lambda_R, \epsilon) \equiv \frac{\mu}{\mu_R} \frac{d\mu_R}{d\mu} = \frac{1}{2} \mu \frac{d \ln(z_3 \bar{z}_0^{-1})}{d\mu} \xrightarrow{\epsilon \rightarrow 0} \gamma_w(\lambda_R) \quad (12)$$

so we write (9) as

$$\left[\mu \frac{\partial}{\partial \mu} + \beta(\lambda_R) \frac{d}{d\lambda_R} + \gamma_w(\lambda_R) \frac{\partial}{\partial \mu_R} - n \gamma(\lambda_R) \right] \Gamma_R^{(n)}(p_i, \lambda_R, \mu_R, \mu) = 0 \quad (13)$$

this is the RGE!

To calculate β we write the perturbation theory result for

$$Z_\lambda = z_3^{-2} z_1 = 1 + \sum_j \frac{a_{2j}(\lambda)}{\epsilon^j}$$

and

$$\partial_\lambda(\lambda) = \sum_j a_{2j} \lambda_R^j$$

(14)

[START BY THE DERIVATION OF eq (10), then do the series expansion!]

$$\lambda_0 = \mu^\epsilon \bar{z}_\lambda \lambda_R \quad (15)$$

Now

$$\mu \frac{d}{d\mu} Z_\lambda = \mu \frac{d\lambda_R}{d\mu} \frac{dZ_\lambda}{d\lambda_R} = \beta(\lambda_R, \epsilon) \frac{dZ_\lambda}{d\lambda_R}$$

that substituted in (10) yields

$$\beta(\lambda_R, \epsilon) Z_\lambda + \epsilon \lambda_R Z_\lambda \cancel{\beta(\lambda_R, \epsilon)} \frac{dZ_\lambda}{d\lambda_R} = 0 \quad (16)$$

At this point, $\beta(\lambda, \epsilon) = \sum_{k=1}^M \beta_k \epsilon^k \quad (17)$

Notice that β is regular at $\epsilon=0$ since the theory is renormalizable!

Now, (14)+(16)+(17) $\Rightarrow$

$$\sum_{k=1}^M \beta_k \epsilon^k \cdot \left(1 + \sum_j \frac{\partial \sigma}{\partial \epsilon^j}\right) + \epsilon \lambda_R \left(1 + \sum_j \frac{\partial \sigma}{\partial \epsilon^j}\right) + \left(\sum_{k=1}^M \beta_k \epsilon^k\right) \sum_j \frac{\lambda_R}{\epsilon^j} \frac{d\sigma}{d\lambda_R} \quad (18)$$

From (18) it follows:

• for $2 \leq k \leq M$ $\beta_k = 0$ from terms exhibiting ϵ^k

• the ϵ^1 component $\Rightarrow \beta_1 + \lambda_R = 0 \Rightarrow \beta_1 = -\lambda_R$

• $\epsilon^0 \Rightarrow \beta_0 + \cancel{\beta_1 a_1} + \cancel{\lambda_R a_1} + \beta_1 \lambda_R \frac{da_1}{d\lambda_R} = 0 \Rightarrow \beta_0 = \beta(\lambda) = \lambda_R^2 \frac{da_1}{d\lambda_R}$

• Notice that $\beta(\lambda_R^{\epsilon}) = -\epsilon \lambda_R + \beta(\lambda_R) \quad (20)$

• From ϵ^{-2} :

$$\beta_0 a_2 + \cancel{\beta_1 a_{2+1}} + \cancel{\lambda_R a_{2+1}} + \lambda_R \beta_0 \frac{da_2}{d\lambda_R} + \cancel{\lambda_R a_{2+1}} \beta_1 \frac{da_{2+1}}{d\lambda_R} = 0$$

$$\Rightarrow \lambda_R^2 \frac{da_{n+1}}{d\lambda_R} = +\beta(\lambda_R) \frac{d}{d\lambda_R} (\lambda_R a_0) \quad (21)$$

Comments:

(i) $\beta(\lambda_R)$ is determined by the residues of the simple poles ~~in~~ in ϵ of Z_λ !

(ii) Bonus: higher order poles of Z_λ can be obtained from $a_\perp$ using (21)! At each order in perturbation theory the new divergence arises in the $\frac{1}{\epsilon}$ form! Higher order terms $\frac{1}{\epsilon^n}$ are determined by lower order perturbation expansion!

Expressions similar to (19) can be obtained for $\gamma(\lambda_R)$ and $\gamma_\omega(\lambda_R)$:

$$Z_3 = 1 + \sum \frac{Z_3^{(n)}}{a^0} \Rightarrow \gamma(\lambda_R) = -\frac{\lambda_R}{2} \frac{d}{d\lambda_R} Z_3^{(1)}(\lambda_R) \quad (22)$$

$$u_B^2 = Z_u u_k^2 \quad \text{with} \quad Z_u = 1 + \sum \frac{Z_u^{(n)}}{\epsilon^0} \Rightarrow \gamma_\omega(\lambda_R) = \frac{\lambda_R}{2} \frac{d Z_u^{(1)}}{d\lambda_R} \quad (23)$$

15.c Solution of the RGE

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Before solving the RGE (13), let's ~~analyze~~ write $\mu = \mu_0 e^t$

$$\Rightarrow \mu \frac{\partial}{\partial \mu} \rightarrow -\frac{\partial}{\partial t} \quad \text{so (13) reads}$$

$$\left[-\frac{\partial}{\partial t} + \beta(\lambda_k) \frac{\partial}{\partial \lambda_k} + \gamma_m(\lambda_k) u_k \frac{\partial}{\partial u_k} - n \gamma_p(\lambda_k) \right] T_R^{(u)}(p_i, \lambda_k, u_k, \mu_0 e^t) = 0 \quad (24)$$

As an auxiliary step let's solve

$$\frac{\partial}{\partial t} \bar{\lambda}(t, \lambda_k) = \beta(\bar{\lambda}) \quad (25)$$

$$\text{with } \bar{\lambda}(0, \lambda_k) = \lambda_k$$

$$\text{and } \frac{\partial \bar{u}}{\partial t}(t, \lambda_k, u_k) = \gamma_m(\bar{\lambda}) \bar{u} \quad (26)$$

$$\text{with } \bar{u}(0, \lambda_k, u_k) = u_k$$

$$\text{From (25)} \Rightarrow t = \int_{\lambda_k}^{\bar{\lambda}(t, \lambda_k)} \frac{dx}{\beta(x)} \quad (27)$$

$$\text{Also, integrating (26)} \Rightarrow \bar{u}(t, \lambda_k, u_k) = u_k \exp \left[\int_0^t dt' \gamma_m(\bar{\lambda}(t')) \right] \quad (28)$$

$$\stackrel{\text{"(25)"}}{\Rightarrow} u_k \exp \left[\int_{\lambda_k}^{\bar{\lambda}(t, \lambda_k)} \frac{\gamma_m(x)}{\beta(x)} dx \right] \quad (29)$$

