

Analytic Properties of the Scattering Amplitude - Antonio Pedrosa

1. Scattering operator and partial-wave expansion

1.1 Møller wave operators, scattering operator

In scattering theory we're concerned with the study of asymptotic states. If the system is described by $H = H_0 + V$, where H_0 is the free hamiltonian (that may be channel dependent, but in the simplest elastic two body case is just the relative motion kinetic energy) and V is an interaction, we may define the general evolution operator $U(t) = e^{iHt}$ (using $\hbar=1$) and the free (asymptotic) evolution operator as $U_0(t) = e^{iH_0 t}$. If a state $|\psi\rangle = |\psi(t=0)\rangle$ that is in the interaction region at $t=0$ originated from a asymptotic free state $|\psi_{in}\rangle$, we have

$$\lim_{t \rightarrow -\infty} U(t) |\psi\rangle = \lim_{t \rightarrow -\infty} U_0(t) |\psi_{in}\rangle, \quad \text{and}$$

So we define the Møller operator as the limit $\Omega_+ := \lim_{t \rightarrow -\infty} U^\dagger(t) U_0(t)$, that "evolves"

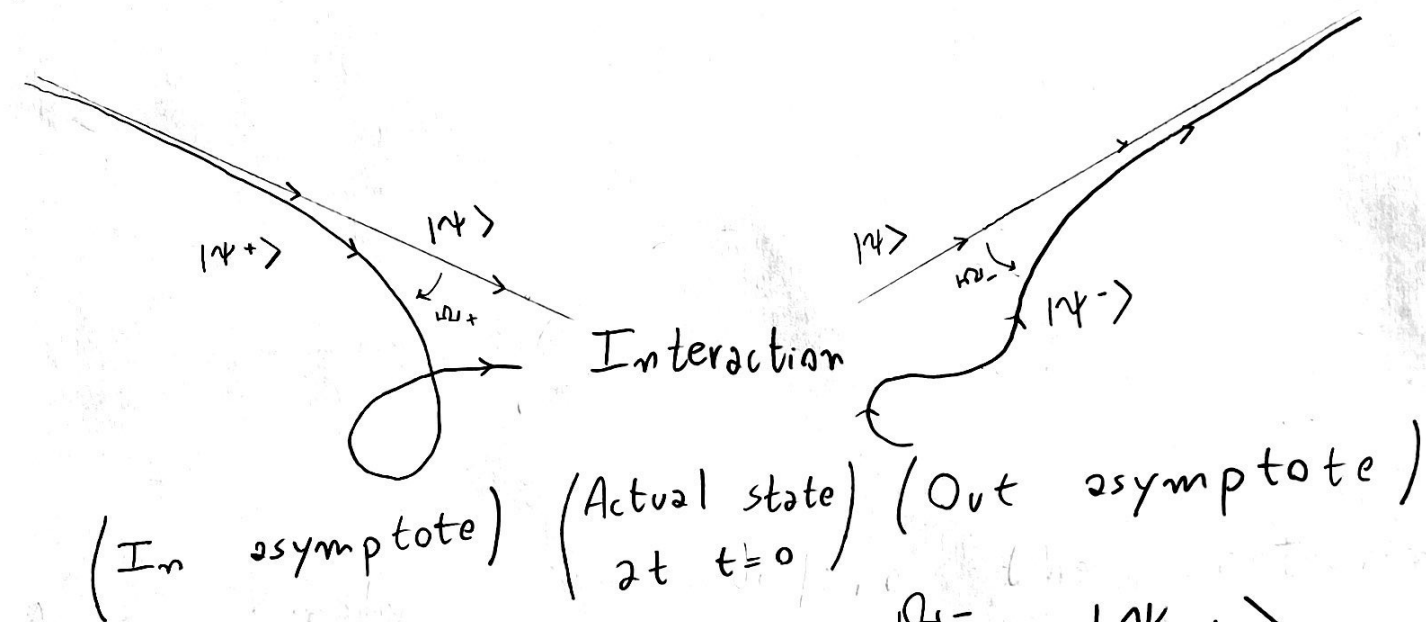
an asymptotic state into the interaction

region: $|\psi\rangle = \Omega_+ |\psi_{in}\rangle \equiv |\psi^+\rangle$. And

similarly $\Omega_- := \lim_{t \rightarrow +\infty} U^\dagger(t) U_0(t)$, that

gives the actual state $\Omega_- |\psi\rangle = |\psi^-\rangle$

at $t=0$ that would evolve to $|\psi\rangle$.



$$|\psi_{in}\rangle \xrightarrow{\Omega_+} |\psi\rangle \xleftarrow{\Omega_-} |\psi_{out}\rangle$$

$$\xrightarrow{\Omega_-^\dagger}$$

($\Omega_\pm$ are isometric). And finally we

define the scattering operator as $S |\psi_{in}\rangle = |\psi_{out}\rangle$, $S = \Omega_-^\dagger \Omega_+$, that's

unitary by construction.

If the interaction is, for example, spherically symmetrical and satisfy the usual assumptions:

- (i) As $r \rightarrow \infty$ $V(r)$ falls faster than r^{-3} .
- (ii) As $r \rightarrow 0$ $V(r)$ is less singular than r^{-2} .
- (iii) $V(r)$ is continuous except for a finite number of finite discontinuities.

then the limits defining $\Omega_{\pm}$ exist.

It's the momentum state $\Omega_{\pm} |\vec{k}\rangle = |\vec{k}^{\pm}\rangle$

defined this way that satisfy the

Lippmann-Schwinger equation

$$|\vec{k}^{\pm}\rangle = |\vec{k}\rangle + \frac{1}{E - H_0 \pm i\epsilon} V |\vec{k}^{\pm}\rangle, \quad (1.1)$$

and with the transition operator

$T |\vec{k}\rangle = V |\vec{k}^{\pm}\rangle$ we've showed that

$$\langle \vec{k}' | S | \vec{k} \rangle = \delta(\vec{k}' - \vec{k}) - 2\pi i \delta(E' - E) \langle \vec{k}' | T | \vec{k} \rangle \quad (1.2)$$

And the scattering amplitude

$$f(\vec{k}' | \vec{k}) = - (2\pi)^2 m \langle \vec{k}' | T | \vec{k} \rangle. \quad (1.3)$$

1.2 Partial-Wave Expansion

Assuming that the interaction is rotational invariant, S is a scalar operator.

To better exploit this fact we use the angular momentum base $|Elm\rangle$, so

the Wigner-Eckart theorem gives:

$$\langle E'l'm' | S | Elm \rangle = \delta(E'-E) \delta_{ll'} \delta_{m'm} S_l(E) \quad (1.4)$$

and since S is unitary $|S_l(E)| = 1$, so

we can define $S_l(E) = e^{2i\delta_l(E)}$ (this

is not always the case if other channels are open at that energy).

Since we can expand the state

$$\langle \vec{x} | Elm \rangle = \sqrt{\frac{2m\kappa}{\pi}} j_l(\kappa r) Y_{lm}(\hat{x}) \quad \text{and}$$

$$\langle \vec{x} | \vec{k} \rangle = \sqrt{\frac{2}{\pi}} \sum_{l,m} j_l(\kappa r) Y_{lm}(\hat{x}) Y_{lm}(\hat{k}), \quad \text{we have}$$

$$\langle \vec{k}' | Elm \rangle = \int d^3x \langle \vec{k}' | \vec{x} \rangle \langle \vec{x} | Elm \rangle =$$

$$= \int d^3x \left(\sqrt{\frac{2}{\pi}} \sum_{l',m'} j_{l'}^*(\kappa' r) Y_{l'm'}^*(\hat{x}) Y_{l'm'}(\hat{k}') \right) \times$$

$$\times \left(\sqrt{\frac{2m\kappa}{\pi}} j_l(\kappa r) Y_{lm}(\hat{x}) \right)$$

$$\left(\text{where } \kappa = \sqrt{2mE} \right)$$

$$= \frac{2}{\pi \kappa} \sqrt{m \kappa} \sum_{l, m'} Y_{l, m'}(\hat{r}) \underbrace{\int dr r^2 \dot{Y}_l(\kappa r) \dot{Y}_l(\kappa r)}_{\frac{\pi}{2\kappa^2} \delta(\kappa' - \kappa)} \underbrace{\int d\hat{x} Y_{l, m'}(\hat{x}) Y_{l, m}(\hat{x})}_{\delta_{l, m'} \delta_{m, m'}}$$

$$= \frac{\sqrt{m \kappa}}{\kappa^2} Y_{l, m}(\hat{r}) \delta(\kappa' - \kappa), \quad \text{let } E' = \sqrt{2m \kappa'},$$

Since $\delta(ax) = \frac{1}{|a|} \delta(x)$ and $\delta(x^2 - a^2) = \frac{\delta(x+a) + \delta(x-a)}{2|a|}$,

$$= 2 \sqrt{\frac{m}{\kappa}} Y_{l, m}(\hat{r}) \frac{1}{2m} \delta(E' - E) = \sqrt{m \kappa} Y_{l, m}(\hat{r}) \delta(E' - E),$$

so $\langle \vec{\kappa}' | E l m \rangle = \sqrt{m \kappa} Y_{l, m}(\hat{r}) \delta(E' - E) \quad (1.5)$

Combining (1.2), (1.3) and $\langle \vec{\kappa}' | \vec{\kappa} \rangle = \delta(\vec{\kappa}' - \vec{\kappa})$,

we see that $\langle \vec{\kappa}' | (S-1) | \vec{\kappa} \rangle = \frac{i}{2\pi m} \delta(E' - E) f(\vec{\kappa}' | \vec{\kappa})$,

but $\langle \vec{\kappa}' | (S-1) | \vec{\kappa} \rangle = \int dE \sum_{l, m} \langle \vec{\kappa}' | (S-1) | E l m \rangle \langle E l m | \vec{\kappa} \rangle$

with (1.4), (1.5) $= \frac{1}{m \kappa} \sum_{l, m} (S_l(E) - 1) Y_{l, m}(\hat{r}) Y_{l, m}(\hat{r}')$,

and with (1.2), (1.3) and the addition formula for spherical harmonics we obtain

$$f(\vec{\kappa}' | \vec{\kappa}) = \frac{1}{2\kappa i} \sum_{l \geq 0} (2l+1) (S_l(E) - 1) P_l(\cos \theta) \quad (1.6)$$

$$= \sum_{l \geq 0} (2l+1) f_l(E) P_l(\cos \theta)$$

where $\cos\theta = \hat{k} \cdot \hat{k}'$, and we've defined the partial wave amplitude as

$$f_l(E) = \frac{1 - S_l(E)}{2ki} = \frac{e^{2i\delta_l(E)} - 1}{2ki} = \frac{e^{i\delta_l(E)} \sin\delta_l(E)}{k} \quad (1.7)$$

2. Radial equation solutions

2.1 Free solutions

Seeking free solutions to a central potential with the form $\Psi(\vec{r}) = \sum_{l,m} c_{lm} y_l(r) Y_{lm}(\hat{r})$ will lead us to the spherical Bessel $j_l(kr)$ and Neumann $n_l(kr)$ functions or the spherical Hankel $h_l^{\pm}(kr) = n_l(kr) \pm i j_l(kr)$ functions. A different approach is to propose solutions with the form $\frac{y_l(r)}{r} Y_{lm}(\hat{r})$, leading us to the free radial equation

$$\left[\frac{d^2}{dr^2} - \frac{l(l+1)}{r^2} + p^2 \right] y(r) = 0 \quad (2.1)$$

We can now introduce the Riccati counterparts of the functions previously mentioned. A clear advantage of, say, the Riccati-Bessel function, is that it is entire.