Notice that

$$\frac{\partial}{\partial \lambda_k} (27) \Rightarrow 0 = -\frac{1}{\beta(\lambda_k)} + \frac{1}{\beta(\bar{\lambda}(\lambda_k, t))} \frac{\partial \bar{\lambda}}{\partial \lambda_k} \stackrel{(25)}{\Rightarrow} \left(-\frac{\partial}{\partial t} + \beta(\lambda_k) \frac{\partial}{\partial \lambda_k} \right) \bar{\lambda}(\lambda_k, t) = 1 \quad (30)$$

(15.7.)

$$\frac{\partial \bar{m}}{\partial \lambda_R} = \bar{m} \left[-\frac{\gamma_w(\lambda_R)}{\beta(\lambda_R)} + \frac{\gamma_w(\bar{\lambda})}{\beta(\bar{\lambda})} \underbrace{\frac{\partial \bar{\lambda}}{\partial \lambda_R}}_{\substack{\downarrow (25) + (30) \\ \frac{\beta(\bar{\lambda})}{\beta(\lambda_R)}}} \right]$$

$$= \bar{m} \left[-\frac{\gamma'_w(\lambda_R)}{\beta(\lambda_R)} + \frac{\gamma_w(\bar{\lambda})}{\beta(\lambda_R)} \right] \quad (30a)$$

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$$T_R^{(n)}(p_i, \gamma_k, w_k, \mu_0 e^{-t}) = T_k^{(n)}(p_i, \bar{\lambda}(t), \bar{w}(t), \mu_0) \quad (31)$$

in fact, substituting $\frac{1}{2}$ into (29) with $\delta = \frac{1}{2}$ we have:

$$\begin{aligned}
 & \left(\frac{\partial}{\partial t} + \beta(\lambda_k) \frac{\partial}{\partial \beta_k} + \gamma_m(\lambda_k) u_k \frac{\partial}{\partial u_k} \right) \Gamma_e^{(n)}(p_i, \bar{\lambda}, \bar{u}, \mu_0) = \\
 & \cdot \left[\frac{\partial \Gamma_R^{(n)}}{\partial \bar{\lambda}} \left(-\frac{\partial \bar{\lambda}}{\partial t} + \beta(\lambda_k) \frac{\partial \bar{\lambda}}{\partial \lambda_k} \right) + \frac{\partial \Gamma_k^{(n)}}{\partial \bar{u}} \left(-\frac{\partial \bar{u}}{\partial t} + \underbrace{\gamma_m(\lambda_k) u_k}_{(29) \bar{u}} \frac{\partial \bar{u}}{\partial u_k} \right) + \beta(\lambda_k) \frac{\partial \bar{u}}{\partial \lambda_k} \right] \\
 & \quad \downarrow (30) \qquad \qquad \qquad (26) \qquad \qquad \qquad \downarrow (30a) \\
 & = \frac{\partial \Gamma_k^{(n)}}{\partial \bar{\lambda}} \left(-\beta(\lambda_k) \frac{d\bar{\lambda}}{dt} + \beta(\lambda_k) \frac{d\bar{\lambda}}{dt} \right) + \frac{\partial \Gamma_k^{(n)}}{\partial \bar{u}} \bar{u} \left(-\gamma_m(\bar{\lambda}) + \gamma_m(\lambda_k) - \gamma_m(\lambda_k) + \gamma_m(\bar{\lambda}) \right) \\
 & \equiv 0!!! \Rightarrow \text{Eq'n (31) is correct!}
 \end{aligned}$$

The sol'n to the full (24) is

$$\Gamma_R^{(n)}(p_i, \lambda_R, \omega_R, \underbrace{\mu_0 e^{-t}}_{\mu}) = \Gamma_R^{(n)}(p_i, \bar{\lambda}_R(t), \bar{\omega}(t), \underbrace{\mu_0}_{\mu e^t}) \exp \left[-n \int_0^t dt' \gamma(\bar{\lambda}(t')) \right] \quad (32)$$

which follows direct from the meaning of $\pi_k^{(n)}(x_i, \bar{\lambda}, \bar{u}, \mu_0)$ and

$$\left(\frac{\partial}{\partial t} + \beta(\lambda_k) \frac{\partial}{\partial \lambda_k} - n \gamma(\lambda_k) \right) e^{-n \int_0^t \gamma(\bar{\lambda}(\epsilon')) d\epsilon'} = e^{-n \int_0^t \gamma(\bar{\lambda}(\epsilon')) d\epsilon'} \left[+n \gamma(\bar{\lambda}_k) - n \beta(\lambda_k) \left(\frac{-\gamma(\lambda_k) + \gamma(\lambda)}{\beta(\lambda_k)} \right) - n \gamma(\lambda_k) \right] = 0 \quad (33)$$

15.d Equation for physical observables

[15.9]

Physical quantities can be evaluated as a function $P(\lambda_k, m_k, \mu)$. However, they do not depend on the μ parameter. Therefore,

$$\mu \frac{dP}{d\mu} = \left[\mu \frac{\partial}{\partial \mu} + \beta(\lambda_k) \frac{\partial}{\partial \lambda_k} + \gamma_m(\lambda_k) m_k \frac{\partial}{\partial m_k} \right] P(\lambda_k, m_k, \mu) = 0 \quad (34)$$

So, we know that P satisfies (31), i.e.,

$$P(p_i, \lambda_k, m_k, \mu_0 e^{-t}) = P(p_i, \bar{\lambda}(t), \bar{m}(t), \mu_0) \quad (35)$$

~~15.e Green's functions for rescaled momenta~~
We want to study

For example, let's consider the ^(Laurent series of) scalar propagator $\Delta(p^2)$:

$$\Delta(p^2) = \frac{R^2}{p^2 - m_F^2} + \bar{\Delta}(p^2, \lambda_k, m_k, \mu) \quad (36)$$

that satisfies

$$\left[\mu \frac{\partial}{\partial \mu} + \beta(\lambda_k) \frac{\partial}{\partial \lambda_k} + \gamma_m(\lambda_k) m_k \frac{\partial}{\partial m_k} + 2\gamma(\lambda_k) \right] \Delta = 0 \quad (37)$$

Notice that we have $2\gamma(\lambda_k)$ since it is the connected Green's function and not $\Gamma_k^{(2)}$! This is due to

$$\Delta_R = (Z_3^{-1/2})^2 \Delta_B \text{ in contrast with } (P)!$$

(37) into (36) $\Rightarrow$

15.10

$$(\mathcal{D} + 2\gamma)\bar{\Delta} + \frac{1}{p^2 m_F^2} (\mathcal{D} + 2\gamma) R^2 + \frac{2m_F (\mathcal{D} m_F)}{(p^2 m_F^2)^2} = 0$$

moment
 $\Rightarrow$
we

$$(\mathcal{D} + 2\gamma)\bar{\Delta} = 0$$

$$(\mathcal{D} + 2\gamma) R^2 = 0$$

$$\mathcal{D} m_F = 0 \quad (38)$$

As expected, the position of the pole (physical mass) satisfies (3)