Riccati-Bessel

Riccati-Neumann

Riccati-Hankel

Definition

$$\hat{j}_\ell = z j_\ell(z)$$

$$= z \sum_{n=0}^{\ell+1} \frac{(-z/2)^n}{(2\ell+2n+1)!! n!}$$

$$\hat{\eta}_\ell(z) = z \eta_\ell(z)$$

$$= z \sum_{n=0}^{\ell} \frac{(-z/2)^n (2\ell-2n-1)!!}{n!}$$

$$\hat{h}_\ell^\pm = \hat{\eta}_\ell \pm i \hat{j}_\ell$$

Behavior

$$\text{as } |z| \rightarrow 0 \quad \frac{z^{\ell+1}}{(2\ell+1)!!} (1+O(z^2))$$

$$z^{-\ell} (2\ell-1)!! (1+O(z^2)) \quad z^{-\ell} (2\ell-1)!! (1+O(z^2))$$

Behavior

$$\text{as } |z| \rightarrow \infty \quad \sin(z - \frac{1}{2}\ell\pi) (1+O(z^{-1}))$$

$$\cos(z - \frac{1}{2}\ell\pi) (1+O(z^{-1})) \quad e^{\pm i(z - \frac{1}{2}\ell\pi)} (1+O(z^{-1}))$$

Both $\{\hat{j}_\ell, \hat{\eta}_\ell\}$ and $\{\hat{h}_\ell^\pm\}$ are linearly independent sets of solutions to (2.1), that span the space of solutions.

It will be handy to establish some bounds to those functions. Both $\hat{j}_\ell$ and $\hat{\eta}_\ell$,

because of the trigonometric behaviour as $|z| \rightarrow \infty$, have a growth dominated by $e^{|\text{Im } z|}$.

As $|z| \rightarrow 0$, $|\hat{j}_\ell| \sim |z|^{\ell+1} \sim \left(\frac{|z|}{1+|z|}\right)^{\ell+1}$, and

since $\left(\frac{|z|}{1+|z|}\right)^{\ell+1} \sim 1$ as $|z| \rightarrow \infty$. Since $\hat{j}_\ell$ is continuous and hence bounded by

a constant in the finite interval between these extremes, we have:

$$|\hat{j}_\ell(z)| \leq C_j \left(\frac{|z|}{1+|z|}\right)^{\ell+1} e^{|\text{Im } z|} \quad (2.2)$$

By similar arguments,

$$|\hat{\eta}_\pm(z)| \leq C_n \left(\frac{|z|}{1+|z|} \right)^{-l} e^{|\operatorname{Im} z|} \quad (2.3)$$

The growth of the Riccati-Hankel function near the origin is dominated by the singular nature of $\eta_\pm \sim z^{-l}$. As $|z| \rightarrow \infty$,

since $|e^{\pm z}| = |e^{i \operatorname{Re} z}| |e^{\mp \operatorname{Im} z}| = e^{\mp \operatorname{Im} z}$, we have

$$|\hat{h}^\pm(z)| \sim e^{\pm \operatorname{Im} z}, \quad \text{so}$$

$$|\hat{h}^\pm(z)| \leq C_n \left(\frac{|z|}{1+|z|} \right)^{-l} e^{\mp \operatorname{Im} z} \quad (2.4)$$

2.2 Normalized solution

In class, we've explored stationary solutions of the form

$$\psi_\kappa(\vec{r}) \xrightarrow{r \rightarrow \infty} (2\pi)^{-3/2} \left(e^{i\vec{\kappa} \cdot \vec{r}} + \frac{e^{i\kappa r}}{r} f(\vec{\kappa}' | \vec{\kappa}) \right), \quad (2.5)$$

constructed as $\psi_\kappa(\vec{r}) = \sum_{\ell, m} \sqrt{\frac{2\ell+1}{2\pi^2}} i^\ell R_\ell(\kappa, r) Y_{\ell m}(\hat{r})$

where $r R_\ell(\kappa, r)$ solves

$$\left[\frac{d^2}{dr^2} - \frac{\ell(\ell+1)}{r^2} + p^2 - (\lambda) U(r) \right] y(r) = 0 \quad (2.6)$$

with boundary conditions defined by (2.5). Those requirements can be combined in an integral equation of the form

$$R_\ell(\kappa, r) = j_\ell(\kappa r) + \int_0^\infty dr' r'^2 G_\ell^\kappa(r, r') U(r') R_\ell(\kappa, r')$$

where $G_l^\kappa(r, r') = -i\kappa j_l(\kappa r) h_l^+(\kappa r')$ is the appropriate Green function, and $U(r) = 2mV(r)$. $\psi_{l,\kappa}(\vec{r})$ can also be expanded as $\sum_{l,m} \sqrt{\frac{2l+1}{2\pi^2}} i^l \frac{\psi_{l,\kappa}(r)}{r\kappa} Y_{lm}(\hat{r})$,

now with $\psi_{l,\kappa}(r) = \hat{j}_l(\kappa r) + \lambda \int_0^\infty dr' \hat{G}_l^\kappa(r, r') U(r') \psi_{l,\kappa}(r')$ (2.7)

and $\hat{G}_l^\kappa(r, r') = -\frac{i}{\kappa} \hat{j}_l(\kappa r) \hat{h}_l^+(\kappa r')$. I've introduced in (2.7) a strength parameter $\lambda V(r)$ to the potential, and notice that $\psi_{l,\kappa}$ solve

(2.7). In fact by construction we have

$$\langle \vec{r} | \Omega_+ | E l m \rangle = \text{const} \times \frac{\psi_{l,\kappa}(\vec{r})}{\kappa r} Y_{lm}(\hat{r}), \quad \text{and since}$$

$$\langle E' l m' | E l m \rangle = \langle E' l m' | \Omega_+^\dagger \Omega_+ | E l m \rangle = \langle E' l m' | E l m \rangle =$$

$$= \delta(E' - E) \delta_{l'l} \delta_{m'm}, \quad \langle \vec{r} | E l m \rangle = \text{const} \times \frac{\hat{j}_l(\kappa r)}{\kappa r} Y_{lm}(\hat{r}),$$

the normalization of $\psi_{l,\kappa}(\vec{r})$ mimic that of $\hat{j}_l(\kappa r)$. If we align $\hat{k}$ and $\hat{r}$,

$$\mathcal{H}_\kappa(\vec{r}) = \langle \vec{r} | \vec{\kappa}^+ \rangle = (2\pi)^{-3/2} \frac{1}{\kappa r} \sum_{l \geq 0} (2l+1) i^l \psi_{l,\kappa}(\vec{r}) P_l(\hat{r} \cdot \hat{\kappa}) \rightarrow$$

$$\xrightarrow{r \rightarrow \infty} (2\pi)^{-3/2} \left(e^{i\vec{\kappa} \cdot \vec{r}} + f(\kappa \hat{r} | \vec{\kappa}) \frac{e^{i\kappa r}}{r} \right) =$$

$$= (2\pi)^{-3/2} \frac{1}{\kappa r} \sum_{l \geq 0} (2l+1) (i^l \hat{j}_l(\kappa r) + \kappa f_l(\kappa) e^{i\kappa r}) P_l(\hat{r} \cdot \hat{\kappa})$$

using (1.6) and $e^{i\vec{\kappa} \cdot \vec{r}} = \sum_{l \geq 0} (2l+1) i^l j_l(\kappa r) P_l(\hat{r} \cdot \hat{\kappa})$

Since $i^2 = e^{2 \log i} = e^{i 2\pi/2}$, comparing we obtain

$$\psi_{\ell, \kappa}(r) \xrightarrow{r \rightarrow \infty} \hat{j}_{\ell}(\kappa r) + \kappa f_{\ell}(\kappa) e^{i(\kappa r - \frac{1}{2}\ell\pi)} \quad \text{or}$$

$$\longrightarrow \hat{j}_{\ell}(\kappa r) + \kappa f_{\ell}(\kappa) \hat{h}_{\ell}^{+}(\kappa r)$$

$$\longrightarrow \frac{i}{2} (\hat{h}_{\ell}^{-}(\kappa r) - s_{\ell}(\kappa) \hat{h}_{\ell}^{+}(\kappa r)), \quad \text{using}$$

the $r \rightarrow \infty$ form of the Riccati-Hankel functions and $\kappa f_{\ell} = (s_{\ell} - 1)/2i$, $\hat{j} = (\hat{h}^{+} - \hat{h}^{-})/2i$.

2.3 Regular solution

The initial conditions defining $\psi_{\ell, \kappa}(r)$

are $\psi_{\ell, \kappa}(0) = 0 = \text{const. } \hat{j}_{\ell}(\kappa r)$, $\psi'_{\ell, \kappa}(0) = \text{const. } \hat{j}'_{\ell}(\kappa r)$,

or $\psi_{\ell, \kappa}(r) \xrightarrow{r \rightarrow 0} \text{const. } \hat{j}_{\ell}(\kappa r)$. We now define a

new solution $\phi_{\ell, \kappa}(r)$ proportional to $\psi_{\ell, \kappa}(r)$, but

now with $\phi_{\ell, \kappa}(r) \xrightarrow{r \rightarrow 0} \hat{j}_{\ell}(\kappa r)$. The boundary

conditions are contained in the following

integral (Volterra) equation:

$$\phi_{\ell, \kappa}(r) = \hat{j}_{\ell}(\kappa r) + \lambda \int_0^r dr' g_{\ell}^{\kappa}(r, r') U(r') \phi_{\ell, \kappa}(r'), \quad (2.8)$$

where $g_{\ell}^{\kappa}(r, r') = \frac{1}{\kappa} (\hat{j}_{\ell}(\kappa r) \hat{\eta}_{\ell}(\kappa r') - \hat{\eta}_{\ell}(\kappa r) \hat{j}_{\ell}(\kappa r'))$.

Let $\frac{d}{dr} = \quad$, since $g' = \frac{1}{\kappa} (\hat{j}' \hat{\eta} - \hat{\eta}' \hat{j}) = \frac{1}{\kappa} \kappa = 1$

and $g'' = \frac{1}{\kappa} (\hat{j}'' \hat{\eta} - \hat{\eta}'' \hat{j}) = 0$, because both

$\hat{j}$ and $\hat{\eta}_l$ solve (2.1). Now (dropping the indexes l, κ) since $g(r, r) = 0$,

$$\phi''(r) = \hat{j}''(r) + \lambda \int_0^r dr' g''(r, r') U(r') \phi(r') + g(r, r) U(r) \phi(r) + g'(r, r) U(r) \phi(r)$$

Again, $\hat{j}, \hat{\eta}$ solve (2.1) so $\hat{j}'' = \left(\frac{l(l+1)}{r^2} - \kappa^2 \right) \hat{j}$

etc, so $g''(r, r') = \left(\frac{l(l+1)}{r^2} - \kappa^2 \right) g(r, r')$, hence

$$\phi''(r) = \left(\frac{l(l+1)}{r^2} - \kappa^2 \right) \hat{j} + \lambda \left(\frac{l(l+1)}{r^2} - \kappa^2 \right) \int_0^r dr' g(r, r') U(r') \phi(r') +$$

$$+ U(r) \phi(r), \quad \text{since } \phi(r) = \hat{j}(r) + \lambda \int_0^r dr' g(r, r') U(r') \phi(r'),$$

$$\phi''(r) = \left(\frac{l(l+1)}{r^2} - \kappa^2 + U(r) \right) \phi(r), \quad \text{so in}$$

fact a solution to (2.8) is a solution to (2.6) with $\phi_{l, \kappa}(r) \xrightarrow{r \rightarrow 0} \hat{j}_l(\kappa r)$, as is clear by the Volterra form of the equation.