15.e Green's functions for repeated momenta

We want to study the effect of $p_i \rightarrow p p_i$ in

$T_R^{(n)}(p p_i, \lambda_R, m_R, \mu)$. We know the canonical dimension of $T_R^{(n)}$:

$[T_R^{(n)}] = [M]^D$. So, using dimensional analysis we

know that

$$T_R^{(n)}(p p_i, \lambda_R, m_R, \mu) = \mu^D f\left(p^2 \frac{p_i \cdot p_i}{\mu^2}, \frac{m_R}{\mu}, \lambda_R\right) \quad (39)$$

Since $T_R^{(n)}$ is a homogeneous function of order D in the variables p, m_R and μ we know that

$$\left(p \frac{\partial}{\partial p} + m_R \frac{\partial}{\partial m_R} + \mu \frac{\partial}{\partial \mu} - D\right) T_R^{(n)}(p p_i, \lambda_R, m_R, \mu) = 0 \quad (40)$$

subtracting $\mu \frac{\partial \Gamma_K^{(u)}}{\partial \mu}$ from (40) and inserting into (41) $\Rightarrow$

$$\left[-\rho \frac{\partial}{\partial \rho} + \beta(\lambda_k) \frac{\partial}{\partial \lambda_k} + (\gamma_m(\lambda_k) - 1) \frac{\partial}{\partial m_k} - \eta \gamma(\lambda_k) + D \right] \Gamma_R^{(u)}(\rho p_i, \lambda_k, \mu_k, \mu) = 0 \quad (41)$$

now $t = \ln \rho \Rightarrow \frac{\partial}{\partial t} = \rho \frac{\partial}{\partial \rho}$. Analogously to the previous case

$$\Gamma_R^{(u)}(e^t p_i, \lambda_k, \mu_k, \mu) = \Gamma_R^{(u)}(p_i, \bar{\lambda}(t), \bar{m}(t), \mu) \exp \left[D t - \eta \int_0^t \gamma(\bar{\lambda}(t')) dt' \right] \quad (42)$$

solving for $\mu_k = \rho = \gamma = 0$ (wavelength non-interaction)

where $\bar{\lambda}$ is given by (25) and

$$\bar{m} = \frac{\bar{m}(t)}{e^t}$$

as in (26).

we could also do (32)

$$\Gamma_R^{(u)}(e^t p_i, \lambda_k, \mu_k, \mu) \stackrel{!}{=} \Gamma_R^{(u)}(e^t p_i, \bar{\lambda}(t), \bar{m}(t), e^t \mu) \exp \left[-\eta \int_0^t \gamma(t') dt' \right]$$

$$= \Gamma_R^{(u)}(e^t p_i, \bar{\lambda}(t), e^t e^{-t} \bar{m}(t), e^t \mu) \exp \left[-\eta \int_0^t \gamma(t') dt' \right]$$

homogeneous

$$= (e^t)^D \Gamma_R^{(u)}(p_i, \bar{\lambda}(t), \bar{m}(t), \mu) \exp \left[-\eta \int_0^t \gamma(t') dt' \right] \quad QED!$$

Anomalous dimensions:

15.11.0

Let's focus on the last term of (42). In $d=4$

$$[\varphi] = m^{-1} \quad [m] = m^1 \quad \text{and the action}$$

$$S = \int d^4x \left[-\frac{1}{2} \varphi (\Box + m^2) \varphi - \frac{\lambda}{4!} \varphi^4 \right]$$

S is invariant under $x^\mu \rightarrow \frac{1}{\lambda} x^\mu$, $m \rightarrow \lambda m$, $\varphi \rightarrow \lambda \varphi$ *

$\partial_\mu \rightarrow \lambda \partial_\mu$, $\lambda \rightarrow \lambda$

This transformation is called Dilation (D)

$$D: \varphi \rightarrow \lambda^{do} \varphi \quad \text{where } do \text{ is the canonical scaling dimension of the field.}$$

Under D : $G_c^{(n)}(x_1 \dots x_n) \rightarrow \lambda^{ndo} G_c^{(n)}(\dots)$

$$\Rightarrow \Gamma^{(n)}(x_1 \dots x_n) \Rightarrow \lambda^{-ndo} \Gamma^{(n)}(\dots)$$

Since we have to amputate n external legs with $G_c^{(2)-1}$! We expect classically that

$$\Gamma^{(n)} \cong m^a \lambda^b x_1^{c_1} \dots x_n^{c_n} \mu^\gamma \quad \text{with } a + c_1 + \dots + c_n = -n. \text{ However, in the}$$

quantum theory we ~~also~~ also have the renormalization scale μ :

$$\Gamma^{(n)} \cong m^a \lambda^b x_1^{c_1} \dots x_n^{c_n} \mu^\gamma$$

hence, we have $D: \Gamma_n \rightarrow \lambda^{n-\gamma} \Gamma_n$ since just m and x are scaled in *. This is the origin of the last term in (42)!

From TQC1, we know that

$$Z_1 = 1 + \frac{3\lambda\kappa}{16\pi^2\epsilon}$$

$$Z_0 = 1 + \frac{\lambda\kappa}{16\pi^2\epsilon} \quad (43)$$

$$Z_3 = 1$$

For the β function: $a_1 = \frac{3\lambda\kappa}{16\pi^2}$
 $Z_\lambda = Z_0^{-2} Z_1$

$$\Rightarrow \beta(\lambda\kappa) = \lambda\kappa^2 \frac{da_1}{d\lambda\kappa} = \frac{3\lambda\kappa^2}{16\pi^2} \quad (44)$$

For the γ function: $Z_3^{(1)} \equiv 0 \xrightarrow{(22)} \gamma = 0 \quad (45)$

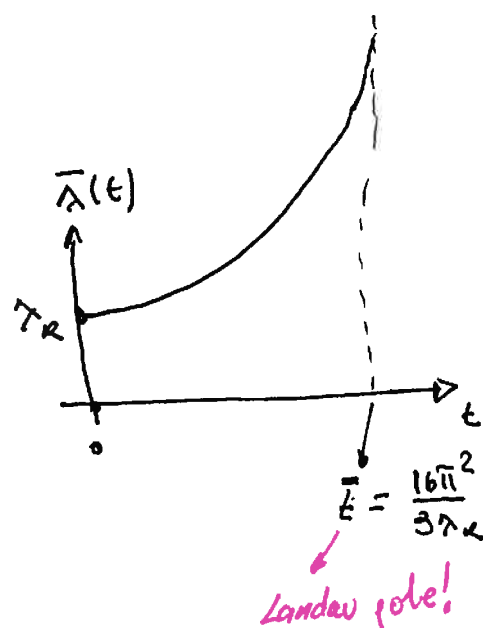
For the γ_m function: $Z_m^{(1)} = \frac{\lambda\kappa}{16\pi^2} \xrightarrow{(23)} \gamma_m(\lambda\kappa) = \frac{1}{2} \frac{\lambda\kappa}{16\pi^2} \quad (46)$

Now turning to the evolution eq'ns:

$$\frac{\partial \bar{\lambda}}{\partial t} = \frac{3\bar{\lambda}^2}{16\pi^2} \Rightarrow \int_{\lambda\kappa}^{\bar{\lambda}(t)} \frac{d\bar{\lambda}}{\bar{\lambda}^2} = \frac{3}{16\pi^2} \int_0^t dt'$$

$$\Rightarrow -\frac{1}{\bar{\lambda}} + \frac{1}{\lambda\kappa} = \frac{3t}{16\pi^2}$$

$$\Rightarrow \bar{\lambda}(t) = \frac{\lambda\kappa}{1 - \frac{3t\lambda\kappa}{16\pi^2}} \quad (47)$$



On the other hand

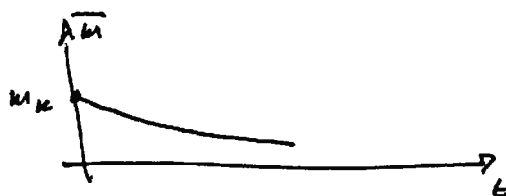
15.15

$$\frac{d\bar{m}}{dt} = \gamma_m(\bar{\lambda}) \bar{m} = \frac{1}{32\pi^2} \frac{\lambda_R}{1 - \frac{3\lambda_R}{16\pi^2} t} \bar{m}$$

$$\int_{m_R}^{\bar{m}} \frac{dy}{y} = \int_0^t dt' \frac{\lambda_R}{32\pi^2} \frac{1}{1 - \frac{3\lambda_R}{16\pi^2} t'}$$

$$\ln\left(\frac{\bar{m}}{m_R}\right) = \frac{\lambda_R}{32\pi^2} \left(-\frac{16\pi^2}{3\lambda_R}\right) \ln\left(1 - \frac{3\lambda_R}{16\pi^2} t\right)$$

$$\bar{m}(t) = m_R \left(1 - \frac{3\lambda_R}{16\pi^2} t\right)^{-1/6} \quad (48)$$



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15.6 Renormalization Group Flow

Starting at small couplings we can have:

i) $\beta > 0$ (like $\lambda\phi^4$, QED), as we saw the coupling grows as indicated in the previous page.

ii) $\beta = 0 \Rightarrow$ the coupling constant is scale independent

iii) $\beta < 0 \Rightarrow$ the coupling decreases at high energies.

Let's make our analyses more general. ~~Let~~ and concentrate on the IR ($t \rightarrow -\infty$) and UV ($t \rightarrow +\infty$) limits. Here, we assume that (25) is valid for $-\infty < t < +\infty$. From ~~the~~ eqn (27)

$$t = \int_{\lambda_R}^{\bar{\lambda}(t, \lambda_R)} \frac{dx}{\beta(x)} \quad (27)$$

for $t \rightarrow \pm\infty$, (27) implies that either

(i) $\bar{\lambda}$ goes to infinity

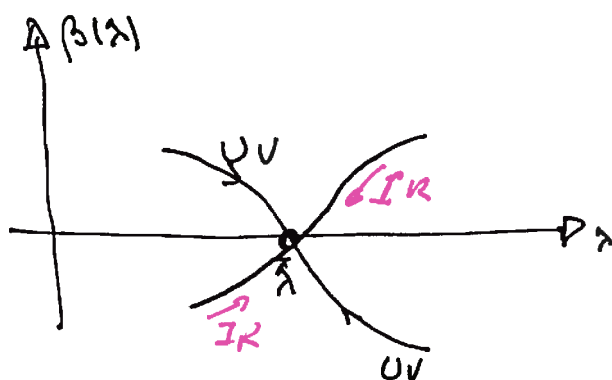
or (ii) $\bar{\lambda}$ approaches a zero of β

A zero of $\beta(x)$ is called a fixed point, that is classified as:

(i) UV stable fixed point if $\lim_{t \rightarrow +\infty} \bar{\lambda}(t, \lambda_+) = \bar{\lambda} \equiv \lambda_+$

(ii) IR stable fixed point for $\lim_{t \rightarrow -\infty} \bar{\lambda}(t, \lambda_-) = \bar{\lambda} \equiv \lambda_-$

Graphically



$$\beta'(\lambda)|_{\bar{\lambda}} < 0 \Rightarrow \text{UV fixed point}$$

$$\Rightarrow \beta'(\lambda)|_{\bar{\lambda}} > 0 \Rightarrow \text{IR fixed point}$$

We use P.T around $\lambda=0$ to obtain β . We classify the theory as

i) asymptotically free theory if $\lambda=0$ is an UV stable fixed point (eg QCD)

ii) IR stable if $\lambda=0$ is an IR stable fixed point. (eg $\lambda\Phi^4$, QED)



Consider QED in the gauge $-\frac{1}{a}(\partial_\mu A^\mu)^2$ with a mass independent renormalization scheme. We have

$$\Gamma_R^{(n_A, n_F)}(p_i, \alpha_R, m_R, \mu, a) = Z_3^{\frac{n_A}{2}} Z_2^{\frac{n_F}{2}} T_B^{(n_A, n_F)}(p_i, \alpha_B, m_B, a_B) \quad (49)$$

with $\alpha_R = \frac{e^2}{4\pi}$ $\alpha_B = \mu^\epsilon Z_3^{-1} \alpha_R$

$$m_B = Z_m^{-1} m_R$$

$$A_{\mu B} = Z_3^{1/2} A_{\mu R} \quad (50)$$

$$\psi_B = Z_2^{1/2} \psi_R$$

$$a = Z_3^{-1} a_B$$

Analogously to the procedure in section (15.5), we have that $\mu \frac{d}{d\mu}$ (49) leads to

$$\left[\mu \frac{\partial}{\partial \mu} + \beta(\alpha_R, a) \frac{\partial}{\partial \alpha_R} + \gamma_m(\alpha_R, a_R) m_R \frac{\partial}{\partial m_R} + \delta(\alpha_R, a_R) \frac{\partial}{\partial a_R} - n_A \gamma_A(\alpha_R, a_R) - n_F \gamma_F(\alpha_R, a_R) \right] \Gamma_R^{(n_A, n_F)}(p_i, \alpha_R, m_R, \mu, a_R) = 0 \quad (51)$$

From now on we drop the subscript R! We have that

$$\beta(\alpha, a) = \lim_{\epsilon \rightarrow 0} \beta(\alpha, a, \epsilon) = \lim_{\epsilon \rightarrow 0} \mu \frac{d}{d\mu} \alpha$$

$$\gamma_m(\alpha, a) = \lim_{\epsilon \rightarrow 0} \mu \frac{d}{d\mu} \ln z_m$$

$$\gamma_a(\alpha, a) = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \mu \frac{d}{d\mu} \ln z_3$$