Theorem 2.1: (2.8) can be solved by iteration with any complex κ, λ , and

the solution $\phi_{l, \kappa, \lambda}(r) = \sum_{n=0}^{\infty} \lambda^n \phi_{l, \kappa}^{(n)}(r)$ with

$$\phi_{l, \kappa}^{(0)}(r) = \hat{j}_l(\kappa r), \quad \phi_{l, \kappa}^{(n+1)}(r) = \int_0^r dr' g_{\kappa}^l(r, r') U(r') \phi_{l, \kappa}^{(n)}(r).$$

is an entire function of κ and l .

Proof: If every $\phi^{(n)}$ is bounded, since $\hat{J}_l(\kappa r)$ is entire in κ , so will be

$$\phi^{(n+1)}(r) = \int_0^r dr' g(r, r') U(r') \phi^{(n)}(r').$$
 Using the

bounds of (2.2) and (2.3) we conclude that $|g(r, r')| \leq \frac{C_g}{|\kappa|} \left(\frac{|\kappa| r_2}{1 + |\kappa| r_2} \right)^{l+1} \left(\frac{|\kappa| r_2}{1 + |\kappa| r_2} \right)^{-l} e^{|Im \kappa| (r_2 - r_2)}$.

Since $\phi^{(n)}(r) = \int_0^r dr_n \int_0^{r_n} dr_{n-1} \dots \int_0^{r_2} dr_1 g(r, r_n) U(r_n) g(r_n, r_{n-1}) \dots U(r_1) \hat{J}_l(\kappa r_1)$,

and $r \geq r_n \geq r_{n-1} \geq \dots \geq r_2 \geq r_1$, using our bounds

$$|\phi^{(n)}(r)| \leq C_g^n \left(\frac{|\kappa| r}{1 + |\kappa| r} \right)^{l+1} e^{|Im \kappa| r} \frac{1}{|\kappa|^n} \int_0^r dr_n \int_0^{r_n} dr_{n-1} \dots \int_0^{r_2} dr_1 \left(\frac{|\kappa| r_n}{1 + |\kappa| r_n} \right)^{-l} e^{-|Im \kappa| r_n}$$

$$|U(r_n)| \left(\frac{|\kappa| r_n}{1 + |\kappa| r_n} \right)^{l+1} \left(\frac{|\kappa| r_{n-1}}{1 + |\kappa| r_{n-1}} \right)^{-l} e^{|Im \kappa| (r_n - r_{n-1})} |U(r_{n-1})| \times \quad (*)$$

$$\times \left(\frac{|\kappa| r_2}{1 + |\kappa| r_2} \right)^{l+1} \left(\frac{|\kappa| r_1}{1 + |\kappa| r_1} \right)^{-l} e^{|Im \kappa| (r_2 - r_1)} |U(r_1)| C_g \left(\frac{|\kappa| r_1}{1 + |\kappa| r_1} \right)^{l+1} e^{|Im \kappa| r_1}$$

$$= C_g^n C_g \left(\frac{|\kappa| r}{1 + |\kappa| r} \right)^{l+1} e^{|Im \kappa| r} \frac{1}{|\kappa|^n} \int_0^r dr_n \int_0^{r_n} dr_{n-1} \dots \int_0^{r_2} dr_1 \prod_{i=1}^n \left(\frac{|\kappa| r_i}{1 + |\kappa| r_i} \right) |U(r_i)|$$

The integral is invariant under the exchange $r_i \leftrightarrow r_j$, and it's performed in the "triangular" region $r \geq r_n \geq \dots \geq r_1$, that is $1/n!$ of the volume of $0 \leq r_1, \dots, r_n \leq r$, where the integrand is certainly bounded.

so we have

$$\int_0^r dr_n \int_0^{r_n} dr_{n-1} \cdots \int_0^{r_2} dr_1 \prod_{i=1}^n \frac{r_i}{1+|\kappa|r_i} |U(r_i)| \leq \int_0^r dr_n \int_0^{r_n} dr_{n-1} \cdots \int_0^{r_2} dr_1 \prod_{i=1}^n r_i |U(r_i)|$$

$$\leq c' \frac{1}{n!} \left(\int_0^r dr' r' |U(r')| \right)^n = c' \frac{\alpha^n}{n!}, \quad \text{since } V(r) = O(r^{-2+\varepsilon})$$

as $r \rightarrow 0$ the integral is convergent.

Now $|\phi^{(n)}(r)| \leq \left(\frac{|\kappa|r}{1+|\kappa|r} \right)^{l+1} e^{|\operatorname{Im} \kappa| r} c' \frac{(C_5 \alpha)^n}{n!}, \quad \text{and}$

$$|\phi(r)| = \left| \sum_{n \geq 0} \lambda^n \phi^{(n)}(r) \right| \leq \sum_{n \geq 0} |\lambda^n| |\phi^{(n)}(r)| \leq$$

$$\leq \left(\frac{|\kappa|r}{1+|\kappa|r} \right)^{l+1} e^{|\operatorname{Im} \kappa| r} c' \sum_{n \geq 0} \frac{(\lambda C_5 \alpha)^n}{n!} =$$

$$= \gamma \left(\frac{|\kappa|r}{1+|\kappa|r} \right)^{l+1} e^{|\operatorname{Im} \kappa| r}, \quad \text{where } \gamma = c' e^{\lambda C_5 \alpha}.$$

We conclude that $\phi_{\ell, \kappa}^{(n)}(r)$ is entire on

κ , and that $\phi_{\kappa, \ell, \lambda}(r) = \sum_{n \geq 0} \lambda^n \phi_{\ell, \kappa}^{(n)}(r)$ is

an absolutely convergent series of analytic

function, so $\phi_{\kappa, \ell, \lambda}(r)$ is entire on

κ and λ , with $|\phi_{\kappa, \ell}(r)| \leq \gamma_\ell \left(\frac{|\kappa|r}{1+|\kappa|r} \right)^{l+1} e^{|\operatorname{Im} \kappa| r}.$

Now we show that, in fact, $\phi_{\ell, \kappa}$

is a solution to (2.8). Since the

series is absolutely convergent, the order

of sum and integration can be exchanged: Since $\phi_{\lambda, \kappa}^{(0)}(r) = \hat{\gamma}_\kappa(\kappa r)$,

$$\begin{aligned}\phi(r) &= \hat{\gamma}(\kappa r) + \lambda \int_0^r dr' g(r, r') U(r') \phi(r') \\ &= \phi^{(0)}(r) + \sum_{n \geq 0} \lambda^{n+1} \underbrace{\int_0^r dr' g(r, r') U(r') \phi^{(n)}(r')}_{\phi^{(n+1)}(r)}, \text{ we}\end{aligned}$$

see the definition of $\phi^{(n+1)}$, hence

$$\phi(r) = \phi^{(0)}(r) + \sum_{n \geq 0} \lambda^{n+1} \phi^{(n+1)}(r) = \sum_{n \geq 0} \lambda^n \phi^{(n)}(r), \text{ as}$$

expected. ■

We highlight that the only physical λ are those on the real line, and the only physical κ are those on the positive real line.

With the definitions of $\hat{\gamma}$, $\hat{\eta}$ and $\hat{h}^\pm$,

$$\begin{aligned}\text{we conclude that } \hat{\gamma}_\kappa(-z) &= (-1)^{l+1} \hat{\gamma}_\kappa(z), \\ \hat{\eta}_\kappa(-z) &= (-1)^l \hat{\eta}_\kappa(z) \text{ so } g_\kappa^{-\kappa}(r, r') = g_\kappa^\kappa(r, r').\end{aligned}$$

So by definition we have $\phi_{\lambda, -\kappa}(r) = (-1)^{l+1} \phi_{\lambda, \kappa}(r)$.

If we consider physical λ and κ , since $\phi_{\lambda, \kappa}(r) \rightarrow \hat{\gamma}_\kappa(\kappa r)$ is real and (2.6) is a real equation, so is $\phi_{\lambda, \kappa}(r)$.

2.4 Jost function

Since $\phi_{\ell, \kappa}(r)$ and $\psi_{\ell, \kappa}(r)$ satisfy proportional initial conditions, they are proportional.

Let $\phi_{\ell, \kappa}(r) = \omega_{\ell}(\kappa) \psi_{\ell, \kappa}(r)$, where we define the Jost function $\omega_{\ell}(\kappa)$. When we consider the strength parameter of the potential, we may write $\phi_{\ell, \kappa, \lambda}(r) = \omega_{\ell}(\kappa, \lambda) \psi_{\ell, \kappa, \lambda}(r)$.

Since $\phi_{\ell, \kappa}(r)$ is real and as $r \rightarrow \infty$ it becomes a solution to

the free equation (2.1), and $\hat{h}^{\pm*} = \hat{h}^{\mp}$ span the solutions of (2.1),

$$\begin{aligned} \phi_{\ell, \kappa}(r) &\xrightarrow{r \rightarrow \infty} \frac{i}{2} (c_{\ell, \kappa} \hat{h}_{\ell}^{-}(\kappa r) - c_{\ell, \kappa}^{*} \hat{h}_{\ell}^{+}(\kappa r)) \quad (\text{real}) \\ &\longrightarrow \omega_{\ell}(\kappa) \frac{i}{2} (\hat{h}_{\ell}^{-}(\kappa r) - S_{\ell}(p) \hat{h}_{\ell}^{+}(\kappa r)) \end{aligned}$$

(asymptotic form of $\psi_{\ell, \kappa}(r)$), we conclude

that $c_{\ell, p} = \omega_{\ell}(\kappa)$, $\phi_{\ell, \kappa}(r) \xrightarrow{r \rightarrow \infty} \frac{i}{2} (\omega_{\ell}(\kappa) \hat{h}_{\ell}^{+}(\kappa r) - \omega_{\ell}(\kappa)^{*} \hat{h}_{\ell}^{-}(\kappa r))$,

so $S_{\ell}(p) = \frac{\omega_{\ell}(\kappa)^{*}}{\omega_{\ell}(\kappa)} = e^{2i\delta_{\ell}(\kappa)}$, so we

have $\omega_{\ell}(\kappa) = |\omega_{\ell}(\kappa)| e^{-i\delta_{\ell}(\kappa)}$. Since

by definition $\hat{\gamma}_{\ell} = \frac{i}{2} (\hat{h}_{\ell}^{-} - \hat{h}_{\ell}^{+})$ and

$\hat{\eta}_{\ell} = \frac{1}{2} (\hat{h}_{\ell}^{-} + \hat{h}_{\ell}^{+})$, we can write

the Green function of (2.8) as

$$g_{e,\kappa}(r, r') = \frac{1}{2\kappa} \left(\hat{h}_e^-(\kappa r) \hat{h}_e^+(\kappa r') - \hat{h}_e^+(\kappa r) \hat{h}_e^-(\kappa r') \right),$$

$$\phi_{e,\kappa}(r) = \hat{g}_e(\kappa r) + \lambda \int_0^r dr' g_{e,\kappa}(r, r') U(r') \phi_{e,\kappa}(r') =$$

$$= \frac{i}{2} \left[\left(1 + \frac{\lambda}{\kappa} \int_0^r dr' \hat{h}_e^+(\kappa r') U(r') \phi_{e,\kappa}(r') \right) \hat{h}_e^-(\kappa r) - (\dots)^* \hat{h}_e^+(\kappa r) \right]$$

$$\xrightarrow{r \rightarrow \infty} \frac{i}{2} \left(\omega_e(\kappa) \hat{h}_e^-(\kappa r) - \omega_e(\kappa)^* \hat{h}_e^+(\kappa r) \right), \quad \text{so we}$$

conclude that

$$\omega_e(\kappa, \lambda) = 1 + \frac{\lambda}{\kappa} \int_0^\infty dr \hat{h}_e^+(\kappa r) U(r) \phi_{e,\kappa,\lambda}(r) \quad (2.9).$$

It will be handy to establish a bound to $\omega - 1$. Using the bounds already established for $\hat{h}^+$ and ϕ , we have:

$$\begin{aligned} |\omega_e(\kappa, \lambda) - 1| &\leq \text{const.} \left| \frac{\lambda}{\kappa} \right| \int_0^\infty dr \left(\frac{|\kappa| r}{1 + |\kappa| r} \right)^{-\ell} \left(\frac{|\kappa| r}{1 + |\kappa| r} \right)^{\ell+1} |U(r)| e^{(\text{Im} \kappa - \text{Im} \kappa) r} \\ &\leq \text{const.} |\lambda| \int_0^\infty dr r |U(r)| e^{(\text{Im} \kappa - \text{Im} \kappa) r} \quad (2.10), \end{aligned}$$

so without further assumptions on V we is bounded on $\text{Im} \kappa \geq 0$ and analytic on $\text{Im} \kappa > 0$ (open region), since the exponential term diverge otherwise.

2.5 Hankel-like solutions.

We need to introduce yet another pair of solutions to (2.6). We now seek solutions of the form $\chi_{\ell, \kappa}^{\pm}(r)$, with

$\chi_{\ell, \kappa}^{\pm}(r) \xrightarrow{r \rightarrow \infty} \hat{h}_{\ell}^{\pm}(\kappa r)$. Luckily, we can use the same Green function $g_{\ell}^{\kappa}(r, r')$, but now with

$$\chi_{\ell, \kappa}^{\pm}(r) = \hat{h}_{\ell}^{\pm}(\kappa r) - \lambda \int_r^{\infty} dr' g_{\ell}^{\kappa}(r, r') U(r') \chi_{\ell, \kappa}^{\pm}(r'). \quad (2.11)$$

It's clear that $\chi^{\pm} \rightarrow \hat{h}^{\pm}$ as $r \rightarrow \infty$, and this definition solves (2.6) by the same arguments given to ϕ . And just like ϕ , we have

Theorem 2.2 Without further assumption on the potential, (2.11) can be solved by iteration with any complex λ and with $\kappa \in \{\text{Im } \kappa \geq 0\}$ in case of χ^{+} or $\kappa \in \{\text{Im } \kappa \leq 0\}$ in case of χ^{-} . The

solution $\chi^{\pm} = \sum_{n \geq 0} \lambda^n \chi^{\pm(n)}$ ($\chi^{\pm(0)}(r) = \hat{h}_{\ell}^{\pm}(\kappa r)$)

$\chi_{\ell, \kappa, \lambda}^{\pm(n+1)} = - \int_r^{\infty} dr' g_{\ell}^{\kappa}(r, r') U(r') \chi_{\ell, \kappa, \lambda}^{\pm(n)}(r')$ is entire

in λ , analytic for $\kappa \in \{\text{Im } \kappa > (<) 0\}$ and

continuous for $k \in \{ \operatorname{Im} k \geq (\leq) 0 \}$.

"Proof": The proof is analogous to that of theorem 2.1. The only difference is that in the cancelations of (*) we have to restrict ourselves to the upper (lower) half plane of k , since $|\hat{h}_\ell^\pm(kr)| \leq C_{h^\pm} \left(\frac{|k|r}{1+|k|r} \right)^{-\ell} e^{\pm \operatorname{Im} k r}$.

2.6 The Wronskian

We introduce the Wronskian of a pair of functions $f(r)$ and $g(r)$ as the

function
$$W(f, g)(r) = \det \begin{bmatrix} f(r) & g(r) \\ f'(r) & g'(r) \end{bmatrix} = f(r)g'(r) - f'(r)g(r).$$

Claim: if f, g are linearly dependent, $W(f, g) = 0$.

We see that if $f = cg$ for some $c \in \mathbb{C}$,
$$W(f, g) = fg' - f'g = cgg' - cg'g = 0.$$

Claim: if f, g solve $\left(\frac{d^2}{dr^2} + F(r) \right) y(r) = 0$,

then $W(f, g)$ is independent of r .

We see that
$$W(f, g)' = f'g' + fg'' - f''g - f'g' = fg'' - f''g = -Ffg + Ffg = 0,$$
 so $W(f, g)$ is independent of r .

It's convenient now to calculate the Wronskian of some solutions to our radial equations (2.1) and (2.6): When we're dealing with solutions to the same equation, we can calculate the Wronskian in some limit since it will be independent of r .

$$W(\hat{\eta}_\ell, \hat{\eta}_\ell) \stackrel{r \rightarrow \infty}{=} W(\sin(\kappa r - \frac{1}{2} \ell \pi), -\cos(\kappa r - \frac{1}{2} \ell \pi)) = -\kappa$$

$$W(\chi_{\ell, \kappa}^+, \chi_{\ell, \kappa}^-) \stackrel{r \rightarrow \infty}{=} W(\hat{h}_\ell^+, \hat{h}_\ell^-) \stackrel{r \rightarrow \infty}{=} W(e^{i(\kappa r - \frac{1}{2} \ell \pi)}, e^{-i(\kappa r - \frac{1}{2} \ell \pi)}) = -2i\kappa$$

$$W(\chi_{\ell, \kappa}^+, \phi_{\ell, \kappa}) \stackrel{r \rightarrow \infty}{=} W(\hat{h}_\ell^+, \frac{i}{2}(\omega_\ell(\kappa)\hat{h}_\ell^- - \omega_\ell(\kappa)^*\hat{h}_\ell^+)) = \kappa \omega_\ell(\kappa),$$

since $W(\hat{h}_\ell^+, \hat{h}_\ell^+) = 0$. We arrived at the important result

$$\omega_\ell(\kappa) = \frac{1}{\kappa} W(\chi_{\ell, \kappa}^+, \phi_{\ell, \kappa}). \quad (2.12)$$

3. Analytic properties of the S-matrix function and of the S-matrix elements

3.1 Continuation of $\omega_\ell(\kappa)$ to the lower half plane, "reasonable" potentials.

The analytic properties of depend on the convergence of the integral (2.10),

$$|\omega_2(k) - 1| \leq \text{const.} \int_0^\infty dr \, r |U(r)| e^{(|\text{Im } k| - \text{Im } k)r},$$

that certainly convergent for $k \in \text{Im } k$. But with further assumptions on $V(r)$, we can make $\omega_2(k)$ analytic in larger regions: We consider the "reasonable" potentials:

(i) $V(r) = 0$ for $r \geq a$, finite range potential:

In this case, the integral of (2.10) is certainly convergent for all k because

it runs from 0 to a (if the other assumptions on V still hold). Then $\omega_2(k)$ is entire on k .

(ii) $V(r) \sim e^{-\nu r}$ with $\nu > 0$.

In this case, on the lower half plane the exponential part of the integral

(2.10) will be $e^{-2\text{Im } k r}$, combining

with the potential we obtain $e^{-(\nu + 2\text{Im } k)r}$,

so the integral converges in the region

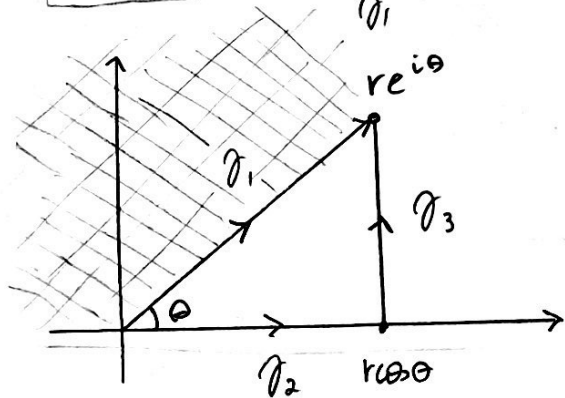
$\text{Im } k \geq -\nu/2$ of the lower half plane,

and $\omega_2(k)$ is analytic for $\text{Im } k > -\nu/2$.

(iii) $V(r)$ is analytic on $r > 0$, and satisfy the usual assumptions on any ray $\{re^{i\theta}; 0 \leq r < \infty\}$ on the half plane $\operatorname{Re} re^{i\theta} > 0$: (A Yukawa potential is an example of this, since in the half plane $\operatorname{Re} re^{i\theta} < 0$ it's divergent).

In this case we can see that since for a contour integral $|\int_{\gamma} f(z) dz| \leq \max_{z \in \gamma} |f(z)| L_{\gamma}$ where γ is a finite path of length L_{γ} , and $\int_{\gamma} dz \hat{h}^+ U \phi$ is path independent,

$$\int_{[0, re^{i\theta}]} dz \hat{h}^+(\kappa z) U(z) \phi(z) = \underbrace{\int_0^r dr \hat{h}^+(\kappa r) U(r) \phi(r)}_{\gamma_2} + \underbrace{\int_{[r e^{i\theta}, re^{i\theta}]} dz \hat{h}^+(\kappa z) U(z) \phi(z)}_{\gamma_3}$$



$$\text{as } r \rightarrow \infty, |U(z)| \sim |z|^{-3-\epsilon} = r^{-3-\epsilon},$$

$$\left| \int_{[r e^{i0}, re^{i\theta}]} dz \hat{h}^+(\kappa z) U(z) \phi(z) \right| \leq \left(\frac{|\kappa| r}{1 + |\kappa| r} \right)^2 e^{\operatorname{Im} \kappa r e^{i\theta}} \frac{1}{r^{3+\epsilon}}$$

$$\left(\frac{|\kappa| r}{1 + |\kappa| r} \right)^{2+\epsilon} e^{|\operatorname{Im} \kappa r e^{i\theta}|} \xrightarrow{r \rightarrow \infty} 0 \quad \text{if } \operatorname{Im}(\kappa e^{i\theta}) > 0,$$

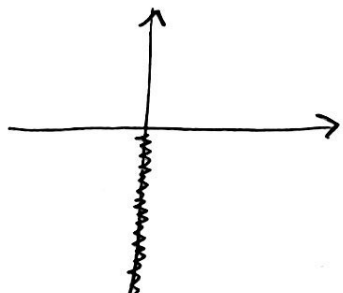
the shaded region, and we have

$$\int_0^{\infty e^{i\theta}} dr \hat{h}^+(\kappa r) U(r) \phi(r) = \int_0^{\infty} dr \hat{h}^+(\kappa r) U(r) \phi(r), \quad \text{and}$$

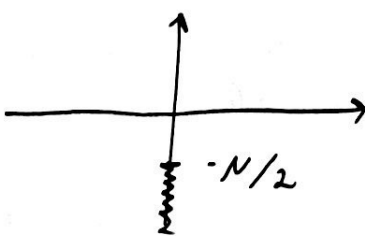
these are analytic functions of κ on $\operatorname{Im}(\kappa e^{i\theta}) > 0$.