(52)

$$\gamma_F(\alpha, a) = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \mu \frac{d}{d\mu} \ln z_c$$

$$\delta(\alpha, a) = \lim_{\epsilon \rightarrow 0} \mu \frac{da}{d\mu} = \lim_{\epsilon \rightarrow 0} \left(-a \mu \frac{d}{d\mu} \ln z_3 \right)$$

The sol'n of (51) ~~requires~~ requires solving

$$\frac{\partial}{\partial t} \bar{\alpha}(t, \alpha) = \beta(\bar{\alpha}) \quad \text{with} \quad \bar{\alpha}(0, \alpha) = \alpha$$

$$\frac{\partial}{\partial t} \bar{m}(t, \alpha, m) = \bar{m} \gamma_m(\bar{\alpha}) \quad \text{with} \quad \bar{m}(0, \alpha, m) = m$$

(53)

$$\frac{\partial}{\partial t} \bar{a}(t, \alpha, a) = \delta(\bar{\alpha}, \bar{a}) \quad \text{with} \quad \bar{a}(0, \alpha) = a$$

to obtain that (for ~~every~~ $\mu = \mu_0 e^{-t}$) $(t = -\ln \frac{\mu}{\mu_0})$

$$\Gamma_R^{(n_A, n_F)}(p_i, \alpha_R, m_R, a_R, \mu_0 e^{-t}) = \Gamma_R^{(n_A, n_F)}(p_i, \bar{\alpha}(t), \bar{m}(t), \bar{a}(t), \mu_0) \quad \textcircled{x}$$

$$\textcircled{x} \exp \left[-n_A \int_0^t dt' \gamma_A(\bar{\alpha}(t'), \bar{a}(t')) - n_F \int_0^t dt' \gamma_F(\bar{\alpha}(t'), \bar{a}(t')) \right] \quad (54)$$

$$\beta = \mu \frac{d}{d\mu} \alpha = \mu \frac{d}{d\mu} (\bar{\mu}^{\epsilon} z_3 \alpha_B)$$

$$= -\epsilon \alpha_R + \alpha_R \mu \frac{d}{d\mu} \ln z_3 \quad (10. QED)$$

$$\gamma_A = \frac{1}{2} \mu \frac{d}{d\mu} \ln z_3 \quad \checkmark$$

Since $\alpha_B = \mu^{\epsilon} Z_3^{-1} \alpha$ and comparing (10) with (14) $\Rightarrow Z_{\lambda}^{\text{QED}} = Z_3^{-1}$ (55) 15.17

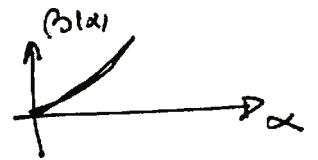
$$\Rightarrow \beta_{\text{QED}} = \alpha_R^2 \frac{d Z_{\lambda}^{\text{QED}(1)}}{d \alpha_R} \quad (56)$$

We know that, at the one-loop level, $Z_{\lambda}^{\text{QED}} = 1 - \frac{\alpha^2}{3\pi\epsilon}$ in MS.

See page B.17, eq (41) $\Rightarrow Z_{\lambda}^{\text{QED}} = 1 + \frac{2\alpha}{3\pi\epsilon}$

$$\Rightarrow \beta_{\text{QED}}^{(\alpha)} = \frac{2\alpha^2}{3\pi} \quad (57)$$

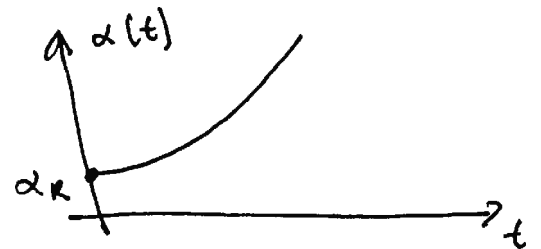
Notice that $\beta_{\text{QED}}^{(\alpha)} > 0 \Rightarrow \text{QED is IR stable.}$



In this case, the RGE for α reads $\Rightarrow$ (53)

$$\frac{d\bar{\alpha}}{dt} = \frac{2\bar{\alpha}^2}{3\pi} \Rightarrow -\frac{1}{2(t)} + \frac{1}{\alpha_R} = \frac{2t}{3\pi}$$

$$\Rightarrow \bar{\alpha} = \frac{\alpha_R}{1 - \frac{2\alpha_R t}{3\pi}} \quad (58) \Rightarrow$$



Compare with eq'n (3) on page (15.1).

Notice: Z_3 is gauge-independent since it renormalises the gauge invariant operator $F_{\mu\nu} F^{\mu\nu}$. Therefore, β_{QED} is also gauge independent; see (55-56)!

Instead of α we could have considered e . 15.18

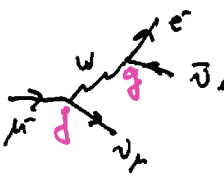
In this case

$$\beta(e) = \mu \frac{de}{d\mu} \Rightarrow \beta(\alpha) = \mu \frac{d\alpha}{d\mu} = \frac{2e}{4\pi} \mu \frac{de}{d\mu} \Rightarrow \beta(e) = \frac{2\pi}{e} \beta(\alpha)$$

$$\Rightarrow \beta(e) \stackrel{\text{1-loop}}{=} \frac{4}{3} \frac{e^3}{16\pi^2} \quad (59)$$

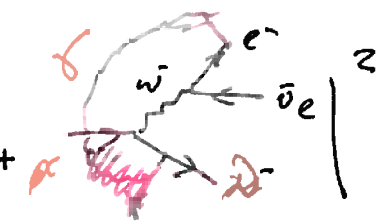
5.1 RGE for higher-order operators

Motivation: In the SM, μ decay rate is given by

$$(\mu^- \rightarrow \bar{\nu}_e e^- \nu_\mu) = \frac{1}{2m_\mu} \int d\phi_3 \left| \text{diagram} \right|^2 = \left(\frac{\sqrt{2} g^2}{8 m_W^2} \right)^2 \frac{m_\mu^5}{192 \pi^3} \quad (59a)$$


Large log's can appear since $m_\mu \gg m_e$. Consider $m_W \gg m_\mu$.