Since θ can take any value from $-\pi$ to π , $\omega_e(k)$ is analytic in this case for $k \in \mathbb{C} \setminus \{bi; b < 0\}$:

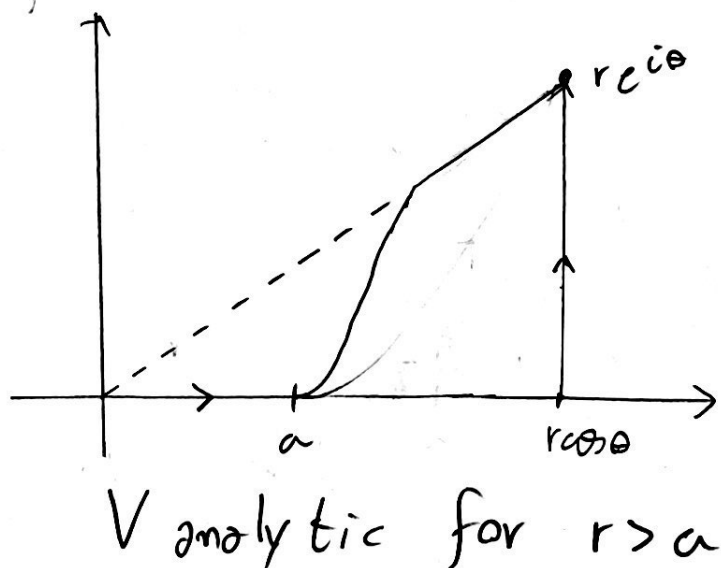
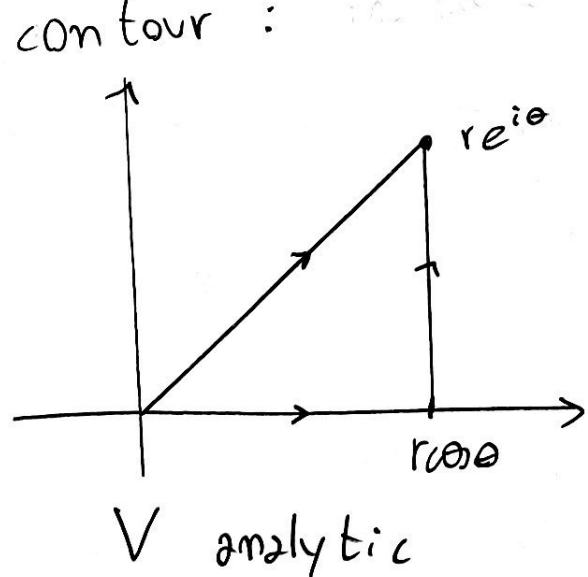
$V(r)$ analytic $\Rightarrow \omega_e(k)$ analytic on



If $V(r) \sim e^{-Nr}$ also, we combine the two conclusions to state that $\omega_e(k)$ is analytic for $k \in \mathbb{C} \setminus \{bi; b < -N/2\}$



We can draw the same conclusion if $V(r)$ is analytic for $r > a$, all we have to do is use a different contour:



It's convenient to highlight that there isn't much physical meaning to the obstacles that prevent us from continuing $\omega_\ell(\kappa)$ to the lower half plane: if, for instance, $V(r) = 0$ for $r > 10 \text{ km}$ the continuation is possible, but in practice a scattering experiment has now way to tell if $V(r) \neq 0$ for r that large. We shall call a potential that allows the continuation of $\omega_\ell(\kappa)$ to the lower half plane (possibly with the exclusion of some region where $\text{Im } \kappa < -V/2$) reasonable potentials.

3.2 S matrix elements as meromorphic functions.

We've seen in section 2.3 that $\phi_{\ell, -\kappa}(r) = (-1)^{\ell+1} \phi_{\ell, \kappa}(r)$, and since $\hat{h}_\ell^+(-\kappa r) = \hat{h}_\ell^{+*}(\kappa r) = (-1)^\ell \hat{h}_\ell^+(\kappa r)$, from 2.9 we have

$$\omega_\ell(-\kappa) = 1 - \frac{1}{\kappa} \int_0^\infty dr \hat{h}_\ell^+(-\kappa r) U(r) \phi_{\ell, \kappa}(r)$$

$$= 1^* + \frac{1}{\kappa^*} \int_0^\infty dr \hat{h}_\ell^{+*}(\kappa r) U(r)^* \phi_{\ell, \kappa}(r)^* = \omega_\ell(\kappa^*)^*$$

for physical $\kappa = \kappa^*$. By the Schwartz reflection principle, if an analytic function $f: G \rightarrow \mathbb{C}$ can be continued to G^* and $f(z) = f(z^*)^*$ in some set with a cluster point on G , then $f(z) = f(z^*)^*$ $\forall z \in G$. Since on the real line we

have $\omega_e(\kappa) = \omega_e(-\kappa^*)^*$. Applying Schwartz

principle for $\kappa = iz$, $\omega_e(z) = \omega_e(z^*)^*$,

so we conclude that $\omega_e(\kappa) = \omega_e(-\kappa^*)^*$

$\forall \kappa$ in the region of analyticity of ω_e .

For a reasonable potential, we see that

since for physical κ we have that

$$S_e(\kappa) = \frac{\omega_e(\kappa)^*}{\omega_e(\kappa)} = \frac{\omega_e(\kappa^*)^*}{\omega_e(\kappa)} = \frac{\omega_e(-\kappa)}{\omega_e(\kappa)}, \quad (3.1)$$

$S_e(\kappa)$ is meromorphic in some region

around the real axis (if, for

example, $\omega_e(\kappa)$ is analytic for $\text{Im} \kappa > -N/2$,

$S_e(\kappa)$ is meromorphic on the strip

$-N/2 < \text{Im} \kappa < N/2$).

3.3 Bound states and zeros of ω_ℓ / poles of S_ℓ .

Suppose that $\omega_\ell(k_0) = 0$ for some k_0 with $\text{Im} k_0 > 0$. If the potential is reasonable, this corresponds to a pole of $S_\ell(k) = \frac{\omega_\ell(k)}{\omega_\ell(k)}$ at $k = k_0$.

According to (2.12), $k\omega_\ell(k) = W(\chi_{\ell,k}^+, \phi_{\ell,k})$,

so $W(\chi_{\ell,k_0}^+, \phi_{\ell,k_0}) = k_0\omega_\ell(k_0) = 0$, so

using our discussion about the Wronskian on section 2.6 we can state that

$\chi_{\ell,k_0}^+(r)$ and $\phi_{\ell,k_0}(r)$ are proportional.

Since $\chi_{\ell,k}^+(r) \xrightarrow{r \rightarrow \infty} \hat{h}_\ell^+(\kappa r)$, we have

$$|\phi_{\ell,k_0}(r)| = |c \chi_{\ell,k_0}^+(r)| \xrightarrow{r \rightarrow \infty} |c \hat{h}_\ell^+(\kappa_0 r)| \leq$$

$$\leq \text{const.} \cdot e^{-\text{Im} k_0 r} \quad \text{according to (2.14), and}$$

since $\text{Im} k_0 > 0$ and $\phi_{\ell,k_0}(r) \xrightarrow{r \rightarrow 0} j_\ell(\kappa_0 r)$,

$\frac{\phi_{\ell,k_0}(r)}{r k_0} Y_{\ell m}(\hat{r})$ is a proper (normalizable)

eigenfunction of the Hamiltonian

with eigenvalue $E = \frac{k_0^2}{2m}$. Since

H is hermitian, $\frac{k_0^2}{2m} \in \mathbb{R}$, but by

assumption $\text{Im } \kappa_0 > 0$, so $\kappa_0 = i\alpha$ and

$$E = \frac{\hbar^2 \kappa_0^2}{2m} = -\frac{\alpha^2 \hbar^2}{2m}, \quad \text{we've constructed a}$$

normalizable, localized state with negative energy: a bound state. On the

other hand, if the Hamiltonian admits a bound state of energy $-\frac{\alpha^2 \hbar^2}{2m}$, the

last function has to vanish at $i\alpha$, because $\phi_{i\alpha, \ell}$ has to be normalizable.

So we've just proved

Theorem 3.1 : There is a one-to-one correspondence between zeros of $\omega_\ell(\kappa_0) = 0$ in the positive imaginary axis and bound states of energy $E = -\frac{\hbar^2 \kappa_0^2}{2m}$. If the potential is reasonable, the correspondence also holds for poles of S_ℓ on the positive imaginary axis, since if $\omega_\ell(\kappa_0) = 0$, $\omega_\ell(-\kappa_0) \neq 0$, otherwise we would have $\phi_{\ell, \kappa_0}(r) = \frac{i}{2} (\omega_\ell(\kappa_0) \chi_{\kappa_0, \ell}^-(r) - \omega_\ell(-\kappa_0) \chi_{\kappa_0, \ell}^+(r)) = 0$.

3.4 Levinson's theorem

In order to prove Levinson's theorem, we first need to prove 4 facts:

Lemma 3.2: The zeros of ω_ℓ at bound states are always simple:

If $\omega_\ell(\kappa_0) = 0$ at the positive imaginary axis, $\phi_{\kappa_0, \ell} = c \chi_{\kappa_0, \ell}^+$, since $\omega_\ell(\kappa) = \frac{1}{\kappa} W(\chi_{\kappa_0, \ell}^+, \phi_{\kappa_0, \ell})$.

Let $\frac{d}{d\kappa} = \dot{}$, $\dot{\omega} = \frac{1}{\kappa} (W(\dot{\chi}^+, \dot{\phi}) + W(\chi^+, \dot{\phi}))$. Since

both ϕ and χ^+ solve (2.6), we have:

$$\begin{aligned} W'(\chi_{\kappa_0}^+, \phi_{\kappa_0}) &= \chi_{\kappa_0}^+ \phi_{\kappa_0}'' - \chi_{\kappa_0}^{+''} \phi_{\kappa_0} = \\ &= \left(\frac{\ell(\ell+1)}{r^2} + U(r) - \kappa^2 \right) \chi_{\kappa_0}^+ \phi_{\kappa_0} - \left(\frac{\ell(\ell+1)}{r^2} + U(r) - \kappa_0^2 \right) \chi_{\kappa_0}^+ \phi_{\kappa_0} \\ &= (\kappa_0^2 - \kappa^2) \chi_{\kappa_0}^+ \phi_{\kappa_0}, \text{ so } W(\chi_{\kappa_0}^+, \phi_{\kappa_0}) = (\kappa_0^2 - \kappa^2) \int_0^r dr' \chi_{\kappa_0}^+ \phi_{\kappa_0}, \end{aligned}$$

$$\text{and } W(\chi_{\kappa_0}^+, \dot{\phi}_{\kappa_0}) = -2\kappa \int_0^r dr' \chi_{\kappa_0}^+ \dot{\phi}_{\kappa_0},$$

$$W(\chi_{\kappa}^+, \dot{\phi}_{\kappa}) = -2\kappa \int_0^r dr' \chi_{\kappa}^+ \dot{\phi}_{\kappa}. \text{ Similarly, we}$$

$$\text{have } W(\dot{\chi}_{\kappa}^+, \phi_{\kappa}) = -2\kappa \int_r^\infty dr' \dot{\chi}_{\kappa}^+ \phi_{\kappa}, \text{ so}$$

$$\dot{\omega}(\kappa_0) = \frac{1}{\kappa_0} \left(-2\kappa_0 \int_0^r dr' \chi_{\kappa_0}^+ \dot{\phi}_{\kappa_0} - 2\kappa_0 \int_r^\infty dr' \dot{\chi}_{\kappa_0}^+ \phi_{\kappa_0} \right)$$

$$= -2c \int_0^r dr' (\chi_{\kappa_0}^+)^2. \quad \chi_{\kappa_0} = i^{\ell} \text{ (real function)}$$

because κ_0 is pure imaginary, $\dot{\omega}(\kappa_0) \neq 0$,

the zero $\omega(k_0) = 0$ is of order 1.

Lemma 3.3: If $V(r) = O(r^{-2+\eta})$ as $r \rightarrow 0$ (for $0 < \eta < 1$), then $|\omega_e(k) - 1| = O(|k|^{-\eta})$ as $|k| \rightarrow \infty$ on the upper half plane:
From (2.10), $|\omega_e(k) - 1| \leq \frac{\text{const.}}{|k|} \int_0^\infty dr \frac{|k|r}{1+|k|r} |U(r)| e^{(1+\eta|k|-1)r}$
 $\leq \text{const.} \left(\int_0^1 dr \frac{r}{1+|k|r} |U(r)| + c' \int_1^\infty dr \frac{r}{1+|k|r} \frac{1}{r^{2-\eta}} \right)$
 $\rightarrow O(|k|^{-1}) + O(|k|^{-\eta}) = O(|k|^{-\eta}).$

Lemma 3.4: If ω_e vanishes at the origin, it does so linearly or quadratically:

For a reasonable potential, ψ_e is analytic at 0, so it assumes a power series expansion $\omega_e(k) = \sum_{n \geq 0} c_n k^n$. We have

$$\omega_e = 1 + \frac{1}{k} \int \hat{h}^+ U \phi = 1 + \frac{1}{p} \int \hat{\eta} U \phi + \frac{i}{p} \int \hat{\gamma} U \phi.$$

Recall that $\hat{\gamma}$ and $\hat{\phi}$ are expanded in powers of k^2 times k^{2l+1} , and $\hat{\eta}$ in powers of k^2 times k^{-l} . So

$$\omega_e(k) = 1 + (\alpha_e + \beta_e k^2 + O(k^4)) + i(\gamma_e k^{2l+1} + O(k^{2l+3}))$$

with real coefficients.

Now for $l=0$, we have

$$\omega_0(\kappa) = 1 + \alpha_0 + i\beta_0 \kappa + O(\kappa^2), \quad (3.2)$$

and if $\omega_0(\kappa=0) = 0$, $1 = -\alpha_0$, and it vanishes linearly. For $l > 1$, we have

$$\omega_l(\kappa) = 1 + \alpha_l + \beta_l \kappa^2 + O(\kappa^3 \text{ or } \kappa^4), \quad (3.3)$$

again, if it vanishes it does so quadratically.

Lemma 3.5: If f is analytic with a simple zero at z_0 , then f'/f has a simple pole of residue 1 at z_0 :

Since f has a simple zero at z_0 , $\exists g$ analytic such that $f(z) = (z - z_0)g(z)$, with $g(z_0) \neq 0$. Now

$$\frac{f'(z)}{f(z)} = \frac{g(z) + (z - z_0)g'(z)}{(z - z_0)g(z)} = \frac{1}{z - z_0} + \frac{g'(z)}{g(z)},$$

since $g'(z)/g(z)$ is analytic at z_0

because $g(z_0) \neq 0$, f'/f has a simple pole at z_0 with residue 1.

Lemma 3.6 : There are no zero energy bound states for $l=0$, and for $l>0$ there is a zero energy bound state of angular momentum l if and only if we have $\psi_l(0) = 0$.

For zero energy the radial equation (2.6) becomes $\left[\frac{d^2}{dr^2} - \frac{l(l+1)}{r^2} - U(r) \right] y(r) = 0$, where

$\frac{l(l+1)}{r^2}$ dominates for small and large r .

Without the potential, the solutions are r^{l+1} and r^{-l} (for small and large r).

In the case $l=0$, $r, 1$ are not normalizable, so a bound state of $l=0$ and zero energy is impossible. For $l>0$,

let $\tilde{\phi}_l := \frac{\phi_l}{r^{l+1}} \xrightarrow{r \rightarrow 0} \frac{\dot{\phi}_l}{r^{l+1}} \xrightarrow{r \rightarrow 0} \frac{r^{l+1}}{(2l+1)!!}$, normalizable at the origin

$\tilde{\chi}_l^\pm := r^l \chi_l^\pm \xrightarrow{r \rightarrow \infty} r^l h_l^\pm \xrightarrow{r \rightarrow \infty} r^{-l}$, normalizable at infinity

appropriate zero energy basis. Since

$$\omega_l(\kappa) = \frac{1}{\kappa} W(\chi^\pm, \phi) = \frac{1}{\kappa} r^{l+1} \frac{1}{r^l} W(\tilde{\chi}^\pm, \tilde{\phi}) = W(\tilde{\chi}^\pm, \tilde{\phi}),$$

we have a normalizable solution if and only if $\omega_l(0) = 0$.

Now we are ready to prove

Theorem 3.6 (Levinson's theorem): For any spherical potential subject to our usual assumptions, the phase shifts satisfy

$\delta_l(0) - \delta_l(\infty) = n_l \pi$, where n_l is the number of bound states with angular momentum l , if $\omega_0(0) \neq 0$. In the case that $\omega_0(0) = 0$, we have $\delta_0(0) - \delta_0(\infty) = (n_0 + \frac{1}{2}) \pi$.

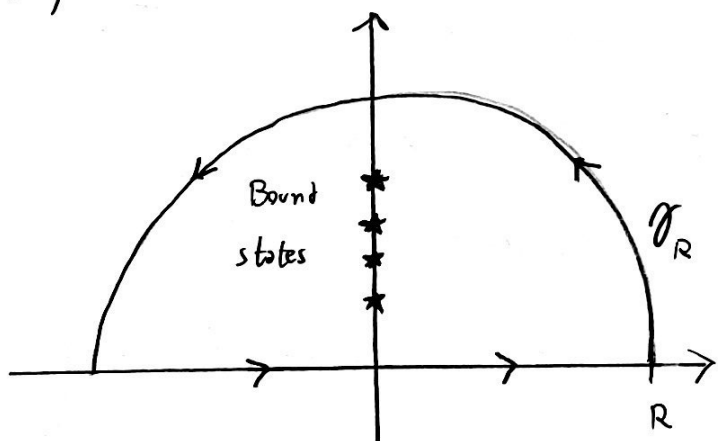
Proof: According to the Lemma 3.3, for large enough $|k|$ the Jost function is zeroless, and according to Lemma 3.4 the zeros don't accumulate at 0. So we conclude that, since zeros of analytic functions are isolated, there is a finite amount n_l of zeros of $\omega_l(k)$ in the upper half plane, all of them in the imaginary axis. Let's suppose that $\omega_l(0) \neq 0$. Then we consider the closed contour integral

$$I = \oint_{\gamma_R} d\kappa \frac{\dot{\omega}_e(\kappa)}{\omega_e(\kappa)}, \quad \text{with a semicircle contour}$$

large enough to contain all the zeros of ω_e . Using Cauchy's theorem and

Lemma 3.5, we see

that $I = 2\pi i m_e$, and the radius of the semicircle can be



taken to infinity without changing this result. Since Lemma 3.3 ensures that the contribution of the semicircle goes to zero, we have

$$2\pi i m_e = \oint_{\gamma_R} d\kappa \frac{\dot{\omega}_e(\kappa)}{\omega_e(\kappa)} = \oint_{\gamma_R} d \ln \omega_e(\kappa) = \lim_{R \rightarrow \infty} \oint_{\gamma_R} d \ln \omega_e(\kappa)$$

$$= \int_{-\infty}^{+\infty} d \ln \omega_e(\kappa) d\kappa = I. \quad \text{Finally, since we}$$

have from our discussion on section 2.4

that $\omega_e(\kappa) = |\omega_e(\kappa)| e^{-i\delta_e(\kappa)}$, and since

for real κ $\omega_e(-\kappa) = \omega_e(\kappa^*)^* = \omega_e(\kappa)^*$,

$$\ln \omega_e(\kappa) = \ln |\omega_e(\kappa)| - i\delta_e(\kappa) \quad \text{for } \kappa \geq 0$$

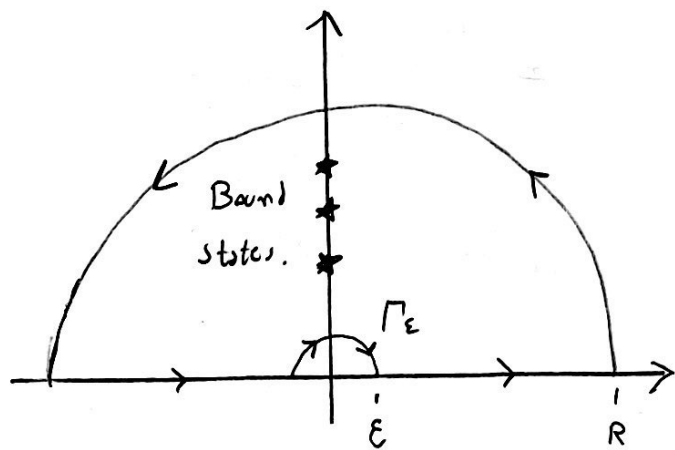
$$\ln \omega_e(\kappa) = \ln |\omega_e(\kappa)| + i\delta_e(\kappa) \quad \text{for } \kappa < 0.$$

Since $I = 2\pi i n_x$ is purely imaginary, the real part contribution vanishes, so

$$I = -2i \int_0^{+\infty} d\delta_x(\kappa) = 2i (\delta_x(0) - \delta_x(\infty)) = 2\pi i n_x,$$

so we finally obtain $\delta_x(0) - \delta_x(\infty) = n_x \pi$.

Now let's suppose that $\omega_x(0) = 0$. The contour has to be modified to exclude the singularity of $\dot{\omega}/\omega$: We can exclude it by adding a smaller semicircle Γ_ϵ , small enough to only contain the $\omega_x(0) = 0$ zero of ω_x , as Lemma 3.4 allows us to do:



$$\text{Now } I = \oint_{\Gamma_{R,\epsilon}} d\kappa \frac{\dot{\omega}_x(\kappa)}{\omega_x(\kappa)},$$

$$I = 2\pi i (\text{zeros in } \text{Im } \kappa > 0) =$$

$$= 2i (\delta_x(0) - \delta_x(\infty)) +$$

$$+ \lim_{\epsilon \rightarrow 0} \int_{\Gamma_\epsilon} d\kappa \frac{\dot{\omega}_x(\kappa)}{\omega_x(\kappa)}. \text{ The}$$

expansions (3.2) and (3.3) guarantee that

the zero of ω_x at 0 is simple if $l=0$ or double if $l>0$ (only for $l=0$ this can be a bound state zero!).