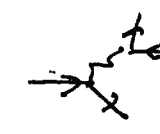
$$(\mu^- \rightarrow \bar{\nu}_e e^- \nu_\mu) = \frac{1}{2m_\mu} \int d\phi_3 \left| \text{diagram} + \text{diagram} \right|^2$$

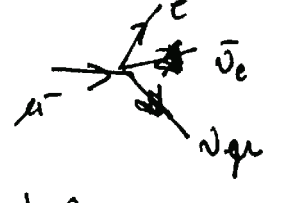
$$= \left(\frac{\sqrt{2} g^2}{8 m_W^2} \right)^2 \frac{m_\mu^5}{192 \pi^3} \left[1 + A \frac{\alpha}{4\pi} \ln \frac{m_W^2}{m_\mu^2} + \dots \right] \quad (60)$$


Let's use the RGE to not only estimate A but also to sum terms

$$\left(A \frac{\alpha}{4\pi} \ln \frac{m_W^2}{m_\mu^2} \right)^n$$

Effective operator (Fermi theory)

Since $m_W \gg m_\mu$, the propagator $\frac{i}{p^2 - m_W^2}$ in  is effectively $-\frac{i}{m_W^2}$

leading to  which is associated to the operator

$$\mathcal{L}_{4F} = \frac{G_F}{\sqrt{2}} \bar{\Psi}_\mu \gamma^\alpha \gamma_5 \Psi_\nu \bar{\Psi}_e \gamma_\alpha \gamma_5 \Psi_e + h.c. \quad (61)$$

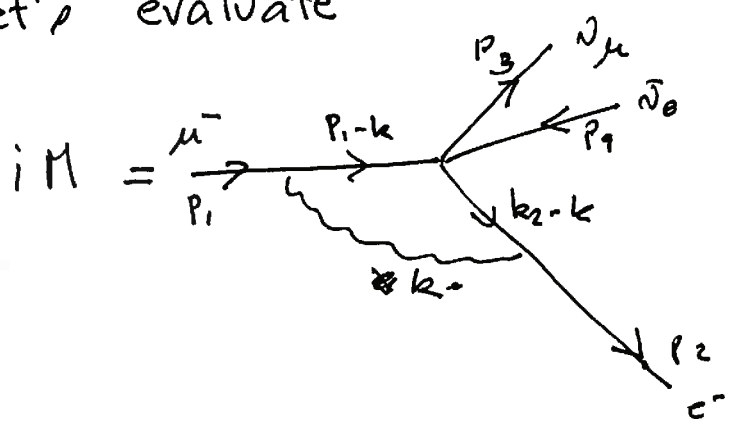
with $\gamma_5 = \frac{1-\gamma_5}{2}$, $G_F = \frac{\sqrt{2} g^2}{8 m_W^2} = 1.166 \times 10^{-5} \text{ GeV}^{-2}$. From (61) we can obtain easily (59a). Do it! (61) is the Fermi 4-fermion model!

to simplify our life let's consider

01/10/19

$$\mathcal{L}_4 = \frac{G}{\sqrt{2}} (\bar{\Psi}_\mu \Psi_{\nu_\mu}) (\bar{\Psi}_e \Psi_e) + h.c. \quad (62)$$

Let's evaluate



$$iM = \frac{G}{\sqrt{2}} e_R^2 \mu \int \frac{d^4 k}{(2\pi)^4} \frac{\bar{u}(p_2) \gamma^\mu (\not{p}_2 - \not{k} + m_e) (\not{p}_1 - \not{k} + m_\mu) \gamma_\mu u(p_1) \bar{u}(p_3) \gamma^\nu \gamma_5 u(p_4)}{[(p_1 - k)^2 - m_\mu^2] [(p_2 - k)^2 - m_e^2] k^2} \quad (63)$$

Since we just need the pole to get the β function we set all masses and external momenta to zero!

$$M = M_0 \left(-i e_R^2 \mu \int \frac{d^4 k}{(2\pi)^4} \frac{1}{k^4} \right) + \text{finite}$$

$$1^{st} M_0 = \frac{G}{\sqrt{2}} \bar{u}(p_2) u(p_1) \bar{u}(p_3) u(p_4)$$

So we have $M = M_0 \left(\frac{e^2}{2\pi^2} \mu^\epsilon \frac{1}{\epsilon} \right) + \text{finite}$ (64)

To remove the divergence we write

$$\mathcal{L} = \frac{G_R}{\sqrt{2}} \bar{\psi}_\mu^R \psi_e^R \bar{\psi}_{\nu e}^R \psi_e^R + h.c.$$

adding $\bar{\mu}^\epsilon$ to $\mu \frac{d}{d\mu} G_{\text{bare}}$

$\Rightarrow$ ϵ term in (64)
that disappears in
the limit $\epsilon \rightarrow 0$!

So to render (64) finite we have

$$Z_0 = 1 - \frac{e^2}{16\pi^2} \frac{8}{\epsilon} \quad (66)$$

Notice that (65) is

$$\mathcal{L} = \frac{G_R}{\sqrt{2}} \frac{Z_0}{(Z_{2\mu} Z_{2e} Z_{2\nu e} Z_{2\nu\mu})^{1/2}} \begin{pmatrix} \bar{\psi}_\mu^B & \psi_e^B \end{pmatrix} \begin{pmatrix} \bar{\psi}_{\nu e}^B & \psi_{\nu\mu}^B \end{pmatrix} \quad (66)$$

Since the $\nu_{e,\mu}$ are neutral $Z_{2\nu e} = Z_{2\nu\mu} = 1$. Moreover,

$$Z_{2e} = Z_{2\mu} = Z_2$$

The RG ϵ group reads

$$0 = \mu \frac{d}{d\mu} \left(\underbrace{\frac{G_R Z_0}{Z_2}}_{G_{\text{bare}}} \right)$$

$\Rightarrow$

$$0 = \frac{G_R Z_0}{Z_2} \left[\frac{\mu}{G_R} \frac{d G_R}{d\mu} + \frac{1}{Z_0} \frac{d Z_0}{d\epsilon} \cdot \mu \frac{d\epsilon}{d\mu} - \frac{1}{Z_2} \frac{d Z_2}{d\epsilon} \cdot \mu \frac{d\epsilon}{d\mu} \right]$$

$$\Rightarrow \frac{\mu}{G_R} \frac{dG_R}{d\mu} = \left(-\frac{1}{z_G} \frac{dz_G}{d\epsilon_R} + \frac{1}{z_2} \frac{dz_2}{d\epsilon_R} \right) \beta(\epsilon_R) \quad (67)$$

Now let's analyze (67) to the lowest order:

$$z_2 = 1 - \frac{\epsilon_R^2}{16\pi^2} \frac{2}{\epsilon}$$

$$z_G = 1 - \frac{\epsilon_R^2}{16\pi^2} \frac{8}{\epsilon}$$

$\Rightarrow$

$$\beta(\epsilon_R) = -\frac{\epsilon}{2} \epsilon_R + \frac{\epsilon_R^3}{12\pi^2}$$

lowest
order β

$$\frac{\mu}{G_R} \frac{dG_R}{d\mu} = \left(-\frac{\epsilon}{2} \epsilon_R \right) \times \left(\cancel{-\frac{2}{\epsilon}} \left(-\frac{\epsilon_R}{16\pi^2} \frac{8}{\epsilon} \right) - 2 \frac{\epsilon_R}{16\pi^2} \frac{2}{\epsilon} \right)$$

$$\Rightarrow \frac{\mu}{G_R} \frac{dG_R}{d\mu} = -\frac{3\epsilon_R^2}{8\pi^2} = -\frac{3\alpha_R}{2\pi} \equiv \gamma_G$$

$$\Rightarrow \frac{dG_R}{G_R} = \gamma_G(\alpha(\mu)) \frac{d\mu}{\mu} = \frac{\gamma(\bar{\alpha})}{\beta(\bar{\alpha})} d\bar{\alpha}$$