We know that $\frac{\dot{\omega}_l(\kappa)}{\omega_l(\kappa)}$ has a simple

pole of residue $m = 1, 2$ depending on

l , because as it's clear from the

expansion (3.3) the zero of ω_l

is either simple or double, so

we have $\frac{\dot{\omega}_l(\kappa)}{\omega_l(\kappa)} = \frac{m}{\kappa} + O(1) \text{ near}$

the origin but $\lim_{\epsilon \rightarrow 0} \int_{\Gamma_\epsilon} \frac{\dot{\omega}_l(\kappa)}{\omega_l(\kappa)} d\kappa =$

$$= \lim_{\epsilon \rightarrow 0} \int_{\Gamma_\epsilon} \frac{m dz}{z} = m \lim_{\epsilon \rightarrow 0} \int_{\pi}^0 \frac{\epsilon i e^{i\theta} d\theta}{\epsilon e^{i\theta}} = m i \int_{\pi}^0 d\theta = -m i \pi, \quad \text{so}$$

for $l > 0$ $I = 2\pi i (m_l - 1) = 2i(\delta_l(0) - \delta_l(\infty)) - 2i\pi$, the

result is the same (there's one missing bound state!). However for $l=0$, we obtain

$$2\pi i m_0 = 2i(\delta_0(0) - \delta_0(\infty)) - i\pi, \quad \text{and we}$$

$$\text{have } \delta_0(0) - \delta_0(\infty) = (m_0 + \frac{1}{2})\pi.$$

3.5 Scattering length

At small energies, the scattering amplitude is dominated by the scattering length.

Recall that the partial-wave amplitude

$$f_l(k) = \frac{S_l(k) - 1}{2ik} = \frac{\omega_l(-k) - \omega_l(k)}{2ik \omega_l(k)}$$

Using the integral form (2.9) of $\omega_l(k)$,

$$\omega_l(k) = 1 + \frac{1}{k} \int dr \hat{h}^+(kr) U(r) \phi_k(r),$$

and since $\hat{h}^+(-kr) \phi_{-k}(r) = (-1)^l \hat{h}^-(kr) (-1)^{l+1} \phi_k(r) = -\hat{h}^-(kr) \phi_k(r)$, we obtain

$$f_l(k) = \frac{1}{\omega_l(k) k^2} \int_0^\infty dr \left(\frac{\hat{h}^+(kr) - \hat{h}^-(kr)}{2i} \right) U(r) \phi_{k,l}(r)$$

$$= \frac{1}{\omega_l(k) k^2} \int_0^\infty dr \hat{\partial}_l(kr) U(r) \phi_{k,l}(r)$$

For physical k , both $|\phi_l|, |\hat{\partial}_l| \leq \left(\frac{|k|r}{1+|k|r} \right)^{l+1}$,

so for a neighborhood of the origin, provided that $f_l(0) \neq 0$,

$$|\psi_\ell(k)| \leq \frac{\text{const.}}{|\psi_\ell(0)|} \frac{1}{k^2} \int_0^\infty dr |U(r)| \left(\frac{|k|r}{1+|k|r} \right)^{2\ell+2} \leq \frac{\text{const.}}{|\psi_\ell(0)|} k^{1/2\ell} \int_0^\infty dr r^{2\ell+2} |U(r)|$$

so provided that the integral converges (V vanish eventually, V exponentially bounded, etc.), the expansion of ψ_ℓ around $k=0$ will have the form

$$\psi_\ell(k) = -a_\ell k^{2\ell} + O(k^{2\ell+1}), \quad \text{where } a_\ell$$

is the definition of scattering length.

4. Resonances

4.1 Resonances vs. poles of the S matrix elements

Let's assume that the potential is reasonable, and that $\psi_\ell(k_0) = 0$ for some k_0 with $\text{Im } k_0 < 0$, $k = k_R - i k_I$. By exactly the same arguments used to prove theorem 2.1, we can see that $\phi_{k_0, \ell}$ is not normalizable, so it can't represent a bound state. Further, suppose that k_0 is a simple zero

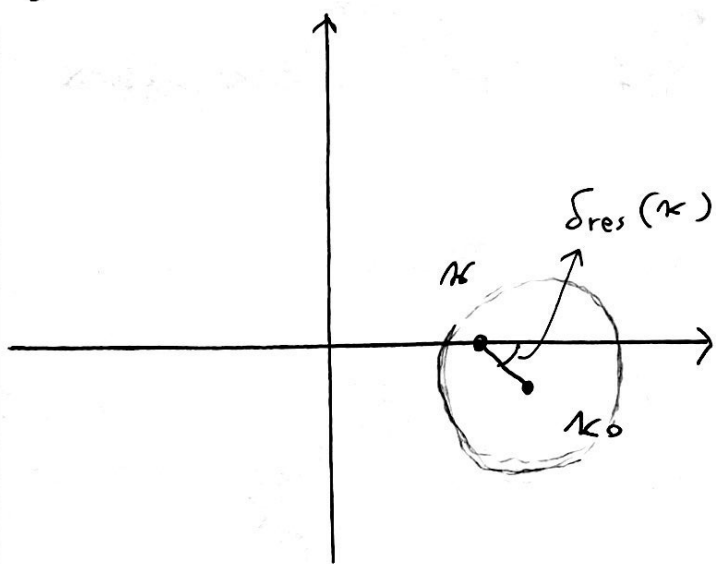
So since ω_l is analytic at κ_0 , we can approximate $\omega_l(\kappa) \approx \dot{\omega}_l(\kappa_0)(\kappa - \kappa_0)$ near κ_0 . We've already seen that

$\arg(\omega_l(\kappa)) = -\delta_l(\kappa)$, so we can approximate the phase shift to

$$\delta_l(\kappa) \approx -\arg \dot{\omega}_l(\kappa_0) - \arg(\kappa - \kappa_0) \quad (4.1)$$

$$\equiv \delta_{bg} + \delta_{res}(\kappa),$$

where $\delta_{res}(\kappa) = -\arg(\kappa - \kappa_0)$. If κ_0 is close to the real line this approximation is good for physical κ near κ_R :



As $\kappa < \kappa_R$ approaches κ_R , $\delta(\kappa)$ increases

fast from δ_{bg} to

$\delta_{bg} + \pi$. This

rapid increase defines

a resonance of angular momentum l .

Notice that

$$\sin \delta_{res}(\kappa) = \frac{\kappa_{\mp}}{\sqrt{(\kappa - \kappa_R)^2 + \kappa_{\mp}^2}} \quad (4.2)$$

so for a small δ_{bg} (pure Breit-Wigner)

we have $\sigma_e(k) \propto \sin^2 \delta_e(k) = \frac{k_I^2}{(k - k_R)^2 + k_I^2}$

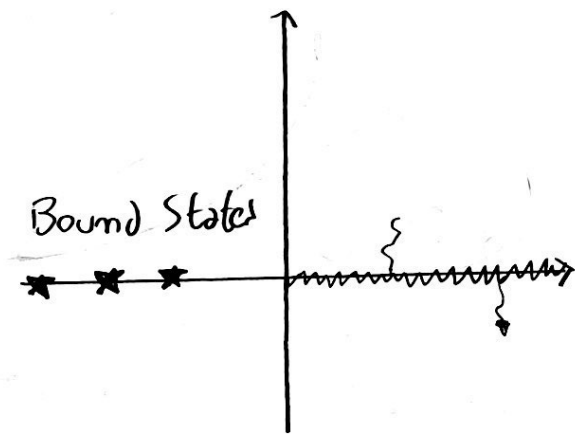
With a peak at $k = k_R$.

Notice however that not always there is a correspondence between zeros of $\omega_e(k)$ on $\text{Im } k < 0$ (or poles of s_e) and resonances. If the zero is far away from the real axis it will have no physical effect whatsoever.

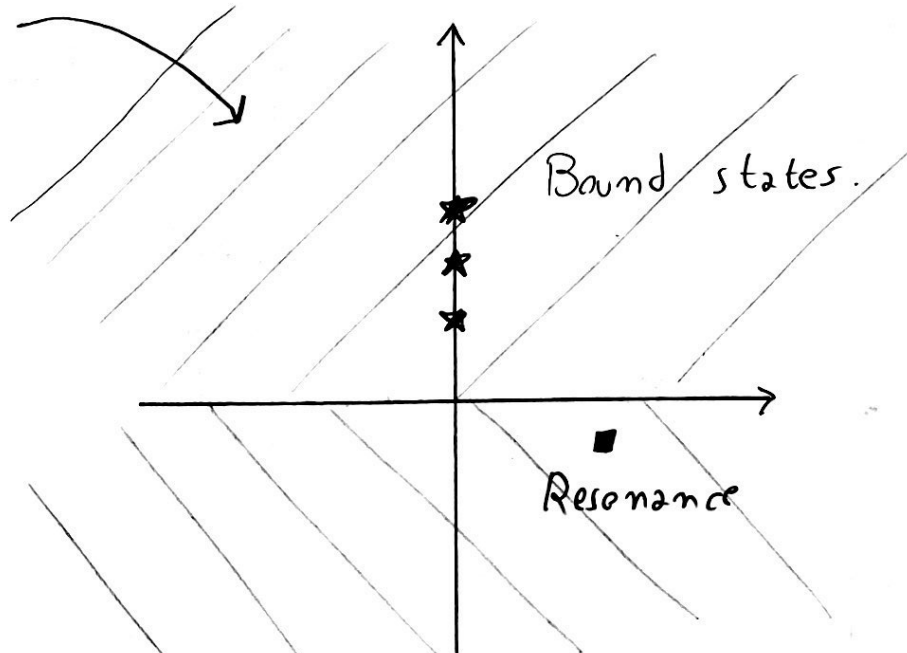
4.2 Expressing σ_e in terms of E .

Since $k = \sqrt{2mE}$ and the square root is in fact a double valued function, it is defined on a Riemann surface $\mathcal{R}$, that can be covered by two sheets:

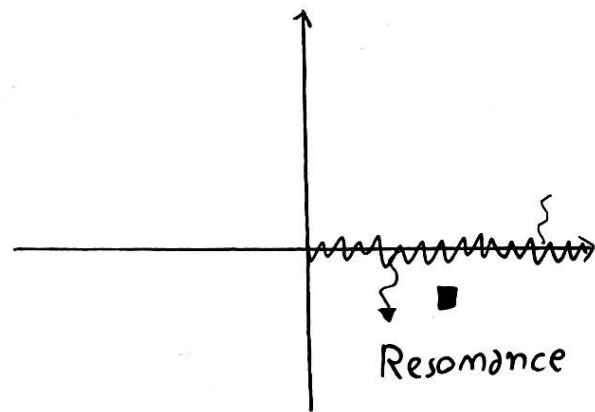
Physical Sheet



$$re^{i\theta} \rightarrow \sqrt{r} e^{i\theta/2} \quad \mathbb{K} \quad \text{Plane}$$



Second sheet



$$re^{i\theta} \rightarrow -\sqrt{r} e^{i\theta/2} = \sqrt{r} e^{i(\theta+2\pi)/2}$$

So in terms of E , $\omega_e(E)$ is defined

on a Riemann surface $\omega_e: \mathbb{R} \rightarrow \mathbb{C}$.