$d\bar{\alpha} = \beta(\bar{\alpha}) \frac{d\mu}{\mu}$

$$\stackrel{\mu}{d\mu} \Rightarrow \Rightarrow G_R(\mu) = G_R(\mu_0) \exp \left[\int_{\alpha(\mu_0)}^{\bar{\alpha}(\mu)} d\bar{\alpha} \frac{\gamma(\bar{\alpha})}{\beta(\bar{\alpha})} \right] \Rightarrow \frac{G_R(\mu)}{G_R(\mu_0)} = \exp \int_{\alpha(\mu_0)}^{\alpha(\mu)} d\alpha \frac{-\frac{3\alpha}{2\pi}}{\frac{2\alpha^2}{3\pi}}$$

As seen, $\beta(\alpha) = \frac{2\alpha^2}{3\pi}$

$$\Rightarrow G(\mu) = G(\mu_0) \left(\frac{\alpha(\mu)}{\alpha(\mu_0)} \right)^{-\frac{7}{4}} \quad (68)$$

15.22

15.1 Wilsonian RGE

Basic idea: remove "high energy" modes while preserving their effect in the "low energy" dynamics. To make the ~~very~~ distinction between low and high energy we work in the euclidean space. For the ϕ^4 model

$$S[\phi] = \int d^4x \left\{ \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - \frac{\lambda}{4!} \phi^4 \right\} \quad (69)$$

we perform the Wick rotation $x_4 = i x_0 \Rightarrow dx_0 = -i dx_4$ and (69) reads

$$S_E[\phi] = i \int d^4x \left\{ \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + \frac{m^2}{2} \phi^2 + \frac{\lambda}{4!} \phi^4 \right\} \quad (70)$$

is the hamiltonian of the system ~~apart~~ from the dx_4 integration!
now, we write the generating functional as

$$Z = \int \mathcal{D}\phi \, e^{-S_E[\phi]} \quad (71)$$

so, Z is the "partition function" of the system! To continue, we consider an overall cutoff Λ

$$Z = \int [\mathcal{D}\phi]_\Lambda \, e^{-S_E[\phi]} \quad (72)$$

with

$$[\mathcal{D}\phi]_\Lambda = \prod_{|k| < \Lambda} d\phi(k) \quad (73)$$

Fourier modes of the system

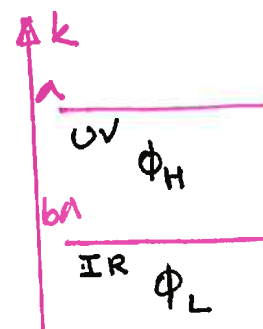
: assume that $\Lambda \gg m$. We can separate the low and high energy modes $\Rightarrow$

$$\phi(k) = \phi_L(k) + \phi_H(k) \quad (73)$$

where $\phi(x) = \int \frac{d^d k}{(2\pi)^d} e^{ikx} \phi(k)$ (74)

and $\phi_H(k) \equiv \begin{cases} \phi(k) & \text{for } |k| > b\Lambda \\ 0 & \text{for } |k| < b\Lambda \end{cases}$ (75)

$$\phi_L(k) \equiv \begin{cases} \phi(k) & \text{for } |k| < b\Lambda \\ 0 & \text{for } |k| > b\Lambda \end{cases} \quad (76)$$



with $\alpha b < 1$.

Idea: we split the integration

$$[\mathcal{D}\phi]_\Lambda = \mathcal{D}\phi_L \mathcal{D}\phi_H \quad (77)$$

and we integrate out the high-energy modes! In practical terms,

$$Z = \int \mathcal{D}\phi_L \int \mathcal{D}\phi_H e^{-\int d^d x \left\{ \frac{1}{2} (\partial_\mu \phi_L + \partial_\mu \phi_H)^2 + \frac{1}{2} m^2 (\phi_L + \phi_H)^2 + \frac{\lambda}{4!} (\phi_L + \phi_H)^4 \right\}}$$

due to the orthogonality conditions in momentum space, terms like $\partial_\mu \phi_L \partial_\mu \phi_H$ and $\phi_L \phi_H$ do not contribute and we are left with:

$$Z = \int \mathcal{D}\phi_L e^{-S_E[\phi_L]} \int \mathcal{D}\phi_H e^{-\int d^d x \left\{ \frac{1}{2} (\partial_\mu \phi_H)^2 + \frac{m^2}{2} \phi_H^2 + \lambda \left(\frac{1}{6} \phi_L \phi_H^3 + \frac{1}{4!} \phi_L^2 \phi_H^2 + \frac{1}{6} \phi_L \phi_H^3 + \frac{1}{4!} \phi_H^4 \right) \right\}} \quad (78)$$

In principle, we can perform the $\mathcal{D}\phi_H$ integral obtaining that

$$Z = \int \mathcal{D}\phi_L e^{-S_E^{\text{eff}}[\phi_L]} \quad (79)$$

here

[15.24]

$$S_E^{\text{eff}}[\phi_L] = \int d^d x \left\{ \frac{1}{2} \partial_\mu \phi_L \partial_\mu \phi_L + \frac{m^2}{2} \phi_L^2 + \frac{\lambda}{4!} \phi_L^4 + \mathcal{O}(\lambda) \text{ corrections} \right\} \quad (80)$$

from $\mathcal{D}\phi_H$!