On the second sheet, we define

a resonance via $E_0 = \frac{k_0^2}{2m} = E_R - i\frac{\Gamma}{2}$,

a pure Breit-Wigner has the form (4.2)

$$\sigma_e(E) \propto \frac{(\Gamma/2)^2}{(E-E_R)^2 + (\Gamma/2)^2}$$

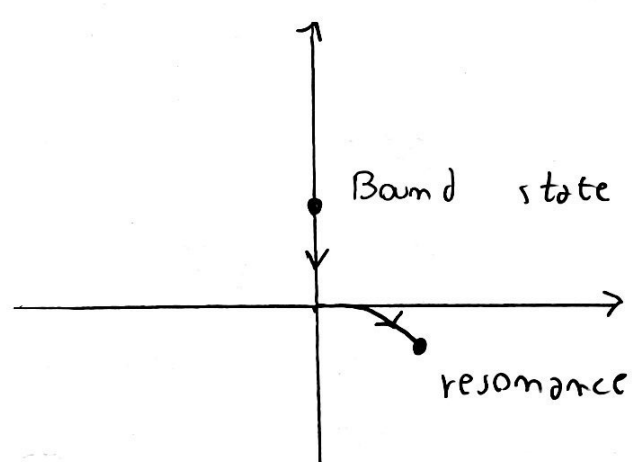
4.3 Bound states and resonances:

We will now use the dependence of $\omega_e(k, \lambda)$ on the strength parameter λ to show that resonances can be regarded as "would-be bound states". For a reasonable potential, $\omega_e(k, \lambda)$ is analytic at $k=0$ and entire on λ . Let's move λ around physical (real) values near λ_0 , and expand $\omega_e(k, \lambda)$ as

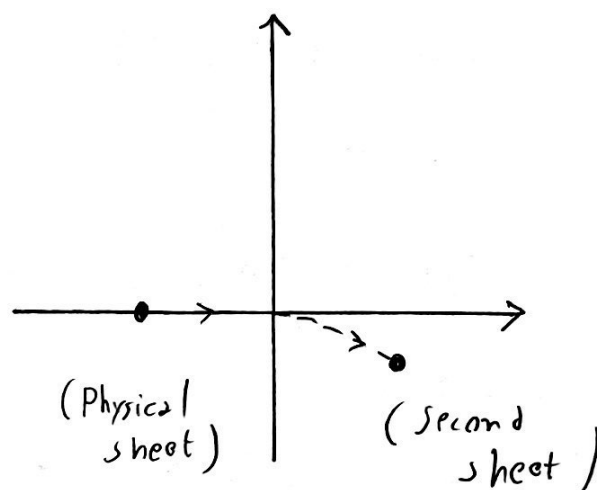
$$\omega_e(k, \lambda) = \sum_{m, n \geq 0} \alpha_{mn} k^m (\lambda - \lambda_0)^n$$

Now suppose that $\lambda > 0$, and for $\lambda = \lambda_0$, $\omega_e(0, \lambda_0) = 0$. Using the expansion (3.3), $\omega_0(p, \lambda_0) = \alpha k^2 + \beta (\lambda - \lambda_0) + \dots$, so for λ close to λ_0 the zeros are located at $k_0 \approx \pm i \sqrt{\frac{\beta}{\alpha} (\lambda - \lambda_0)}$. On the k plane, as $\lambda > \lambda_0$ approaches a bound state (zero on $\text{Im } k > 0$) approaches the origin, and splits into two zeros, one of them possibly a resonance.

On the Riemann surface, this corresponds to a migration of the zero E_0 from the physical sheet from a resonance position on the second sheet:



k Plane



R surface.

For those resonances close to the origin "emerging" from bound states, δ_{bg} is always small. Since as long κ_0 represent a bound ~~state~~ state we have κ_0 pure imaginary, we see that $\omega(\kappa_0) = \omega(-\kappa_0^*)^* = \omega(\kappa_0)^*$, $\omega(\kappa_0)$ is real, so $\dot{\omega}(\kappa_0)$ also is.

As λ varies slowly, and κ_0 becomes a resonance, $\dot{\omega}(\kappa_0)$ remains predominantly real, so $\delta_{bg} = -\arg \dot{\omega}(\kappa_0)$ is small.

4.4 Resonances viewed as "unstable bound states"

Let's study the scattering of a wave package $|\psi_{in}\rangle = \int d^3k \phi(\vec{k}) |\vec{k}\rangle$, where $\phi(\vec{k})$ is peaked about $\vec{k} = \vec{k}_0$. The scattering wave function will be (recall our discussion on section 1.1)

$$\begin{aligned} \psi(\vec{r}, t) &= \langle \vec{r} | U(t) | \psi \rangle = \langle \vec{r} | U(t) U_+ | \psi_{in} \rangle = \\ &= \langle \vec{r} | U(t) U_+ \int d^3k \phi(\vec{k}) |\vec{k}\rangle = \\ &= \langle \vec{r} | U(t) \int d^3k \phi(\vec{k}) |\vec{k}_+\rangle = \int d^3k e^{-iE_k t} \phi(\vec{k}) \langle \vec{r} | \vec{k}_+ \rangle \\ &\xrightarrow{r \rightarrow \infty} \frac{1}{(2\pi)^{3/2}} \int d^3k \phi(\vec{k}) e^{-iE_k t} \left(e^{i\vec{k} \cdot \vec{r}} + f(k \hat{r} | \vec{k}) \frac{e^{i\vec{k} \cdot \vec{r}}}{r} \right) \end{aligned}$$

$= \psi_{in}(\vec{r}, t) + \psi_{sc}(\vec{r}, t)$, a free propagating wave package and an outgoing spherical wave package. We now suppose that the only appreciable contribution to the scattering amplitude comes from a pure Breit-Wigner resonance

$E_0 = E_R - i \frac{\Gamma}{2}$, with width much smaller than the spread ΔE of $\phi(\vec{\kappa})$,

$\Gamma \ll \Delta E$. In that case,

$$f(\kappa \hat{r} | \vec{\kappa}) = \sum_{l \geq 0} (2l+1) f_l(E) P_l(\cos \theta)$$

$$= (2l+1) f_l(E) P_l(\cos \theta)$$

$$= -\frac{1}{2\kappa} (2l+1) \frac{\Gamma}{E - E_R + i\Gamma/2} P_l(\cos \theta)$$

So we'll have:

$$\begin{aligned} \psi_{sc}(\vec{r}, t) &\xrightarrow{r \rightarrow \infty} \text{const} \times \int d^3\kappa \phi(\vec{\kappa}) \bar{e}^{iE_\kappa t} f(\kappa \hat{r} | \vec{\kappa}) \frac{e^{i\kappa r}}{r} \\ &= \text{const} \times \frac{\Gamma}{r} \int d^3\kappa \frac{\phi(\vec{\kappa}) P_l(\cos \theta)}{\kappa(E_\kappa - E_R + i\Gamma/2)} e^{i(\kappa r - E_\kappa t)} \end{aligned}$$

Aligning $\vec{\kappa}$ and $\hat{z}$, and assuming $\phi(\vec{\kappa})$ invariant under rotations about $\hat{z}$,

we must have $\phi(\vec{\kappa}) = \sum_{l \geq 0} \phi_l(\kappa) P_l(\cos \theta) =$

$$= \frac{1}{\sqrt{m\kappa}} \sum_{l \geq 0} \phi_l(E_\kappa) P_l(\cos \theta) \quad \text{An integration}$$

on the angular part of $d^3\kappa$ will

then select only $\phi_l(E_\kappa)$ where

l is the angular momentum of the

$E_0 = E_R - i \frac{\Gamma}{2}$, with width much smaller than the spread ΔE of $\phi(\vec{\kappa})$, $\Gamma \ll \Delta E$. In that case,

$$\begin{aligned} f(\kappa \hat{r} | \vec{\kappa}) &= \sum_{l \geq 0} (2l+1) f_l(E) P_l(\cos \theta) \\ &= (2l+1) f_l(E) P_l(\cos \theta) \\ &= -\frac{1}{2\kappa} (2l+1) \frac{\Gamma}{E - E_R + i\Gamma/2} P_l(\cos \theta) \end{aligned}$$

So we'll have:

$$\begin{aligned} \psi_{sc}(\vec{r}, t) &\xrightarrow{r \rightarrow \infty} \text{const} \times \int d^3\kappa \phi(\vec{\kappa}) \bar{e}^{iE_{\kappa}t} f(\kappa \hat{r} | \vec{\kappa}) \frac{e^{i\kappa r}}{r} \\ &= \text{const} \times \frac{\Gamma}{r} \int d^3\kappa \frac{\phi(\vec{\kappa}) P_l(\cos \theta)}{\kappa(E_{\kappa} - E_R + i\Gamma/2)} e^{i(\kappa r - E_{\kappa}t)} \end{aligned}$$

Aligning $\vec{\kappa}$ and $\hat{z}$, and assuming $\phi(\vec{\kappa})$ invariant under rotations about $\hat{z}$,

We must have $\phi(\vec{\kappa}) = \sum_{l \geq 0} \phi_l(\kappa) P_l(\cos \theta) =$

$$= \frac{1}{\sqrt{m\kappa}} \sum_{l \geq 0} \phi_l(E_{\kappa}) P_l(\cos \theta) \quad \text{An integration}$$

on the angular part of $d^3\kappa$ will

then select only $\phi_l(E_{\kappa})$ where

l is the angular momentum of the

resonance. So we're left with

$$\psi_{sc}(\vec{r}, t) \xrightarrow{r \rightarrow \infty} \text{const} \times \frac{\Gamma P_2(\cos\theta)}{r} \int_0^\infty dk \, k^{-1/2} \frac{\phi_2(E_k) e^{i(kr - E_k t)}}{E_k - E_R + i\Gamma/2}$$

Since the integrand is mostly appreciable near E_R , we can approximate

$$k \approx k_R + \left. \frac{dk}{dE} \right|_{E_R} (E_k - E_R) = k_R - \frac{E_k - E_R}{v_R}, \quad \text{hence}$$

$$e^{i(kr - E_k t)} \approx e^{i(k_R r - E_R t)} e^{-i(E_k - E_R)(t - \frac{r}{v_R})}$$

And since $\Delta E \gg \Gamma$, $\phi_2(E_k)$ varies

slowly near E_R , so we can

take it out of the integral as

$\phi_2(E_R)$ (the point where the integrand

is most prominent).

$$\psi_{sc}(\vec{r}, t) \xrightarrow{r \rightarrow \infty} \text{const.} \times \Gamma P_2(\cos\theta) \phi_2(E_R) \frac{e^{i(k_R r - E_R t)}}{\sqrt{k_R} r} \times$$

$$\times \int_0^\infty dE \frac{e^{-i(E - E_R)(t - r/v_R)}}{E - E_R + i\Gamma/2}$$

This integral can be extended to $-\infty$ without harm, since the integrand is only appreciable near E_R .

We now study the integral

$$\int_{-\infty}^{+\infty} dz \frac{e^{-iz\tau}}{z + i\Gamma/2}$$

It can be solved by contour integration in

the complex plane of z . At $z = -i\Gamma/2$,

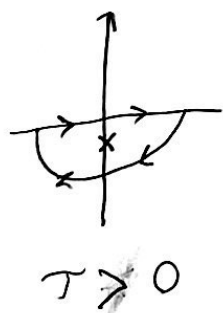
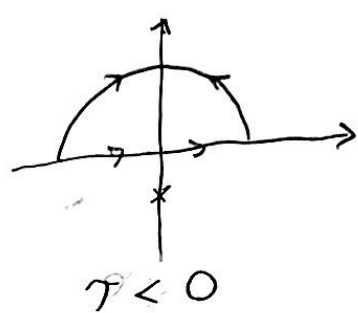
the integrand has a simple pole, with residue $e^{-i(-i\Gamma/2)\tau} = e^{-\Gamma/2\tau}$. If $\tau > 0$,

we have to close the contour in the lower half plane, and by Cauchy's

theorem $\int_{-\infty}^{+\infty} dz \frac{e^{-iz\tau}}{z + i\Gamma/2} = -2\pi i e^{-\Gamma/2\tau}$. And

for $\tau < 0$ we close the contour

on the upper half plane where the integrand is analytic, so the integral vanishes.



$$\int_{-\infty}^{+\infty} dz \frac{e^{-iz\tau}}{z + i\Gamma/2} = \Theta(\tau) (-2\pi i e^{-\Gamma/2\tau})$$