How do we evaluate that?

since $\Lambda \gg m^2$ we can neglect the $m^2 \phi_H^2$ term in (78), thus the quadratic term reads

$$\begin{aligned} \int d^d x \frac{1}{2} \partial_\mu \phi_H \partial_\mu \phi_H &= \frac{1}{2} \int d^d x \int \frac{d^d k}{(2\pi)^d} \int \frac{d^d k'}{(2\pi)^d} \phi_H(k) (-ik) \phi(-ik') \phi_H(k') e^{i(k+k') \cdot x} \\ &= \frac{1}{2} \int \frac{d^d k}{(2\pi)^d} \phi_H(k) k^2 \phi_H(-k) \end{aligned} \quad (81)$$

for a free theory

$$Z_{\text{free}} = \int \mathcal{D}\phi_H e^{- \int \frac{d^d k}{(2\pi)^d} \left\{ \frac{1}{2} \phi_H(k) k^2 \phi_H(-k) + \int d^d x \phi_H(x) \right\}} \quad (82)$$

hence, the propagator in momentum space is

$$D_F(p+p) = \frac{(2\pi)^d}{k^2} \delta^{(d)}(k+p) \quad (83)$$

expanding the exponential in (78) we get terms like:

$$i) - \frac{\lambda}{4} \int d^d x \phi_L^2(x) \phi_H(x) \phi_H(x) \Rightarrow \int \mathcal{D}\phi_H e^{-\frac{1}{2} \int \frac{d^d k}{(2\pi)^d} \phi_H(k) k^2 \phi_H(-k)} \phi_H(k) \phi_H^{(0)} \rightarrow D_F(p+p)$$

$$\Rightarrow - \frac{\lambda}{4} \int \frac{d^d k_1}{(2\pi)^d} \phi_L(k_1) \phi_L(-k_1) \int \frac{d^d k_2}{(2\pi)^d} \frac{1}{k_2^2} = - \frac{\mu}{2} \int d^d x \phi_L^2(x)$$

$$\text{with } \mu = \frac{\lambda}{2} \int \frac{d^d k_2}{(2\pi)^d} \frac{1}{k_2^2} = \lambda \frac{1}{(2\pi)^d \Gamma(d/2)} \frac{(1-S^{d-2}) \Lambda^{d-2}}{d-2} \quad (84)$$

To see that write (82) as

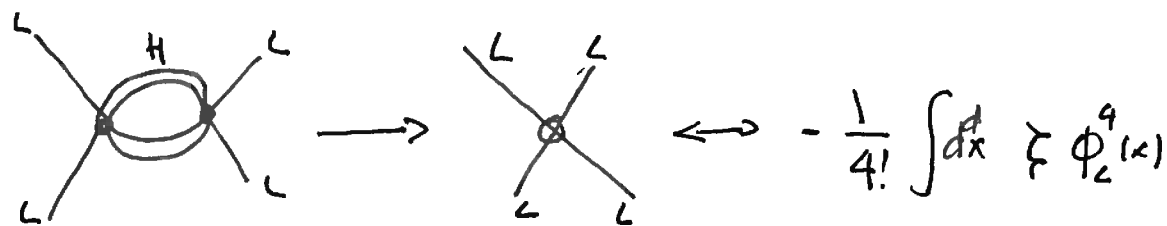
$$\int \frac{d^d k}{(2\pi)^d} \int \frac{d^d k'}{(2\pi)^d} \frac{1}{2} \phi_H(k) \phi_H(k') \text{ with } \phi = k^2 \phi_H(k)$$

This corresponds to ~~the~~ start from



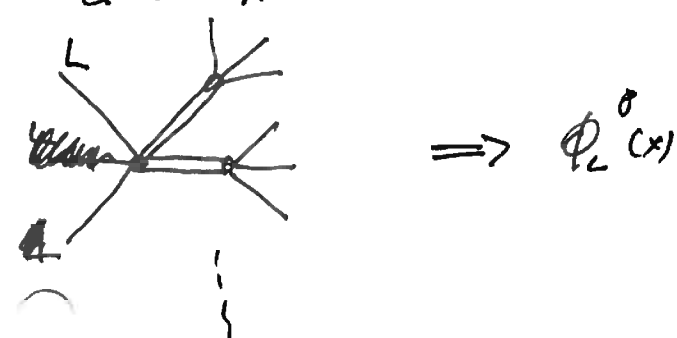
Furthermore, this one contribution to the expansion $e^{\frac{-i}{\hbar} \int d^d x \mathcal{L}_2(x)}$

ii) Analogously, in order λ^2 we have contributions of the form



with $\mathcal{L} = - \frac{3\lambda^2}{(4\pi)^{d/2}} \frac{1}{\Gamma(\frac{d}{2})} \frac{(1 - 5^{d-4})}{d-4} \lambda^{d-4}$

iii) Another possibility for order λ^2 is



At the end we can write

$$S_E^{eff}[\phi_L] = \int d^d x \left\{ \frac{1}{2} (1 + \Delta z) (\partial_\mu \phi_L)^2 + \frac{1}{2} (m^2 + \Delta m^2) \phi_L^2 + \frac{1}{4!} (\lambda + \Delta \lambda) \phi_L^4 + \Delta c (\partial_\mu \phi_L \partial_\mu \phi_L^2) + \Delta d \phi_L^6 + \dots \right\}$$

here the integration over ϕ_k gives rise to $\Delta z, \Delta m^2, \Delta \lambda, \Delta c, \Delta D, \dots$

Notice, that we started with a renormalizable theory but S_E^{eff} is not renormalizable in the usual sense!

Renormalization flow: let's analyze the effects of changing b ; for that

$$k' \equiv \frac{k}{b} \quad \text{and} \quad x' = bx \quad (\text{with this } |k'| < \Lambda)$$

$\Downarrow$

$$d^d x = b^{-d} d^d x' \quad \frac{\partial}{\partial x^\mu} = \frac{b}{\partial x'^\mu} \equiv b \partial'_\mu \quad (86)$$

$\Downarrow$

$$S_E^{\text{eff}}[\phi_L] = \int d^d x' b^{-d} \left\{ \frac{1}{2} (1 + \Delta z) b^2 (\partial'_\mu \phi_L)^2 + \frac{1}{2} (m^2 + \Delta m^2) \phi_L^2 + \frac{1}{4!} (\lambda + \Delta \lambda) \phi_L^4 \right. \\ \left. + \Delta c b^2 (\partial'_\mu \phi_L)^2 + \Delta D \phi_L^6 + \dots \right\} \quad (87)$$

to impose that the field ϕ_L is canonically normalized we do define

$$\phi'_L = [b^{2-d} (1 + \Delta z)]^{1/2} \phi_L \quad (88)$$

so that $(88) + (87) \Rightarrow$

$$S_E^{\text{eff}}[\phi'_L] = \int d^d x' \left\{ \frac{1}{2} (\partial'_\mu \phi'_L)^2 + \frac{1}{2} m'^2 \phi'^2_L + \frac{\lambda'}{4!} \phi'^4_L + c' (\partial'_\mu \phi'_L)^2 + D' \phi'^6_L + \dots \right\} \quad (89)$$

with

$$\begin{aligned} m'^2 &= (m^2 + \Delta m^2) (1 + \Delta z)^{-1} b^{-2} \\ \lambda' &= (\lambda + \Delta \lambda) (1 + \Delta z)^{-2} b^{d-4} \\ c' &= \Delta c (1 + \Delta z)^{-2} b^d \\ D' &= \Delta D (1 + \Delta z)^{-3} b^{2d-6} \end{aligned} \quad (90)$$

The combination of the integration ~~and~~ of higher 15.21 energy modes and the re-scaling leads to a continuous transformation ~~that is the RGE~~ of the Lagrangian that is the RGE. Considering b infinitesimally close to 1 we get the RGE.

For further details see Prokin section 12.1.